

Investigating magnetic fluctuations in gyrokinetic simulations of tokamak SOL turbulence

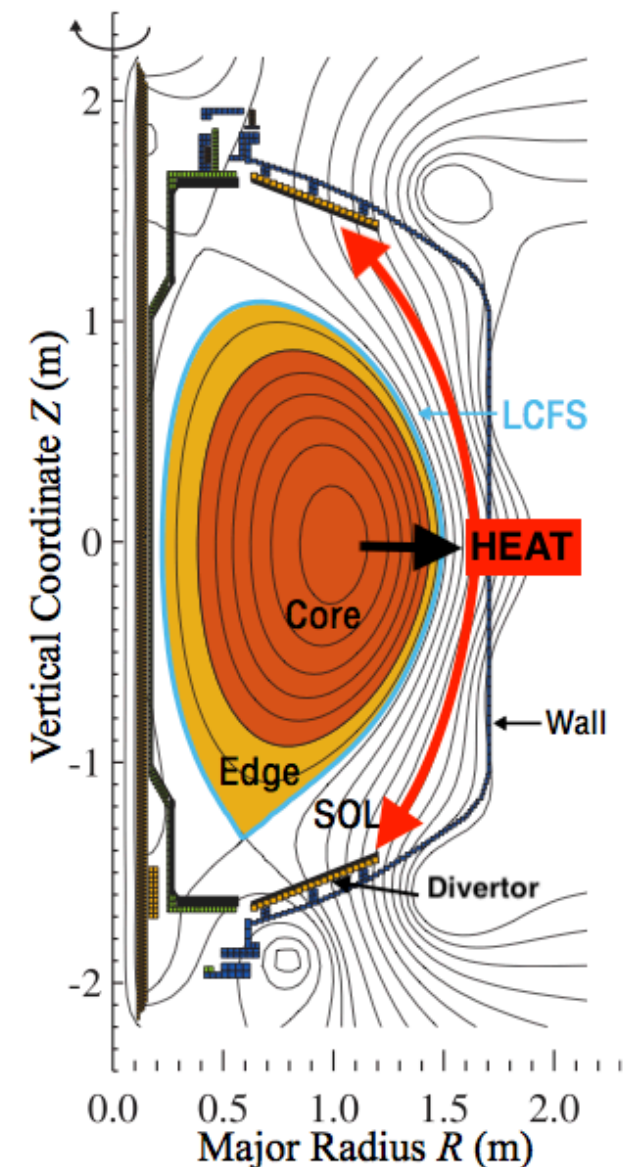
Noah Mandell

PPPL Theory Research & Review Seminar
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Why is SOL turbulence important?

- Plasma properties in the tokamak edge/scrape-off layer (SOL) constrain component lifetime and reactor performance
- Heat exhausted in SOL could damage divertor plates if heat flux width is too narrow
 - Can SOL turbulence broaden the heat flux width?
 - Can electromagnetic effects be important for SOL turbulence?



Modeling the edge/SOL is challenging

- Gyrokinetic (GK) theory and simulation are important first-principles tools for studying turbulence and transport in fusion plasmas, but most present codes optimized for core, small fluctuations (δf)
- Edge/SOL more challenging: large-amplitude fluctuations, open field lines, plasma-wall interactions, X-point geometry, atomic physics, transition from kinetic to fluid regimes → need specialized full-f GK codes
- Including electromagnetic effects (allowing magnetic field to fluctuate) also challenging → all GK SOL results to-date have been electrostatic (no magnetic fluctuations)
- Several GK codes making great progress in edge/SOL
 - XGC1, COGENT, ELMFIRE, GENE, **Gkeyll**, etc
- **Essential to have several independent codes to attack from different perspective and cross-check on difficult turbulence problems!**

Methods for solving gyrokinetic system

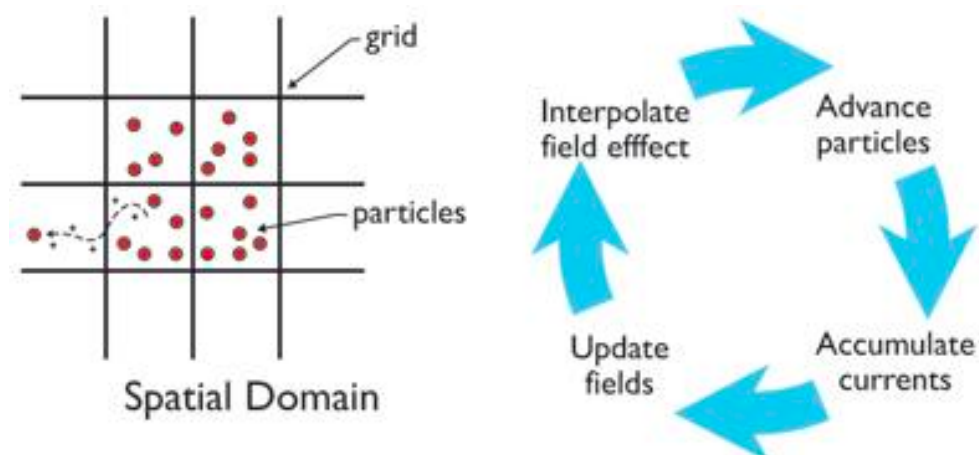
Particle-in-cell (Lagrangian)

Continuum (Eulerian)

Methods for solving gyrokinetic system

Particle-in-cell (Lagrangian)

- Sample phase space with ensemble of N_p ‘superparticles’
- Fields on 3D grid, particles move through grid
- Historically, EM fluctuations challenging in GK PIC codes due to numerical “Ampere cancellation” problem
 - Codes like ORB5 and XGC1 have made good recent progress to mitigate issue

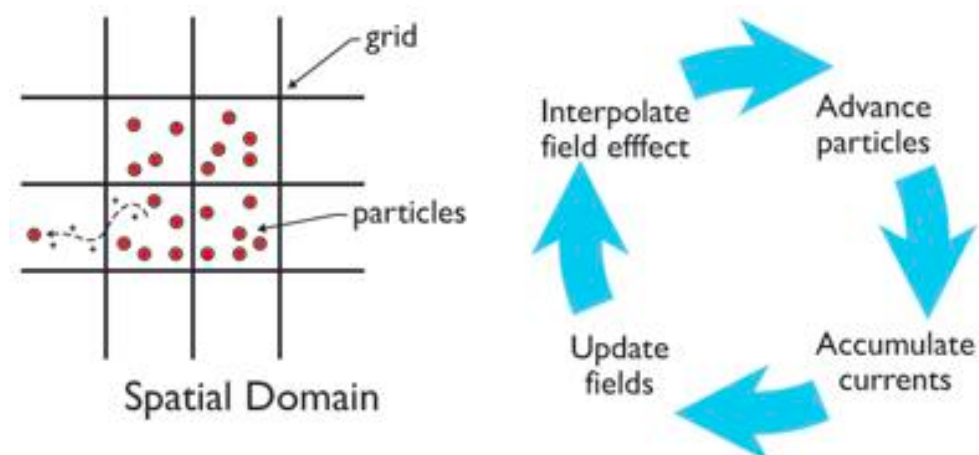


Continuum (Eulerian)

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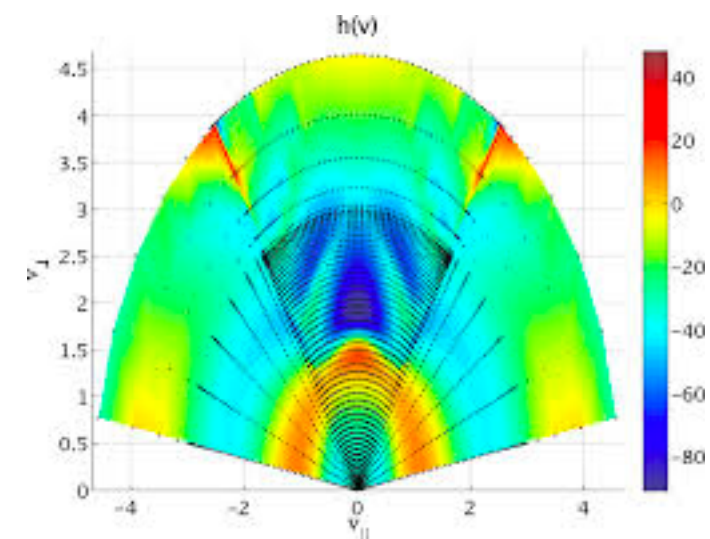
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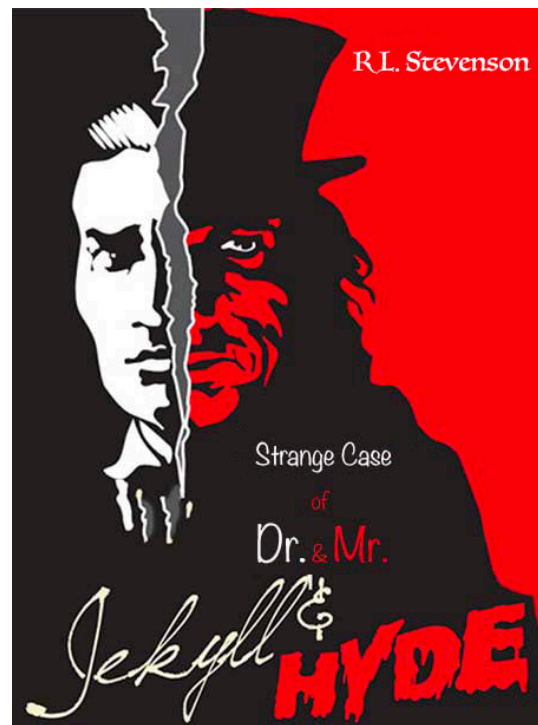
Continuum (Eulerian)

- Discretize distribution function $f(x, y, z, v_{\parallel}, \mu)$ on 5D phase space grid
- Can solve with standard PDE methods, e.g. spectral, finite volume, discontinuous Galerkin, etc.
- Continuum electromagnetic GK codes have mostly avoided the Ampere cancellation problem



Barnes et al, 2010

Methods for solving gyrokinetic system



<https://github.com/ammarrhakim/gkyl/>

<https://gkeyll.readthedocs.io>

Continuum (Eulerian)

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Gkeyll

- First successful continuum GK code on open field lines
- **First electromagnetic GK on open field lines**



Full- f electromagnetic gyrokinetics

EMGK equation, $f_s = f_s(\mathbf{R}, v_{\parallel}, \mu; t)$

$$\frac{\partial f_s}{\partial t} + \dot{\mathbf{R}} \cdot \nabla f_s + \dot{v}_{\parallel} \frac{\partial f_s}{\partial v_{\parallel}} = C[f_s] + S_s$$

with nonlinear phase-space trajectories

$$\begin{aligned} \dot{\mathbf{R}} = \{\mathbf{R}, H_s\} &= \frac{\mathbf{B}_0^* + \delta \mathbf{B}_{\perp}}{B_{\parallel}^*} v_{\parallel} + \frac{\hat{\mathbf{b}}}{q_s B_{\parallel}^*} \times (\mu \nabla B + q_s \nabla \phi) \\ \dot{v}_{\parallel} = \{v_{\parallel}, H_s\} - \frac{q_s}{m_s} \frac{\partial A_{\parallel}}{\partial t} &= - \frac{\mathbf{B}_0^* + \delta \mathbf{B}_{\perp}}{m_s B_{\parallel}^*} \cdot (\mu \nabla B + q_s \nabla \phi) - \frac{q_s}{m_s} \frac{\partial A_{\parallel}}{\partial t} \end{aligned}$$

where $\mathbf{B}_0^* = \mathbf{B}_0 + (m_s v_{\parallel} / q_s) \nabla \times \hat{\mathbf{b}}$ and $\delta \mathbf{B}_{\perp} = \nabla A_{\parallel} \times \hat{\mathbf{b}}$.

- No assumption of scale separation between background and fluctuations
- Taking long-wavelength (drift-kinetic) limit, neglecting gyroaveraging for now
- Using symplectic (v_{\parallel}) formulation of EMGK, so $\frac{\partial A_{\parallel}}{\partial t}$ appears explicitly

Full- f electromagnetic gyrokinetics

Quasineutrality equation (long-wavelength):

$$-\nabla \cdot \sum_s \frac{m_s n_{0s}}{B^2} \nabla_{\perp} \phi = \sum_s q_s \int d^3v f_s \quad (1)$$

Parallel Ampère equation:

$$-\nabla_{\perp}^2 A_{\parallel} = \mu_0 \sum_s q_s \int d^3v v_{\parallel} f_s \quad (2)$$

Can take $\frac{\partial}{\partial t}$ to get an exact Ohm's law:

$$-\nabla_{\perp}^2 \frac{\partial A_{\parallel}}{\partial t} = \mu_0 \sum_s q_s \int d^3v v_{\parallel} \frac{\partial f_s}{\partial t} \quad (3)$$

Writing GK eq. as

$$\frac{\partial f_s}{\partial t} = \frac{\partial f_s^*}{\partial t} + \frac{q_s}{m_s} \frac{\partial A_{\parallel}}{\partial t} \frac{\partial f_s}{\partial v_{\parallel}}, \quad (4)$$

where $\frac{\partial f_s^*}{\partial t}$ denotes all the terms in the gyrokinetic equation except the $\frac{\partial A_{\parallel}}{\partial t}$ term, can write Ohm's law as

$$\left(-\nabla_{\perp}^2 + \sum_s \frac{\mu_0 q_s^2}{m_s} \int d^3v f_s \right) \frac{\partial A_{\parallel}}{\partial t} = \mu_0 \sum_s q_s \int d^3v v_{\parallel} \frac{\partial f_s^*}{\partial t} \quad (5)$$

Ampère cancellation problem

- In p_{\parallel} formulation, Ampère's law:

$$\left(-\nabla_{\perp}^2 + C_n \sum_s \frac{\mu_0 q_s}{m_s} \int d^3p f \right) A_{\parallel} = C_j \mu_0 \sum_s \frac{q_s}{m_s^2} \int d^3p p_{\parallel} f$$

- “Cancellation problem” arises when there are small errors in the calculation of the integrals, represented by C_n and C_j (which should be exactly 1 in the exact system)
- Recall v_{\parallel} formulation Ohm's law... same problem...

$$\left(-\nabla_{\perp}^2 + C_n \sum_s \frac{\mu_0 q_s^2}{m_s} \int d^3v f_s \right) \frac{\partial A_{\parallel}}{\partial t} = C_j \mu_0 \sum_s q_s \int d^3v v_{\parallel} \frac{\partial f_s}{\partial t}^*$$

- The simplest Alfvén wave dispersion relation (slab geometry, uniform Maxwellian background with stationary ions) becomes (with $\hat{\beta} \equiv \frac{\beta_e}{2} \frac{m_i}{m_e}$)

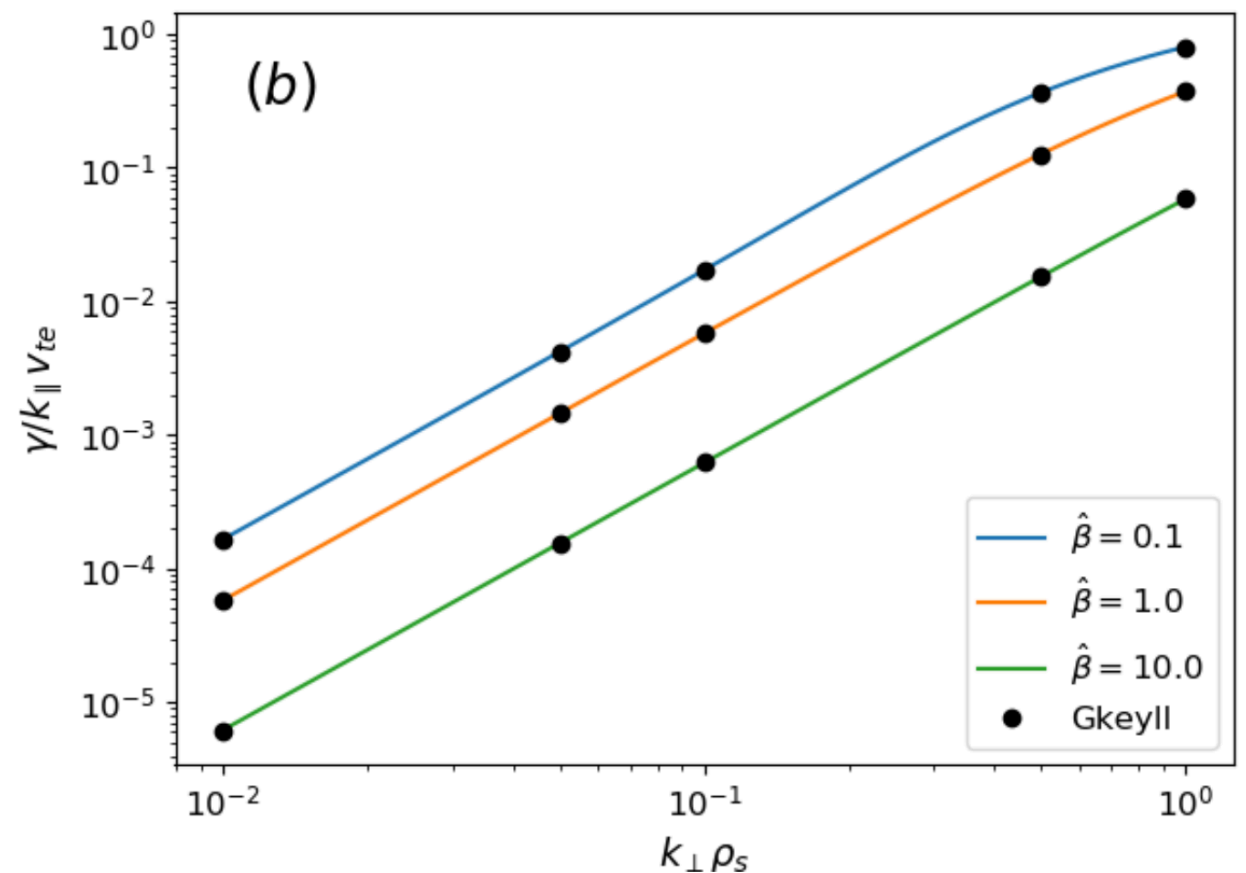
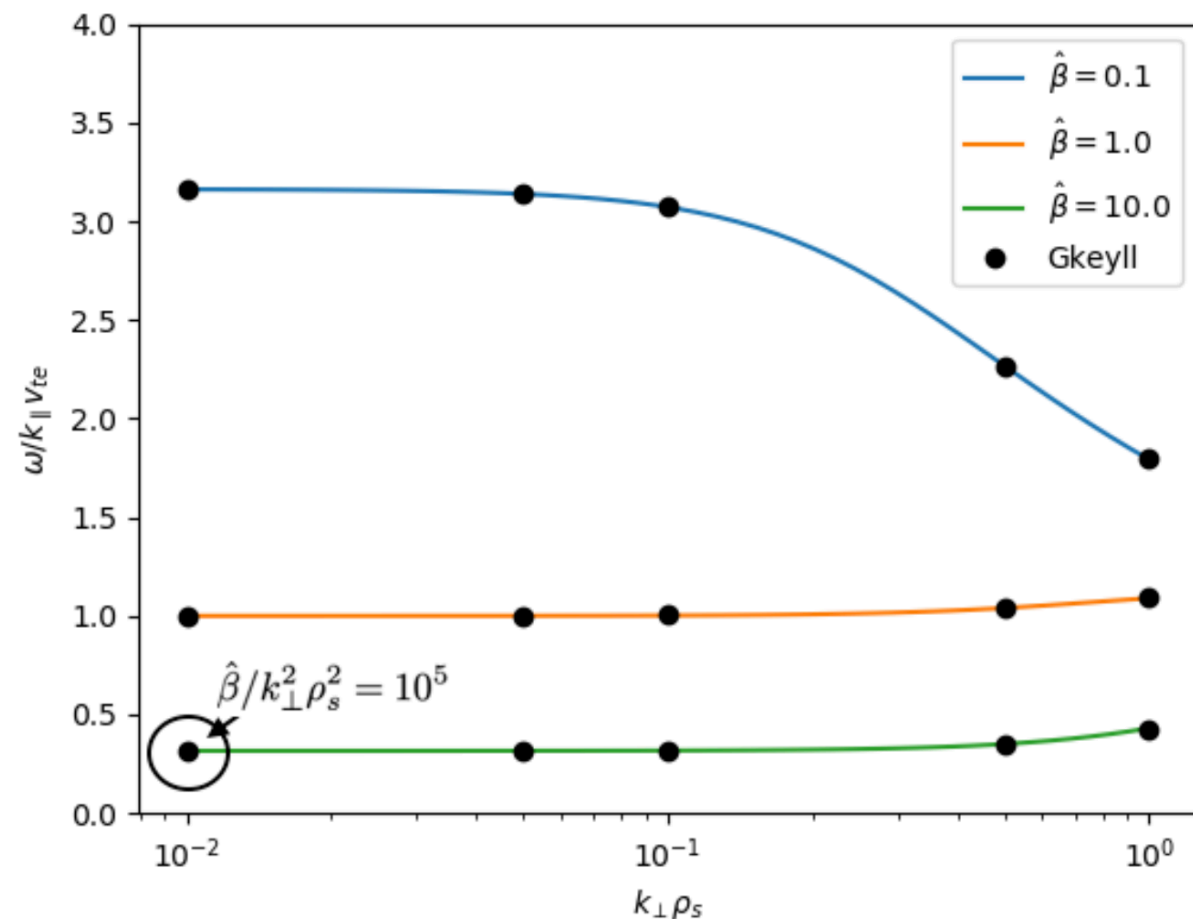
$$\omega^2 = \frac{k_{\parallel}^2 v_A^2}{C_n + k_{\perp}^2 \rho_s^2 / \hat{\beta}} \left[1 + (C_n - C_j) \frac{\hat{\beta}}{k_{\perp}^2 \rho_s^2} \right]$$

- This reduces to the correct result if integrals calculated consistently, so that $C_n = C_j$, but if not there will be errors $\sim \omega_H$ for modes with $\hat{\beta}/k_{\perp}^2 \rho_s^2 \gg 1$.

Gkeyll's DG scheme computes integrals consistently so that errors cancel exactly (appendix of Mandell et al, JPP 2020 shows numerical dispersion relation calculation)

Linear benchmark: kinetic Alfvén wave

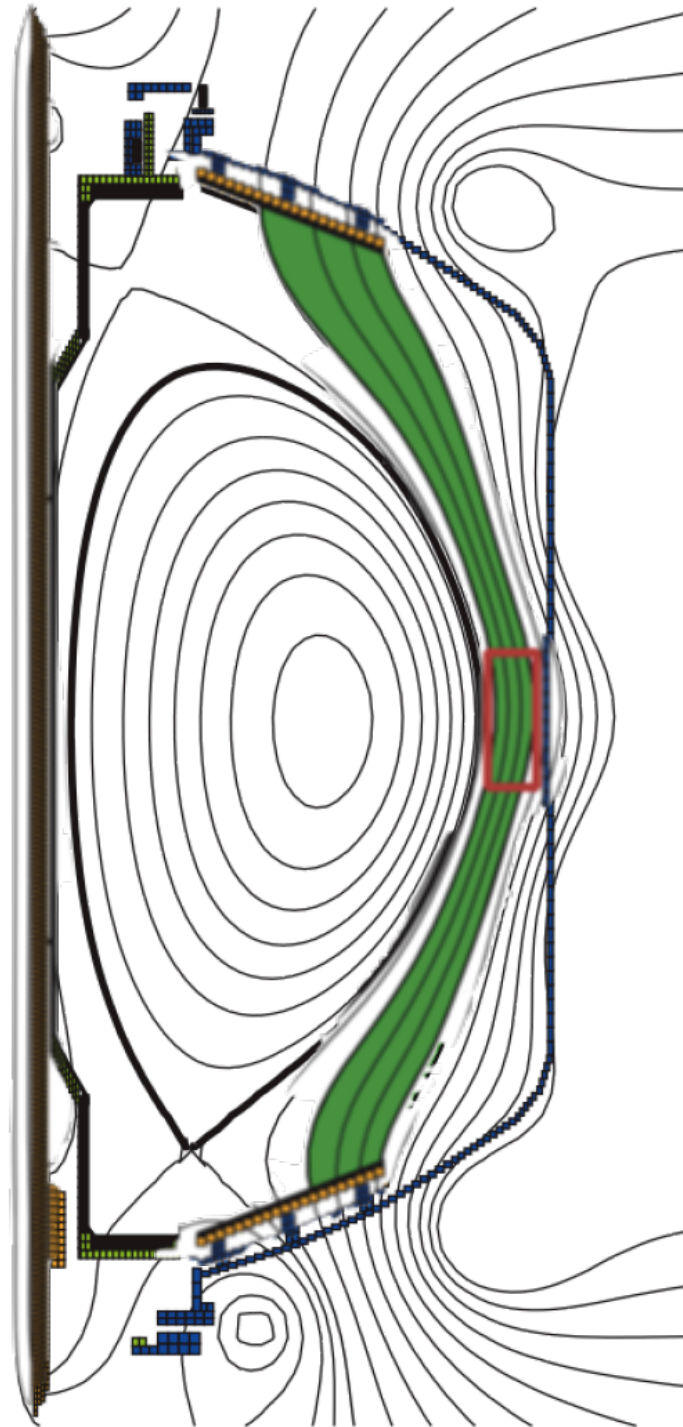
$$\omega^2 = \frac{k_{\parallel}^2 v_A^2}{C_n + k_{\perp}^2 \rho_s^2 / \hat{\beta}} \left[1 + (C_n - C_j) \frac{\hat{\beta}}{k_{\perp}^2 \rho_s^2} \right] \quad (k_{\perp} \rho_s \ll 1)$$



Gkeyll results match theory very well, even for case with $\frac{\hat{\beta}}{k_{\perp}^2 \rho_s^2} = \frac{\beta_e/2}{k_{\perp}^2 \rho_s^2} \frac{m_i}{m_e} = 10^5$

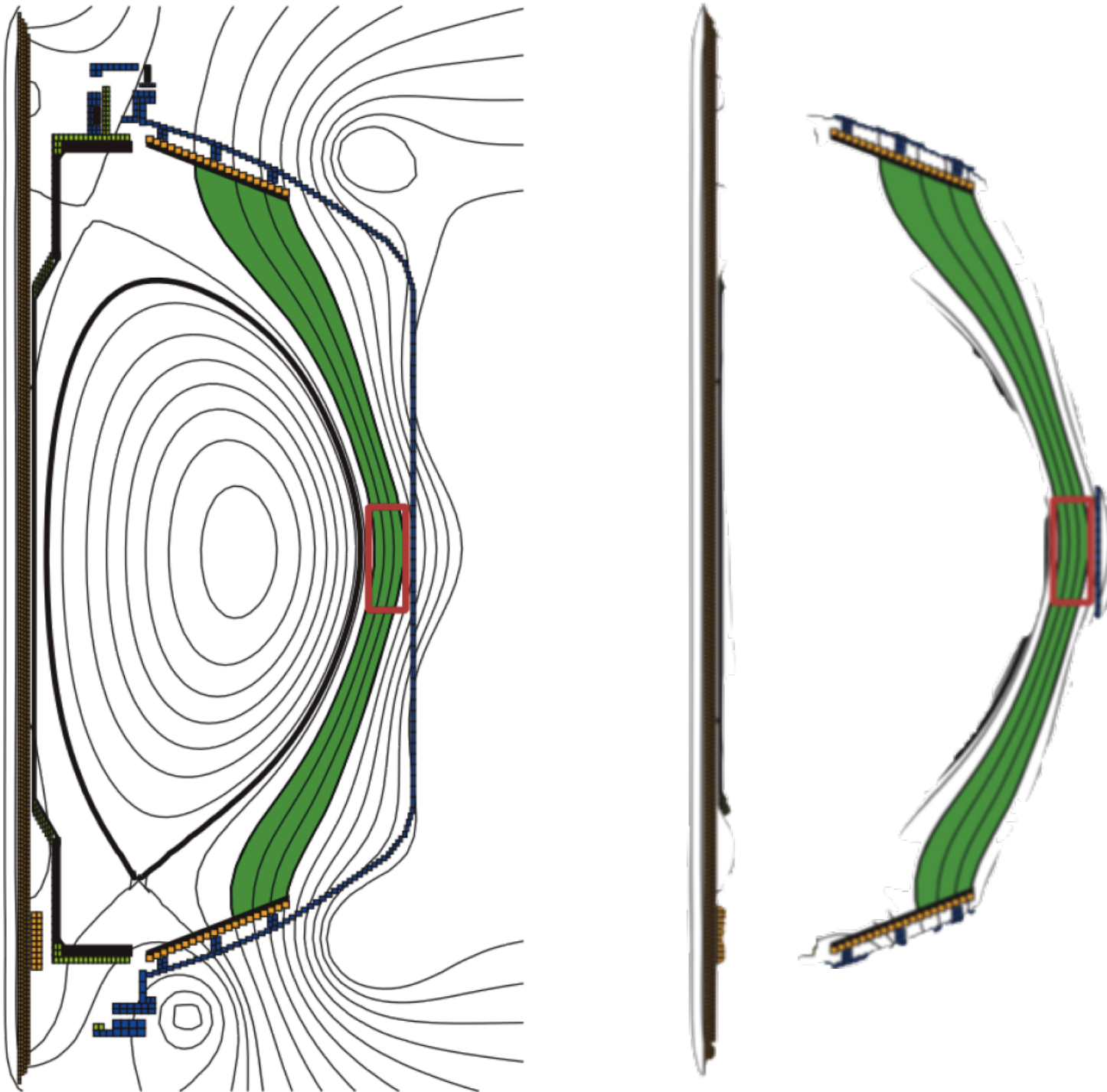
No cancellation problem!

Modeling the NSTX SOL with Gkeyll

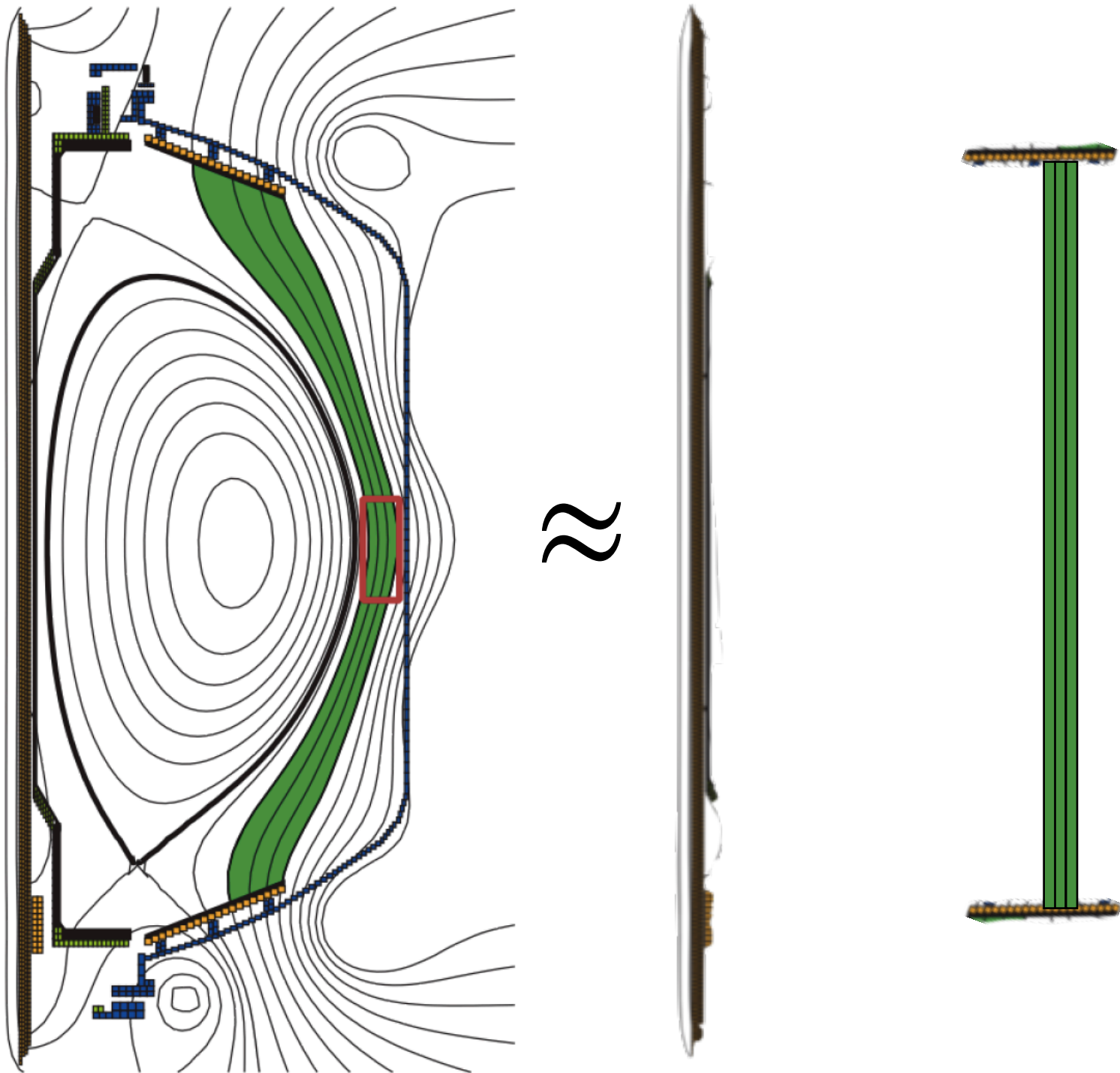


Modeling the NSTX SOL with Gkeyll

- Open-field-line region only

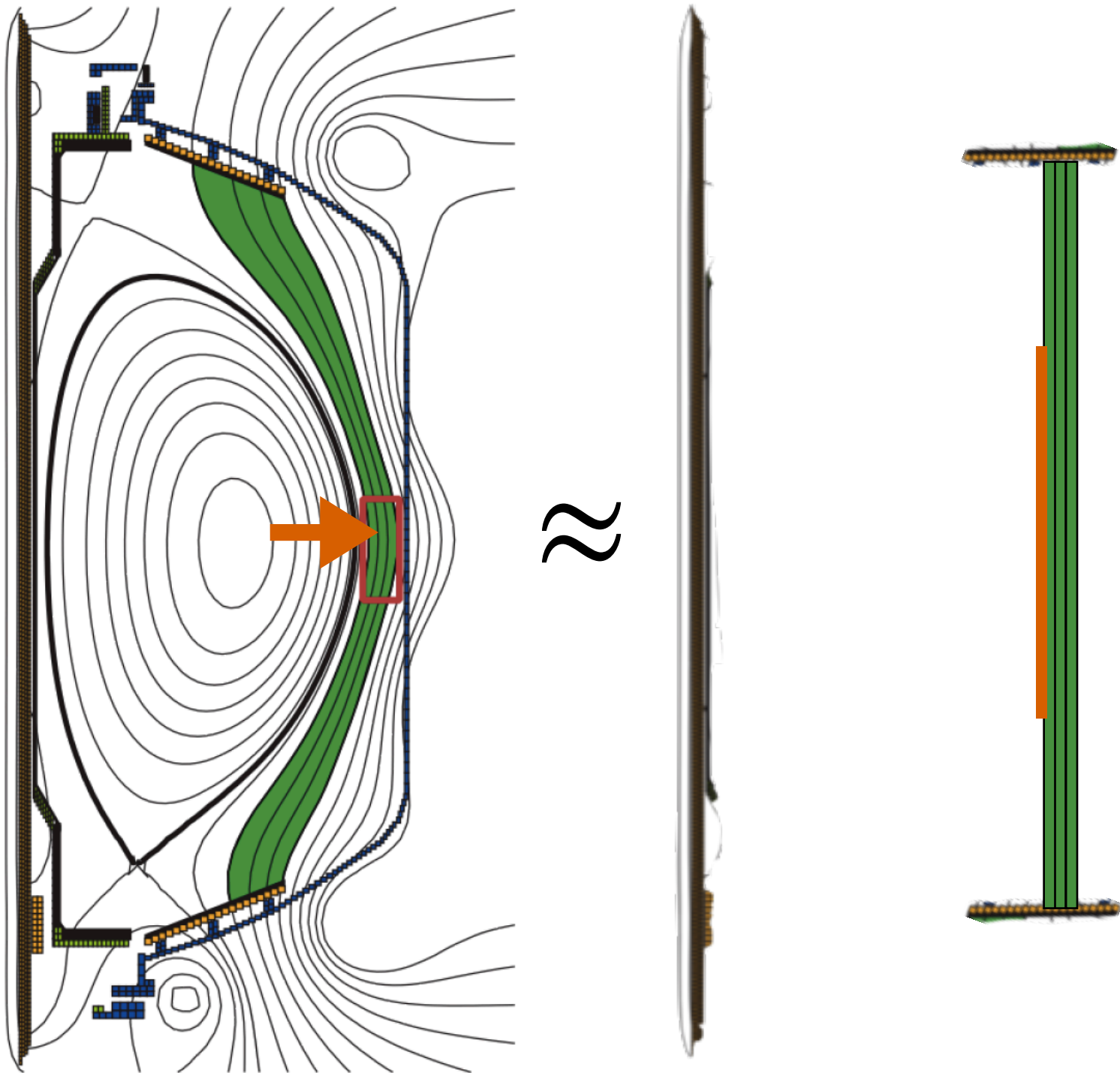


Modeling the NSTX SOL with **Gkeyll**



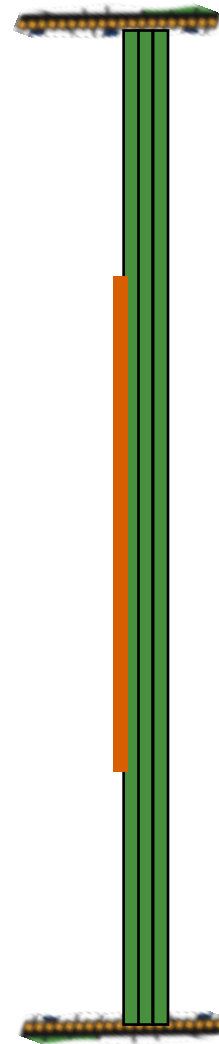
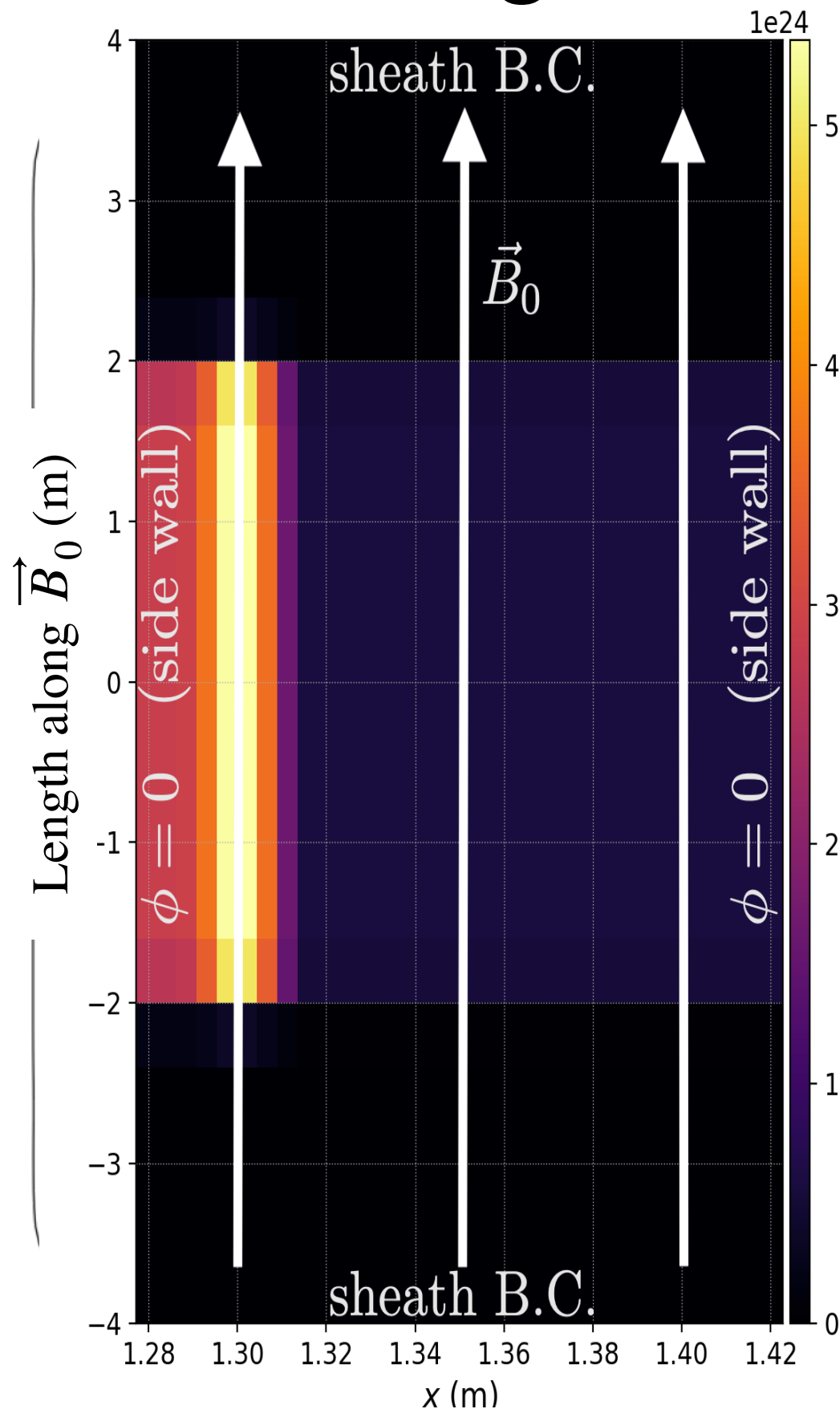
- Open-field-line region only
- Simplified helical geometry with vertical flux surfaces, const curvature and no shear (no X point)

Modeling the NSTX SOL with **Gkeyll**



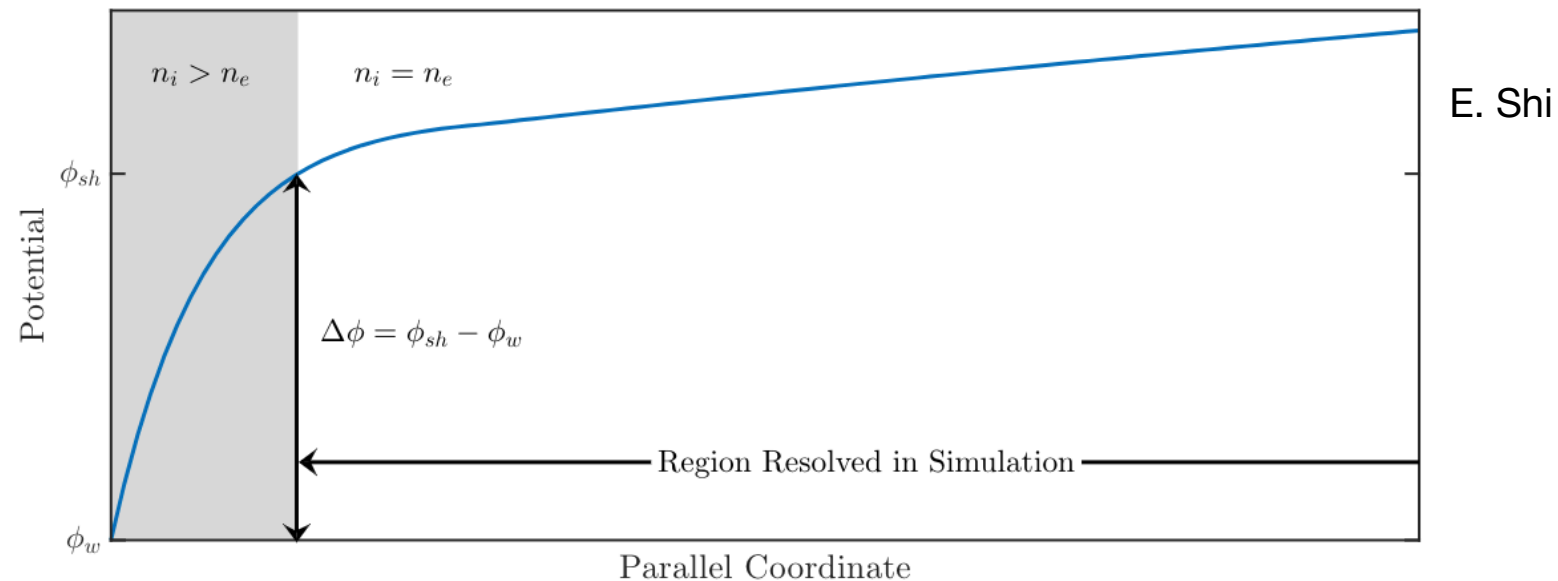
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- Model flux of heat and particles across separatrix with source

Modeling the NSTX SOL with **Gkeyll**



- Open-field-line region only
- Simplified helical geometry with vertical flux surfaces, const curvature and no shear (no X point)
- Model flux of heat and particles across separatrix with source
- Boundary conditions:
 - perfectly conducting walls ($\phi = A_{||} = 0$) in radial direction, x
 - periodic in binormal direction, y
 - conducting sheath model BC along field line, z

Conducting-sheath boundary condition

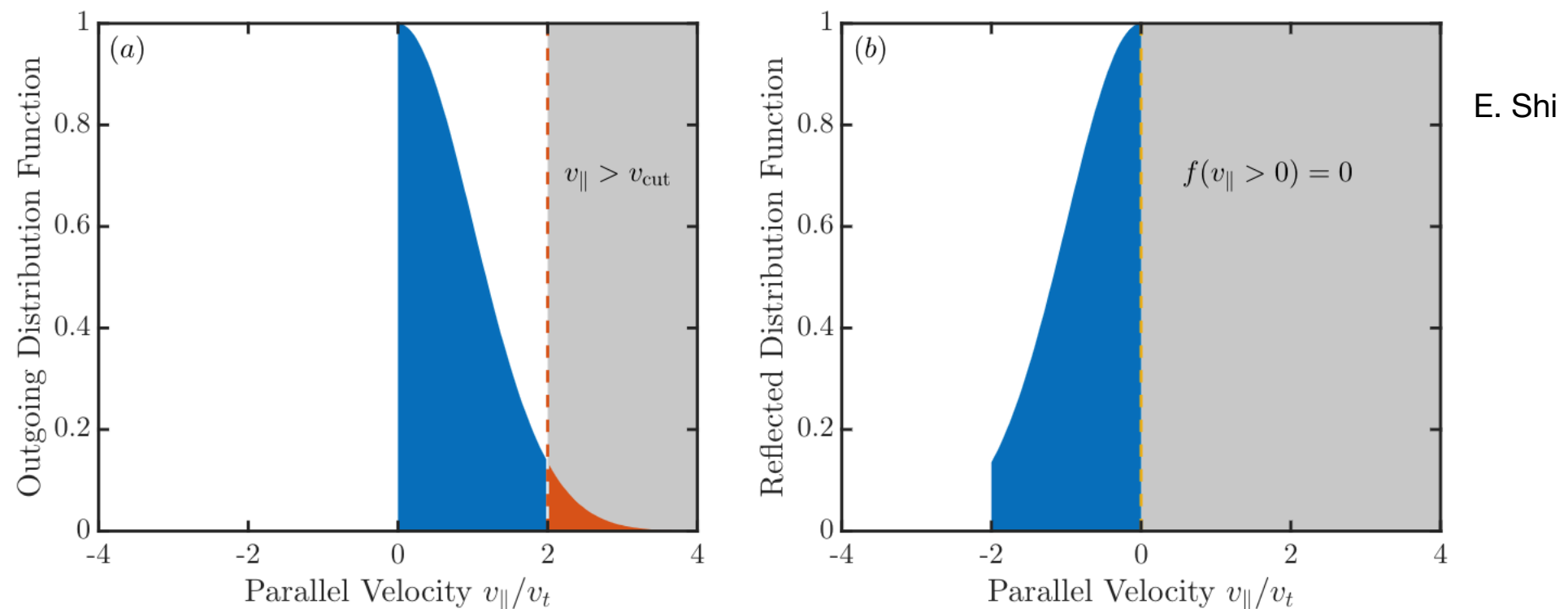


- Need to model non-neutral sheath using BCs (GK assumes quasi-neutrality, cannot resolve sheath)
- Sheath potential should reflect low energy electrons
- Solve Poisson equation on z boundary to get $\phi_{sh}(x, y) \doteq \phi(z = z_{sh})$, then use $\Delta\phi = \phi_{sh} - \phi_w$ to reflect electrons with $mv_{\parallel}^2/2 < |e|\Delta\phi$

$$-\nabla_{\perp} \cdot \sum_s \frac{m_s n_{0s}}{B^2} \nabla_{\perp} \phi(z = z_{sh}) = \sum_s q_s \int d^3v f_s(z = z_{sh})$$

- Potential self-consistently relaxes to ambipolar-parallel-outflow state, and allows local currents in and out of wall (unlike “logical” sheath model)

Sheath boundary condition for electrons

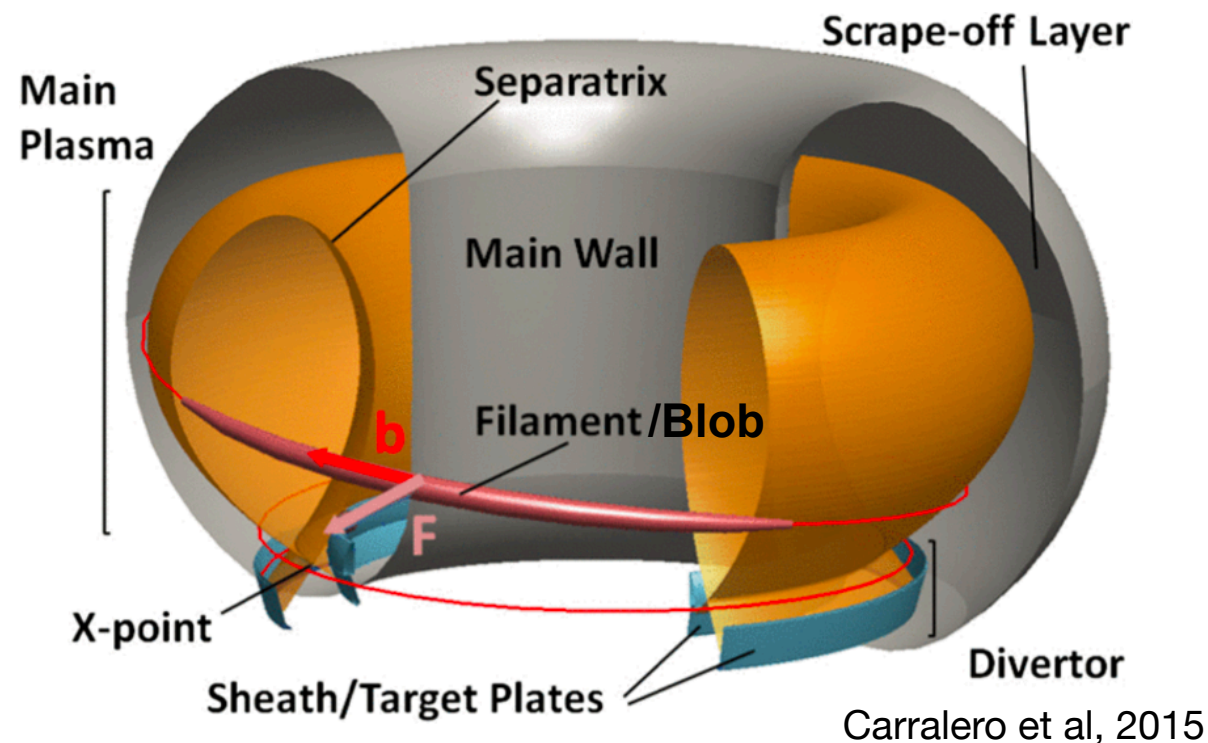


(a) Outgoing electrons with $v_{\parallel} > v_{cut} = \sqrt{2e\Delta\phi/m}$ are lost into the wall

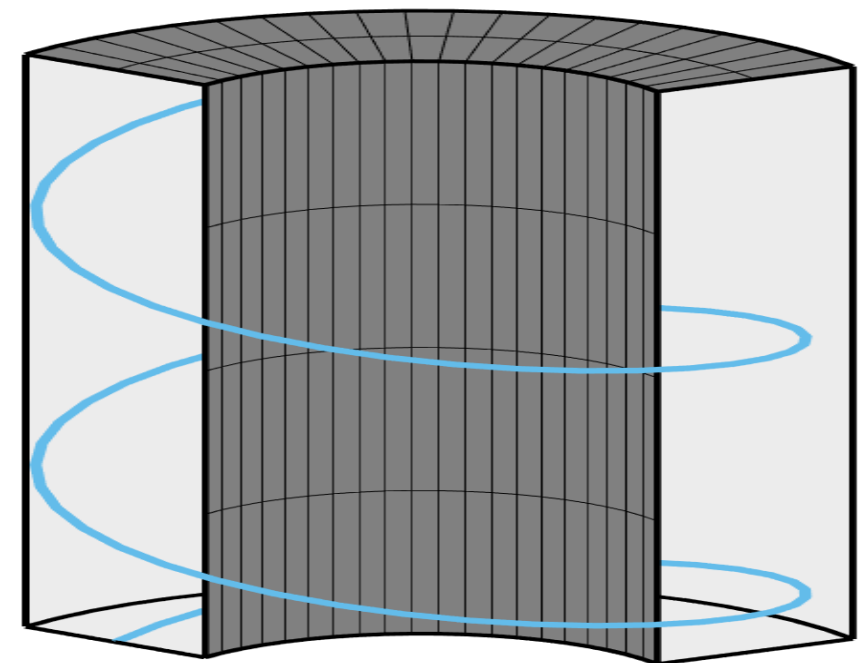
(b) Rest of outgoing electrons $0 < v_{\parallel} < v_{cut}$ are reflected back into plasma

Ions: Assuming positive sheath potential (relative to wall), all ions are lost

Modeling the NSTX SOL with Gkeyll

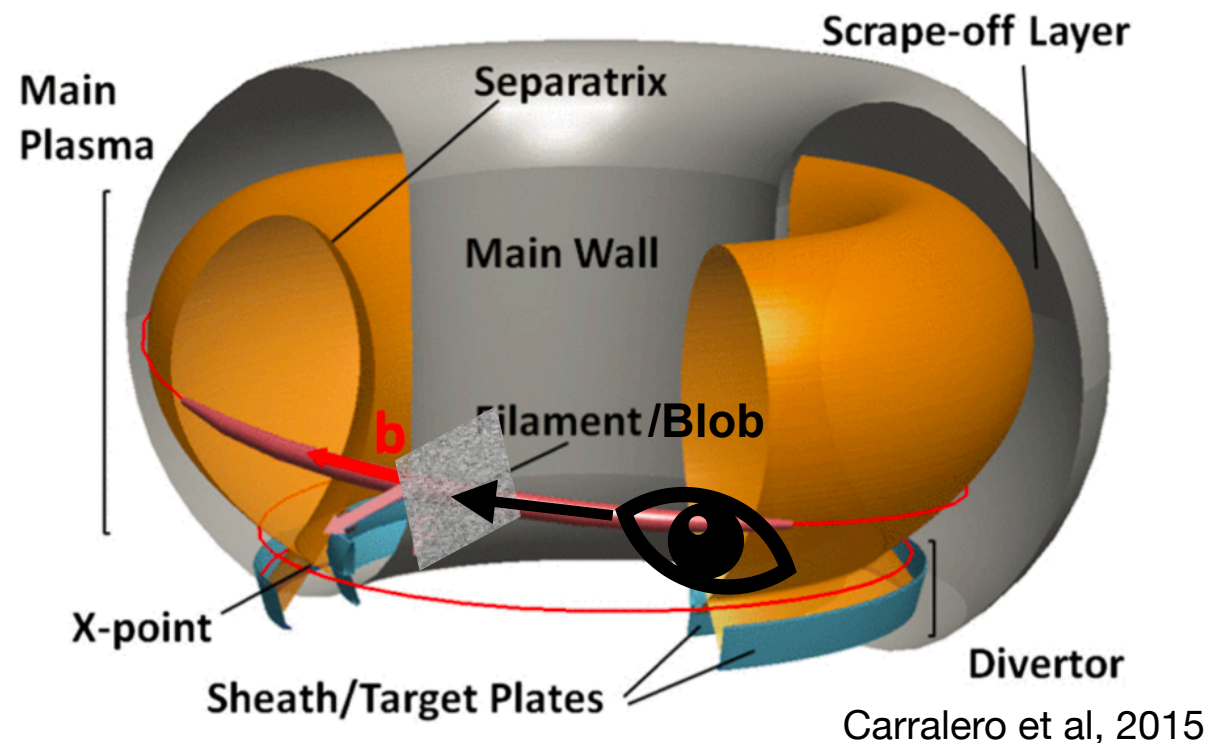


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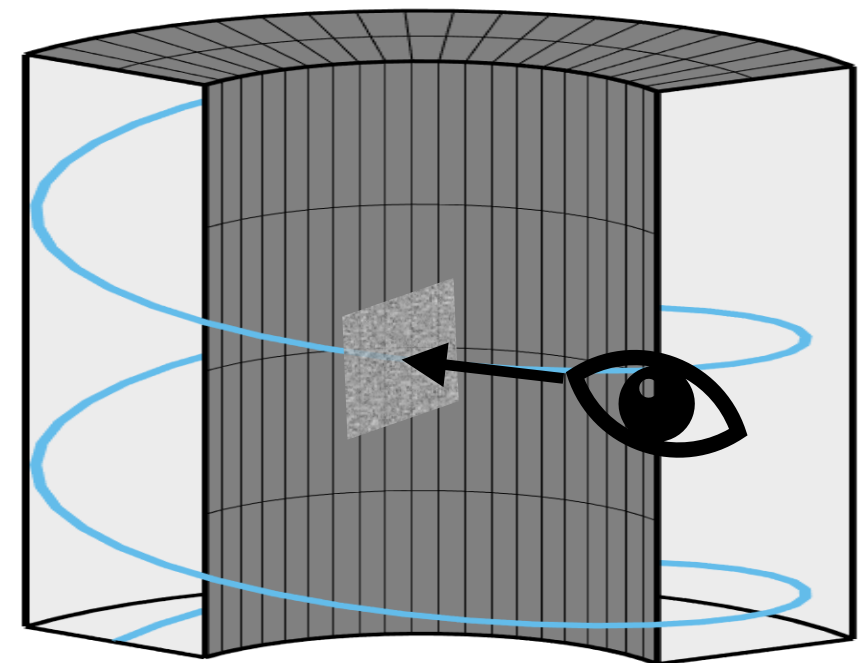


- Simple helical model of NSTX SOL
 - Field-aligned simulation domain that follows field lines from bottom divertor plate, around the torus, to the top divertor plate
 - Length along field line \sim connection length $L_{\parallel} = 8$ m (constant, no shear for now)
 - All bad curvature \rightarrow interchange instability, blob dynamics
 - Real deuterium mass ratio, Lenard-Bernstein collisions

Modeling the NSTX SOL with Gkeyll



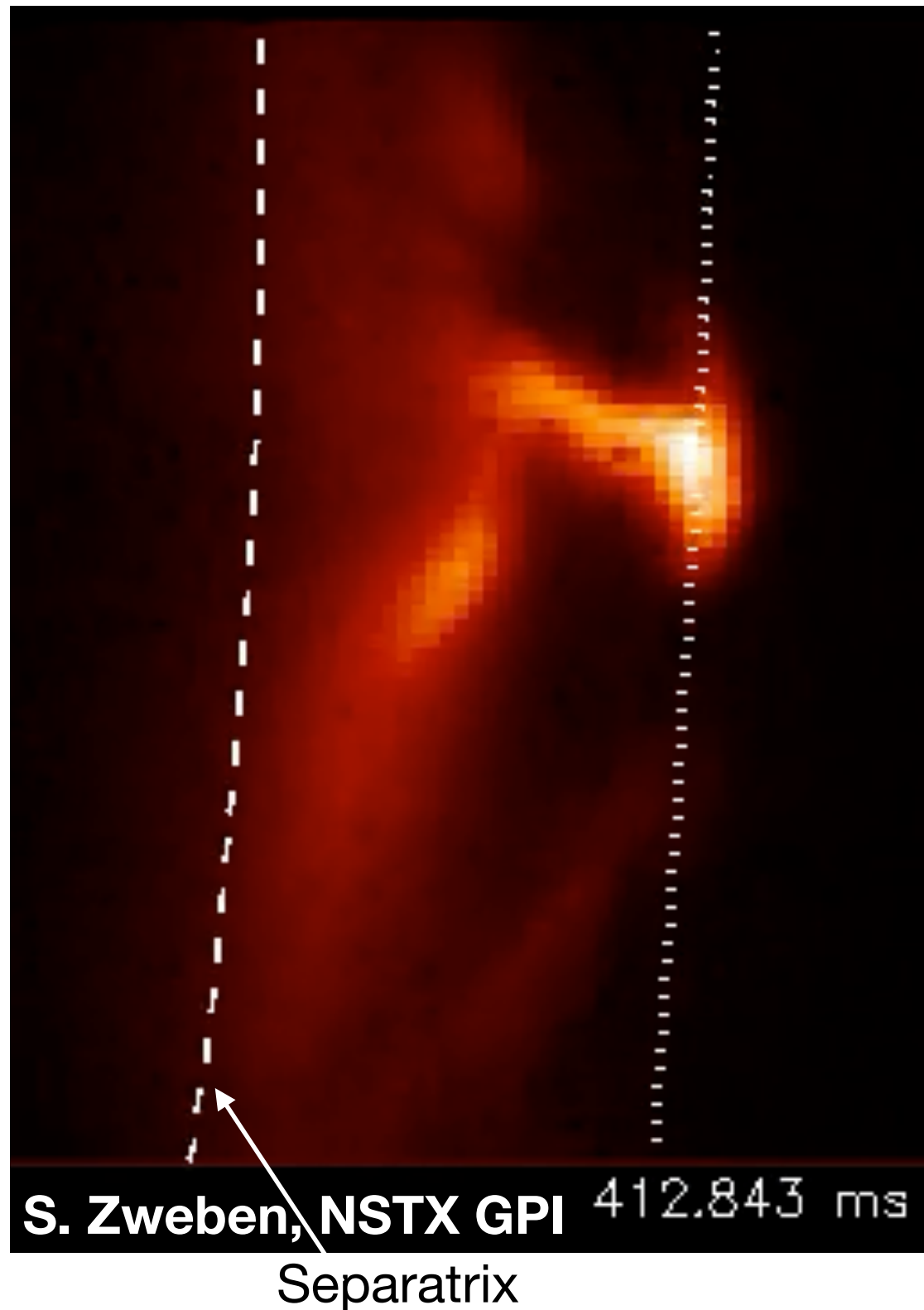
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NSTX GPI

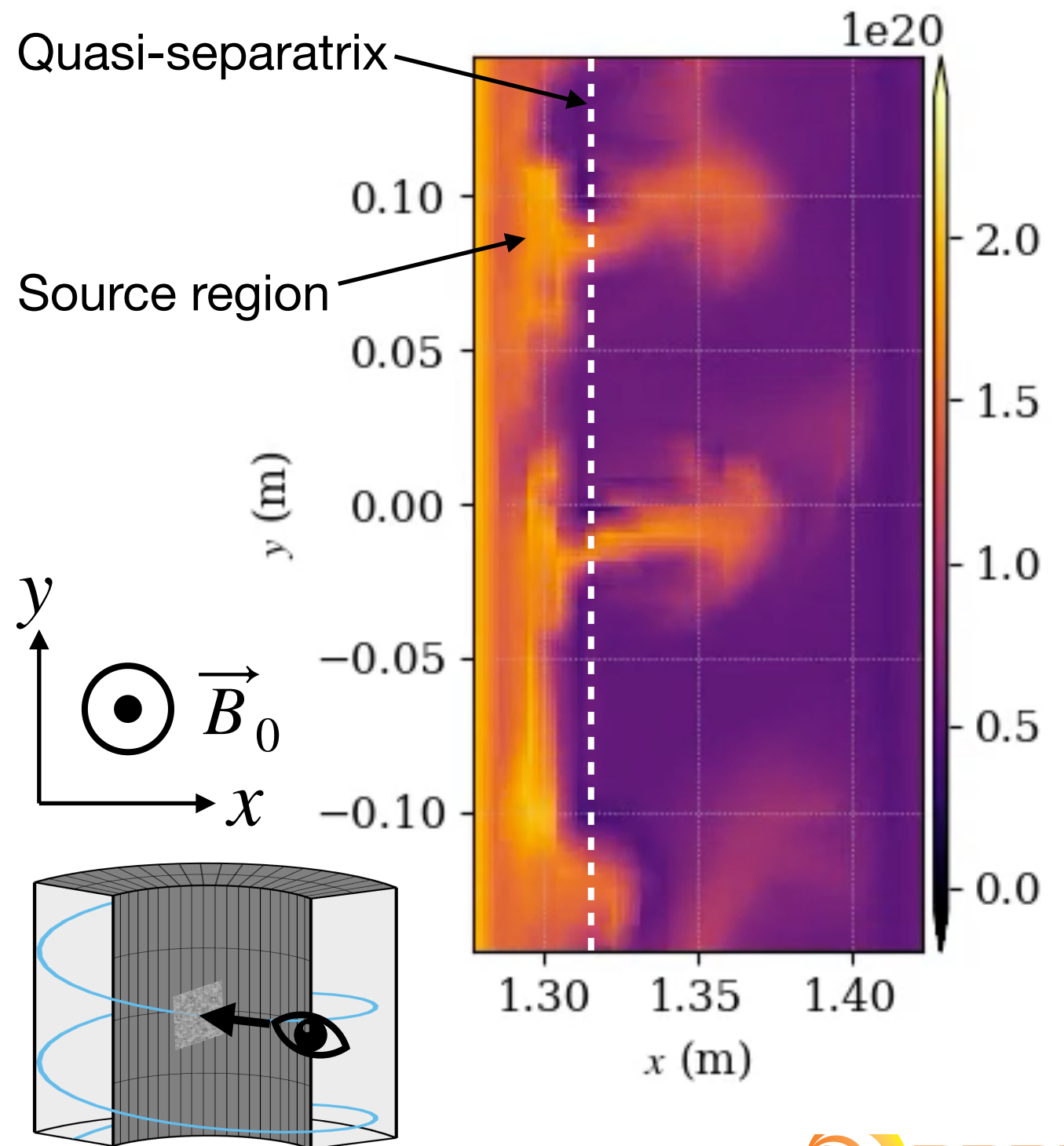
$D\alpha$ signal \sim density at midplane



vs Gkeyll

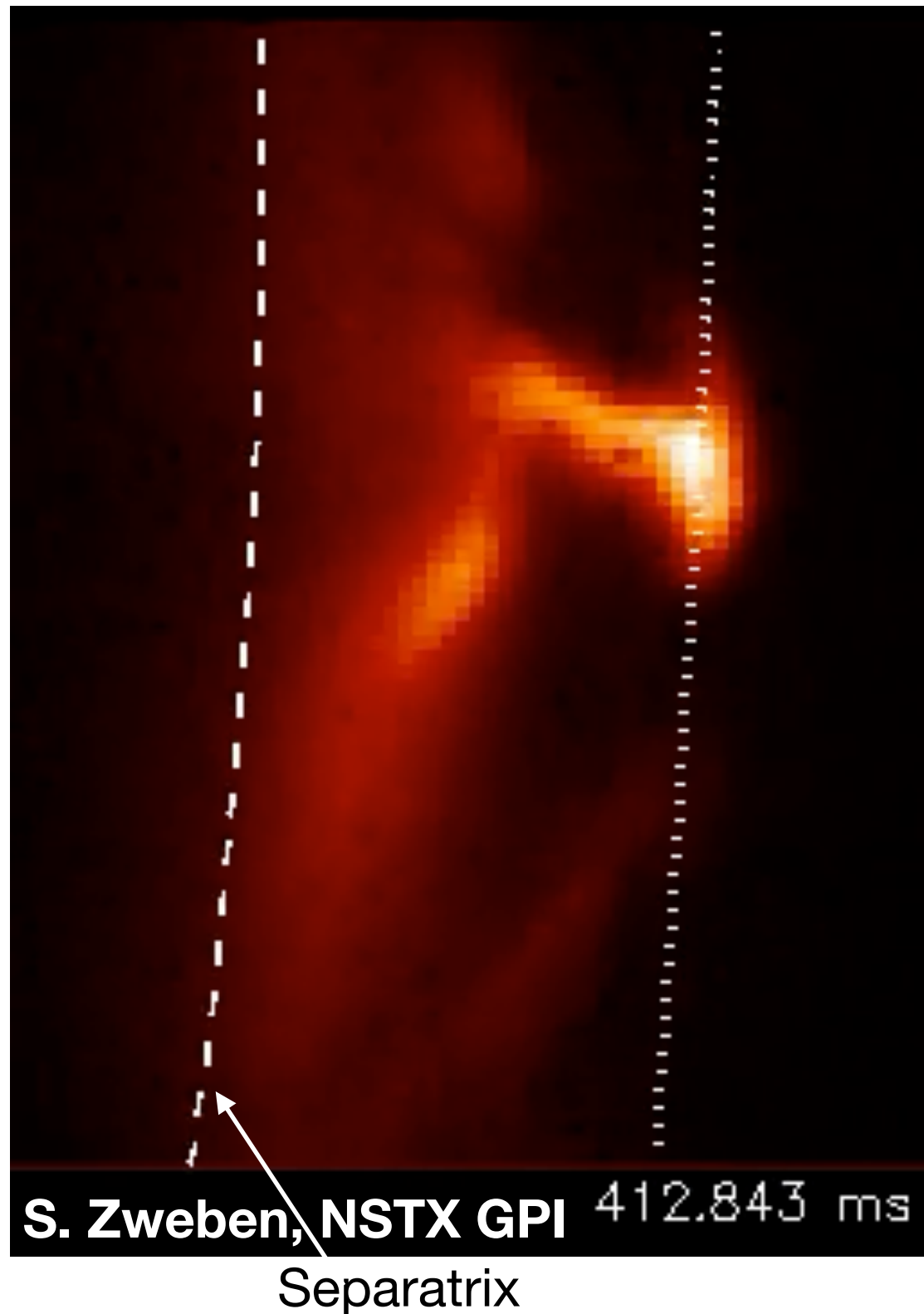
Ion density at midplane

F: 514 T: 5.1400e-04



NSTX GPI

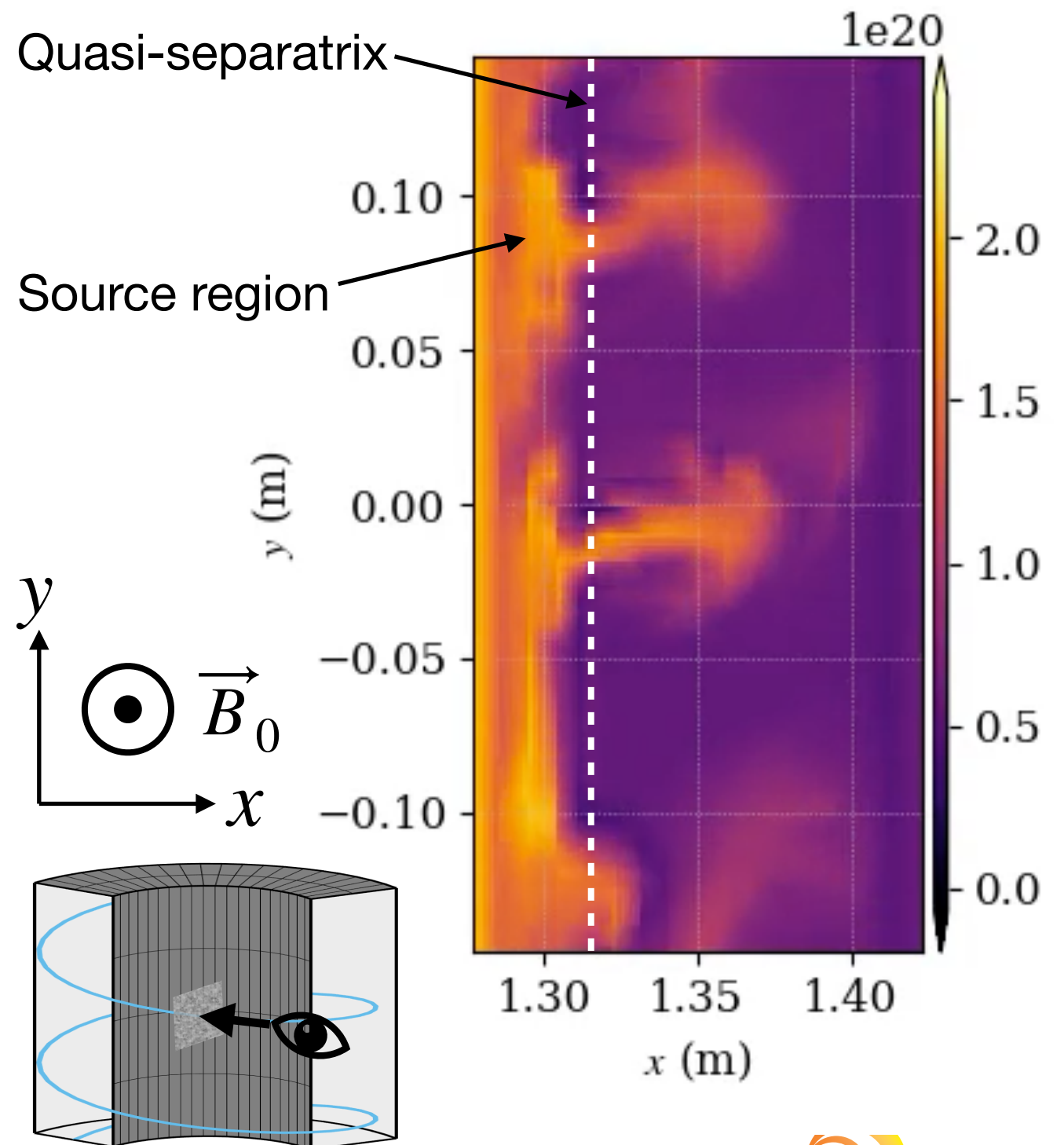
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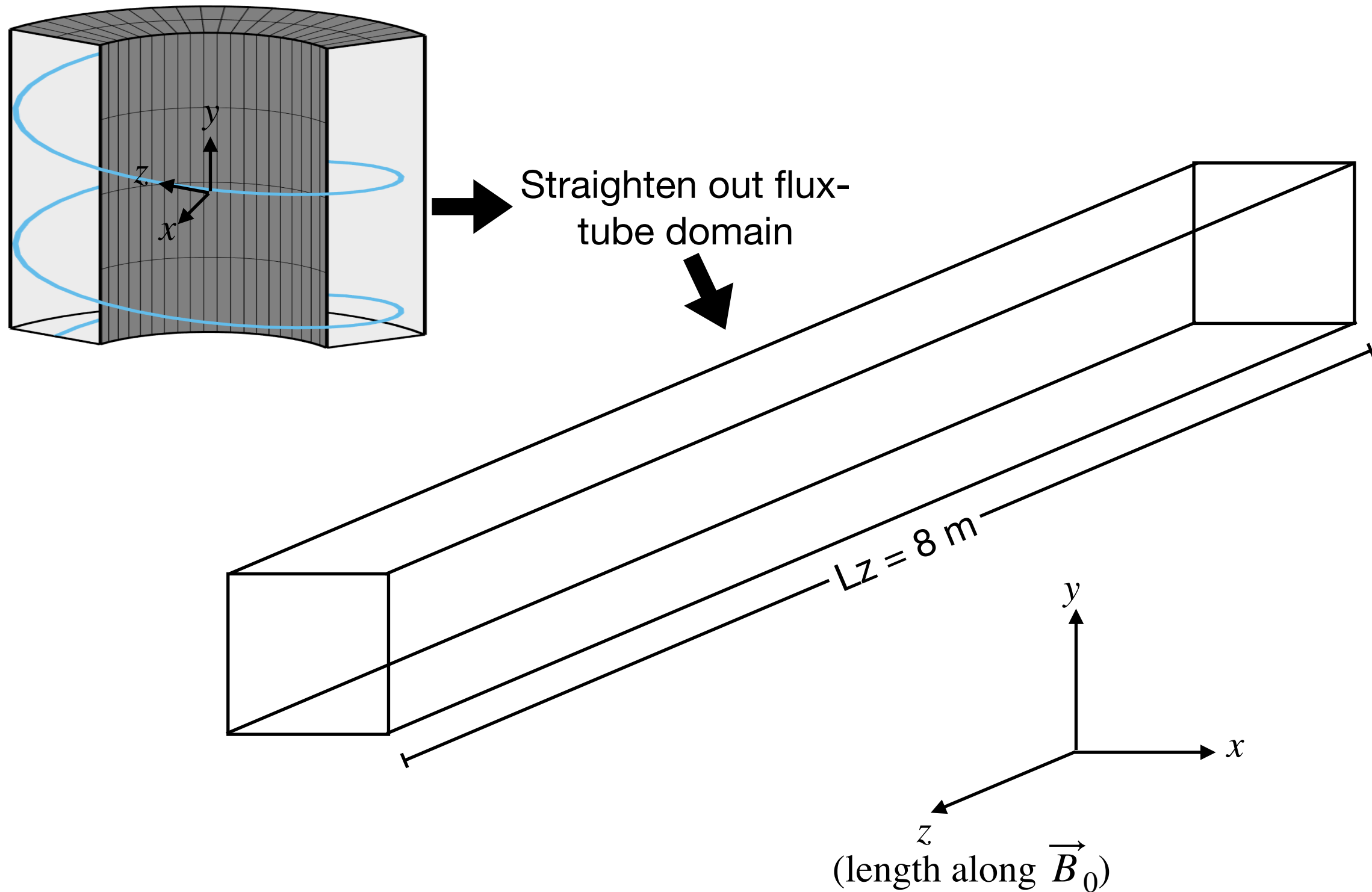
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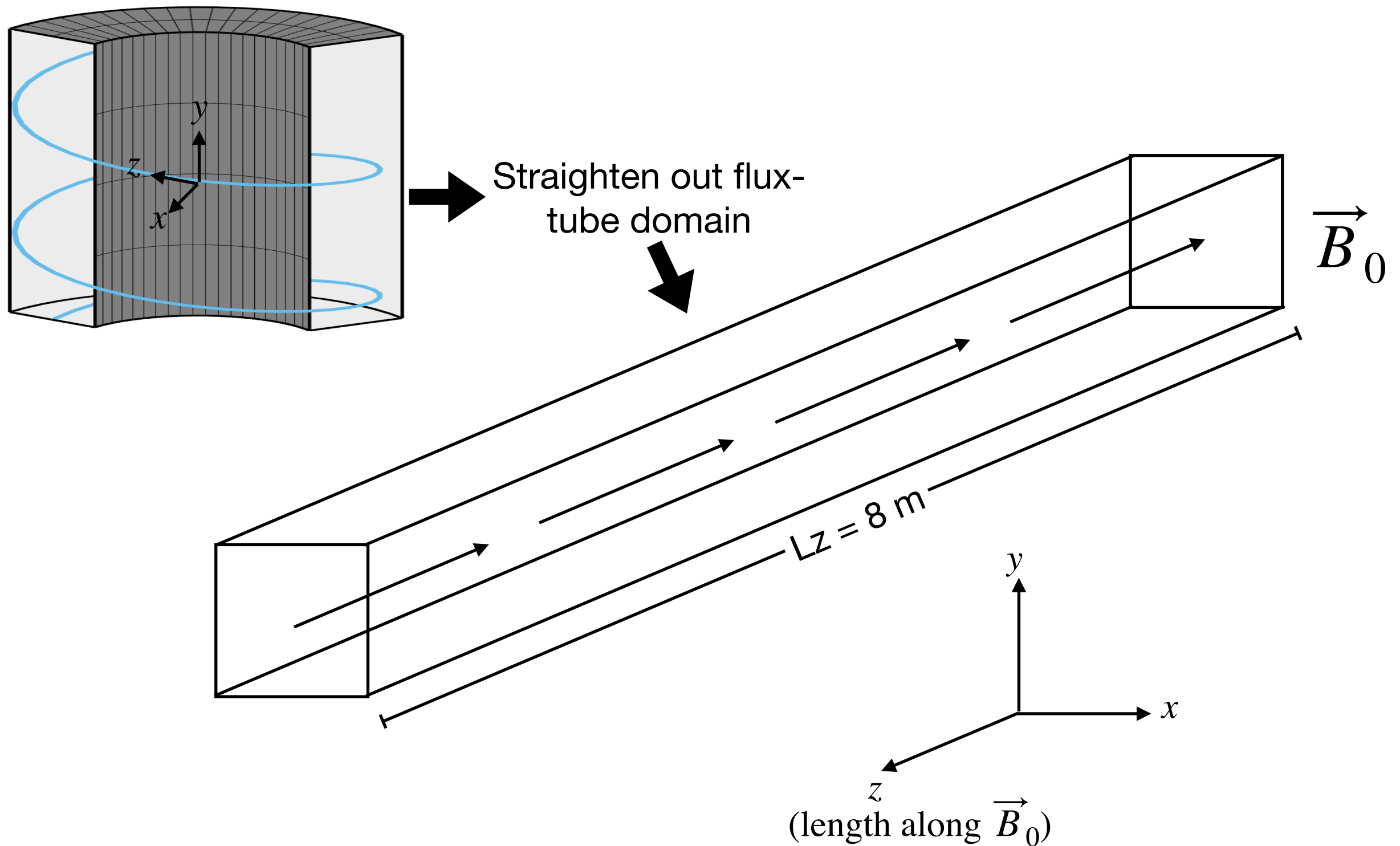
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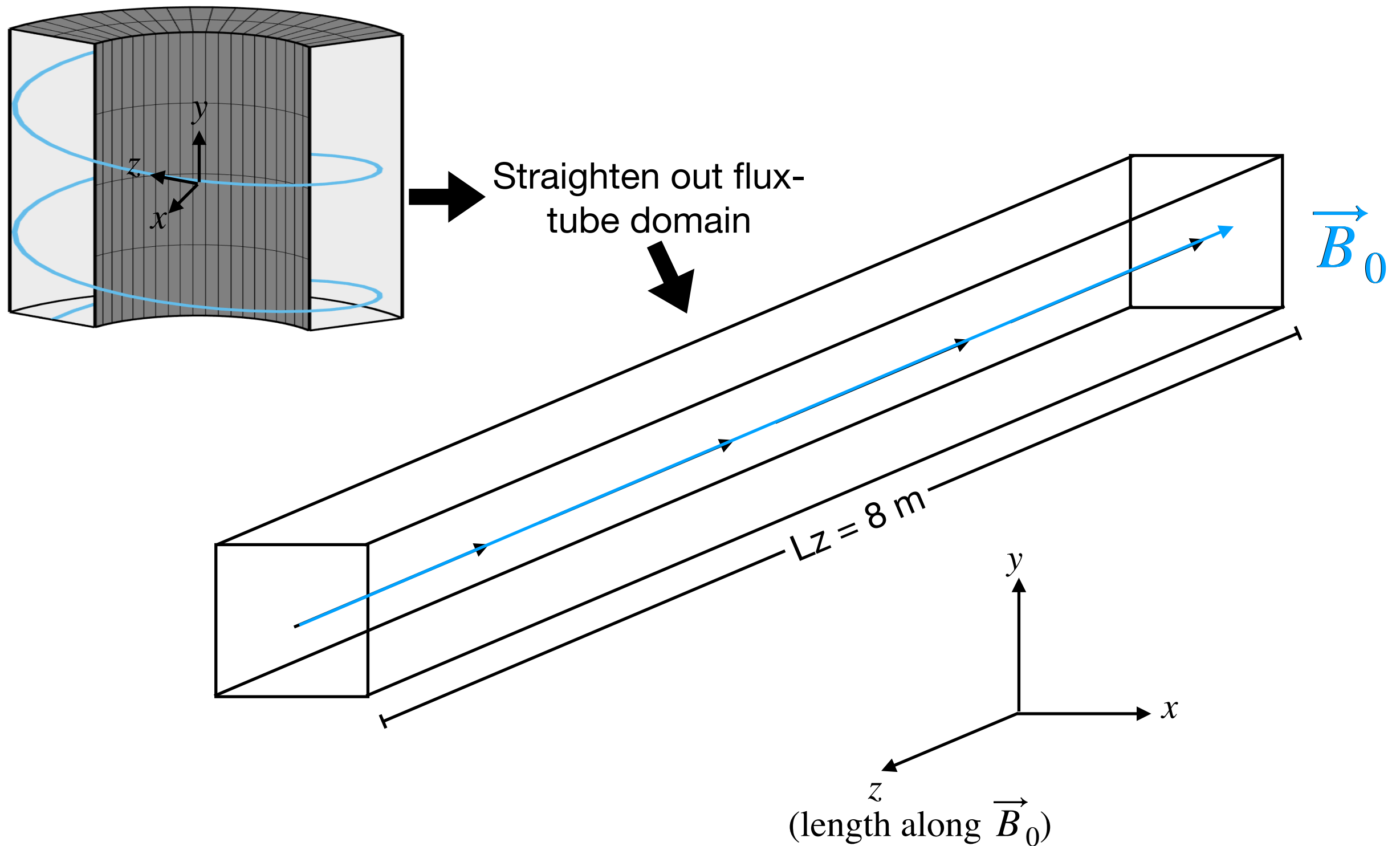
Demonstrating Gkeyll's EMGK capabilities: tracing magnetic field lines



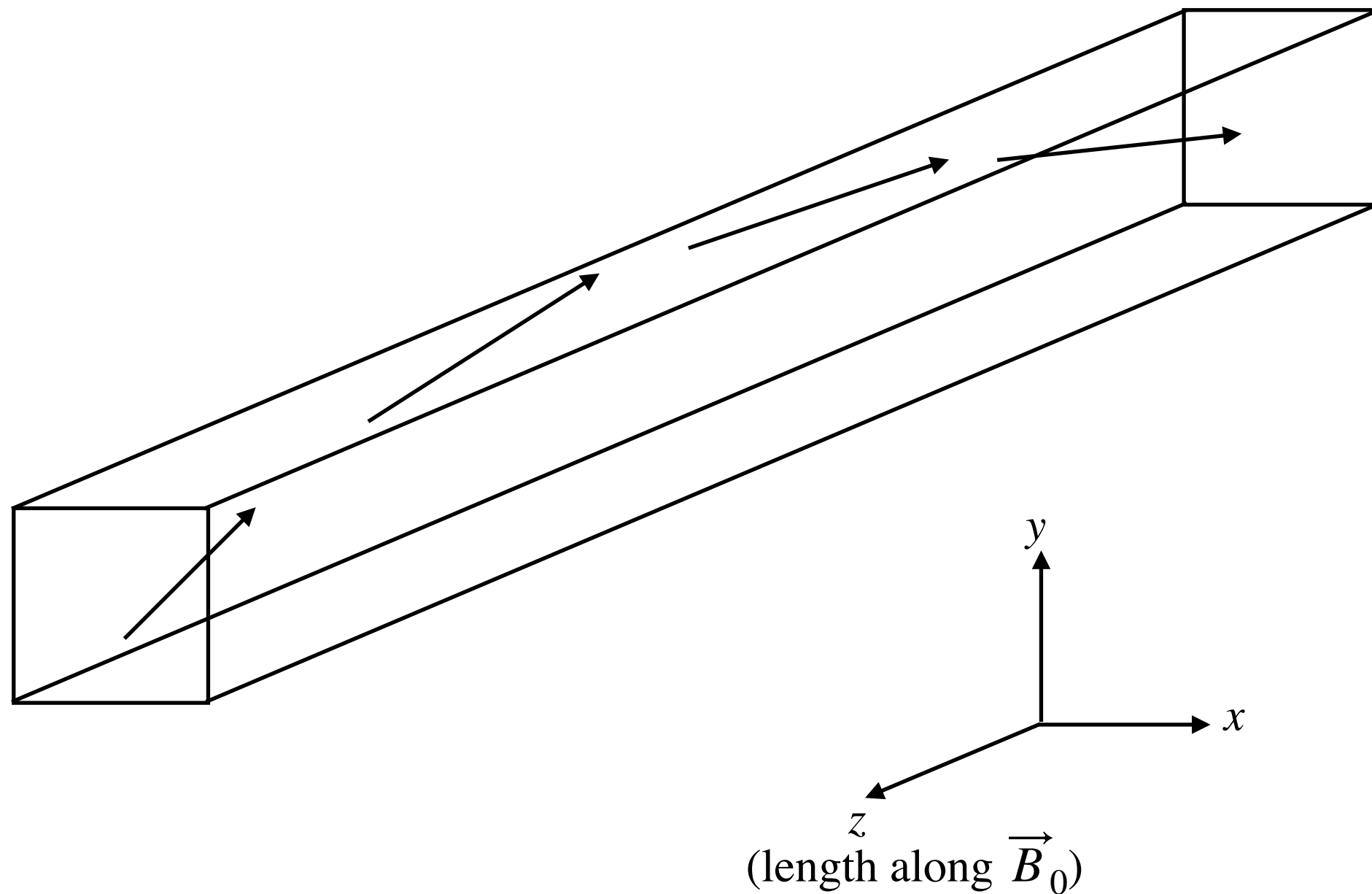
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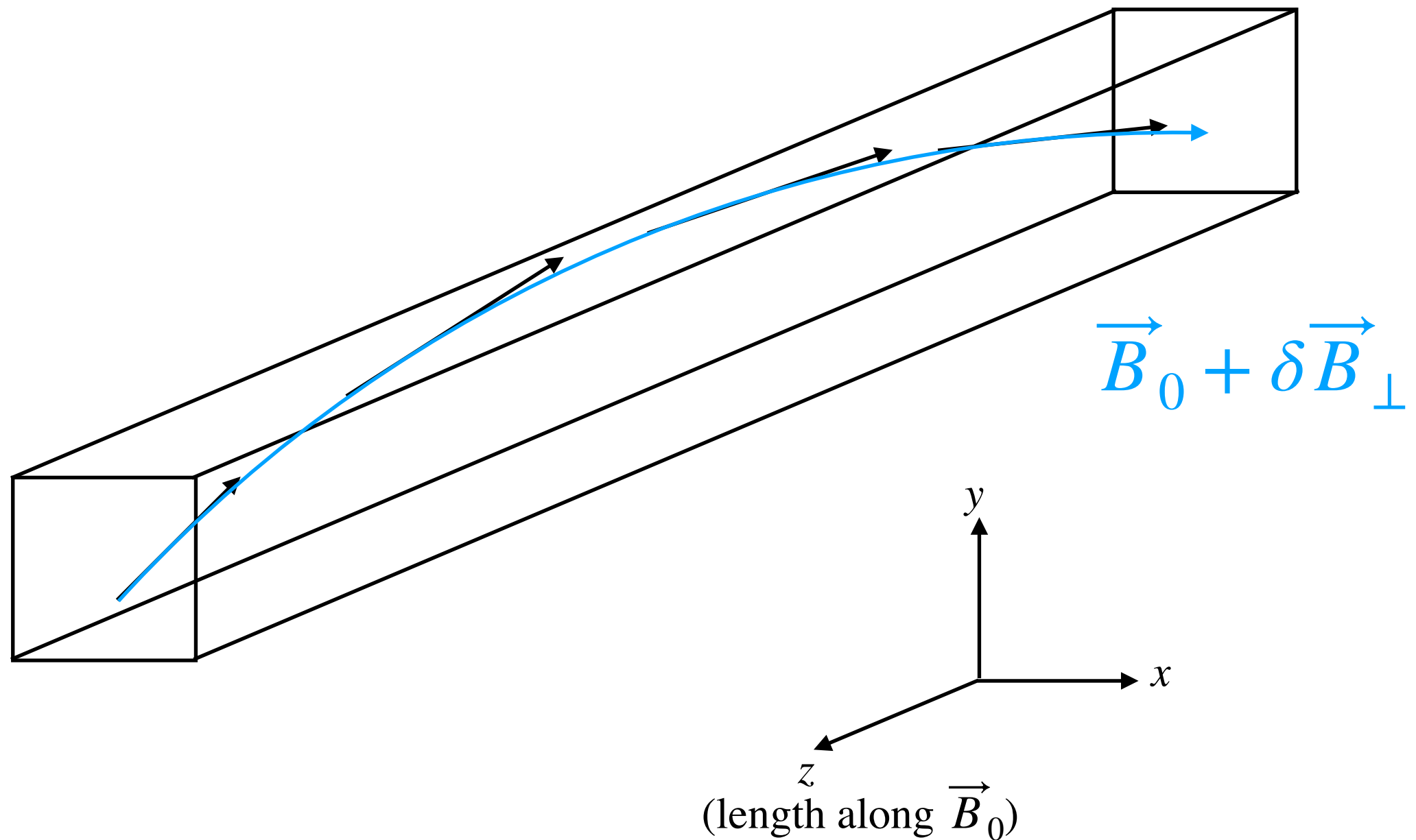
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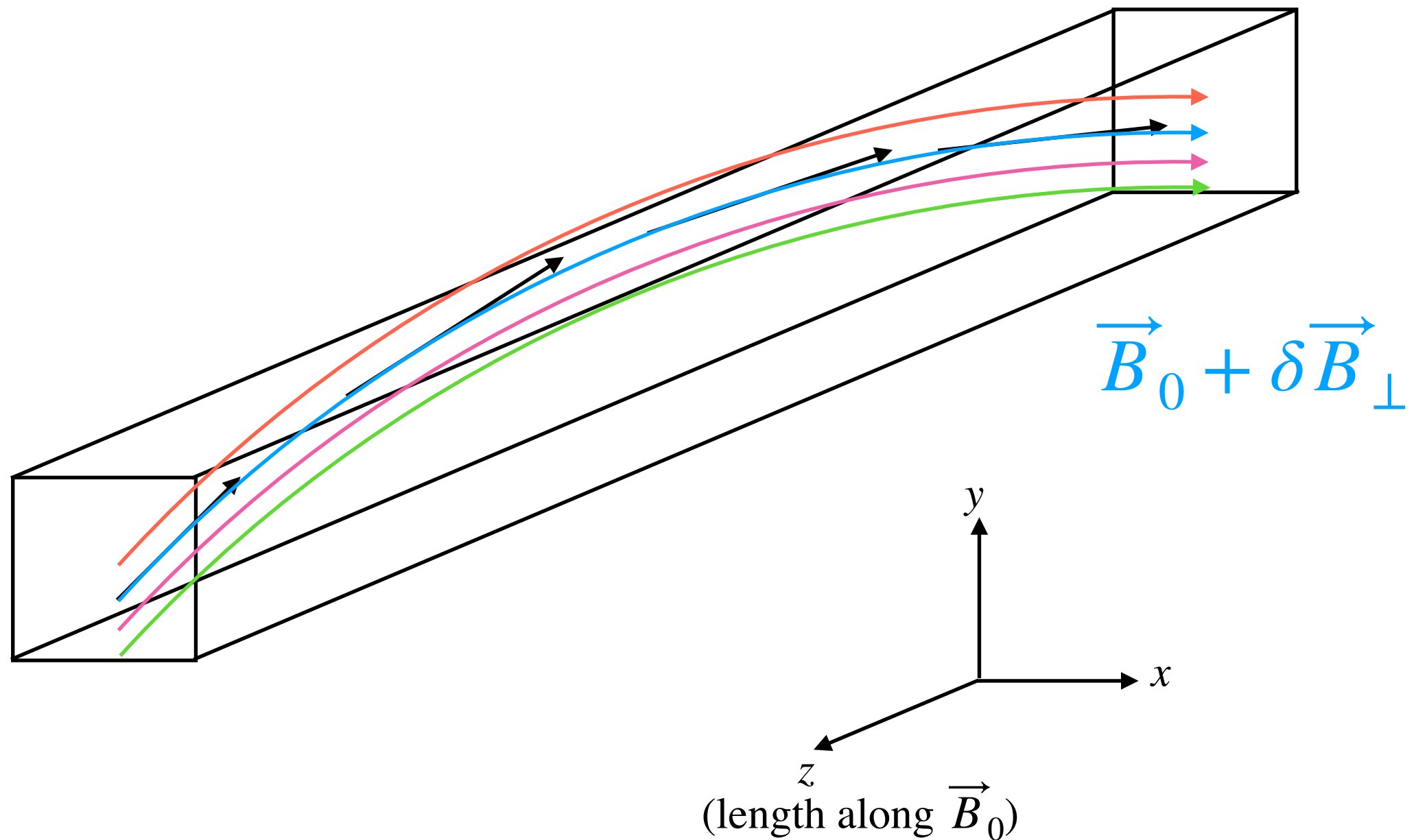
Demonstrating Gkeyll's EMGK capabilities: tracing *fluctuating* magnetic field lines



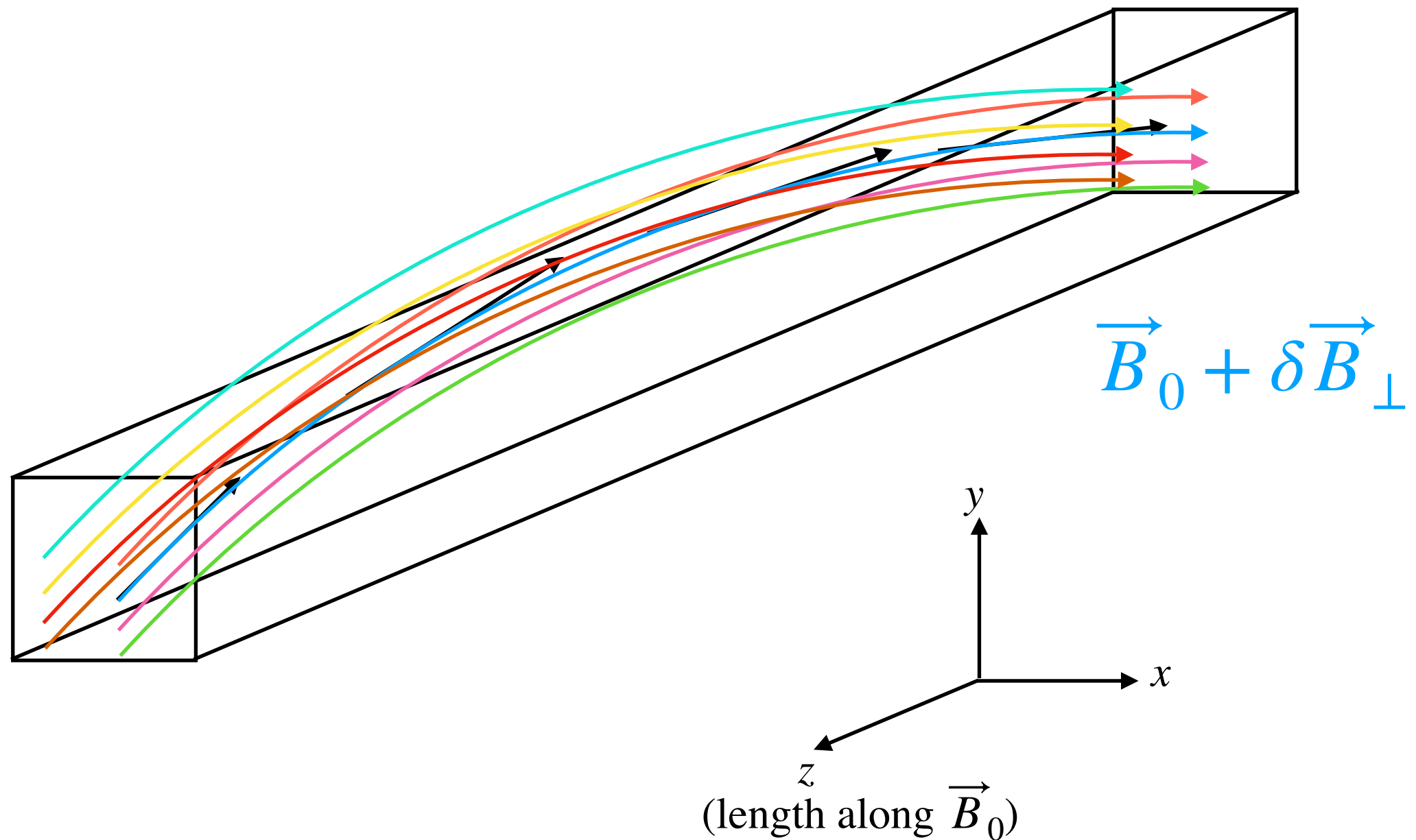
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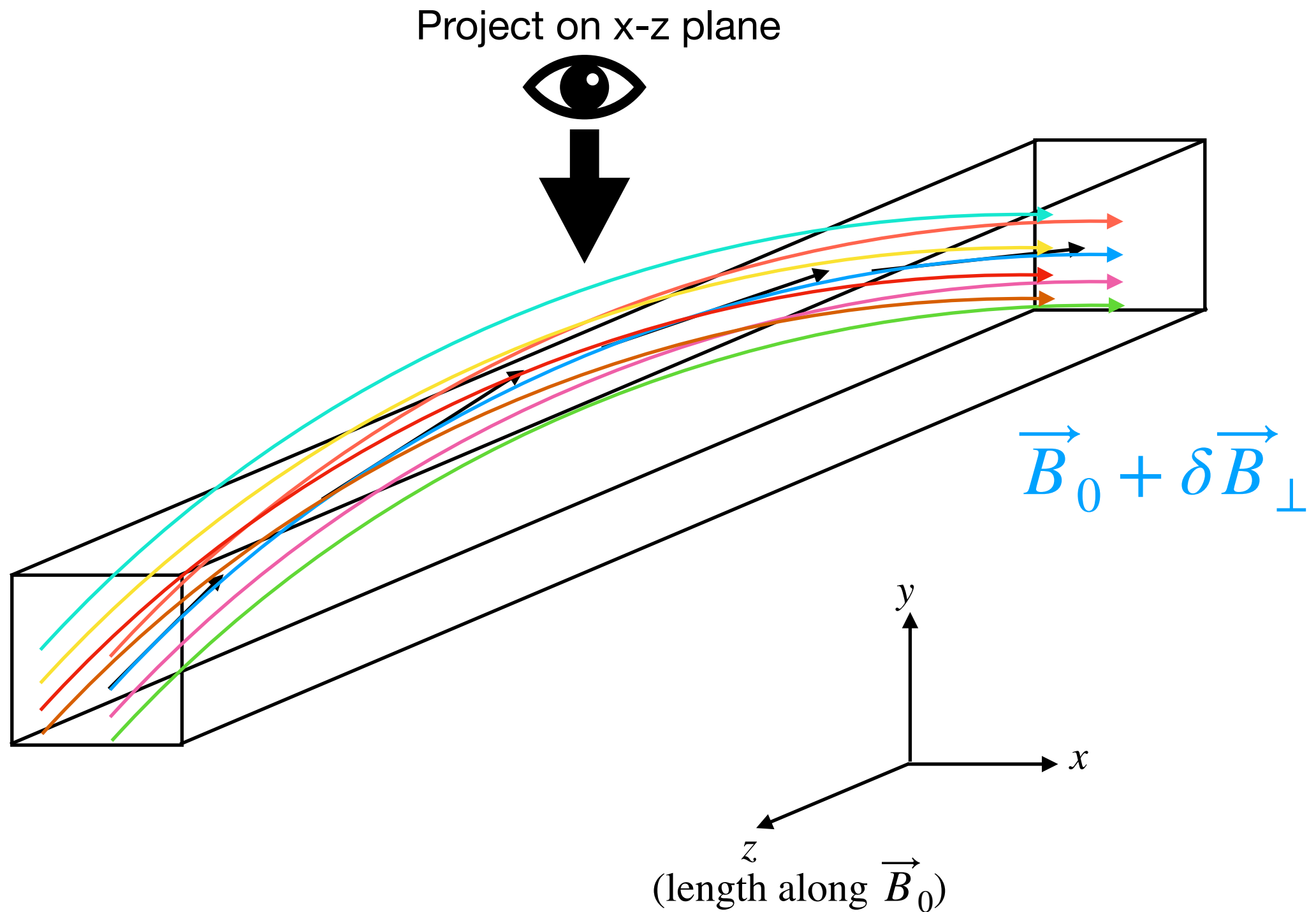
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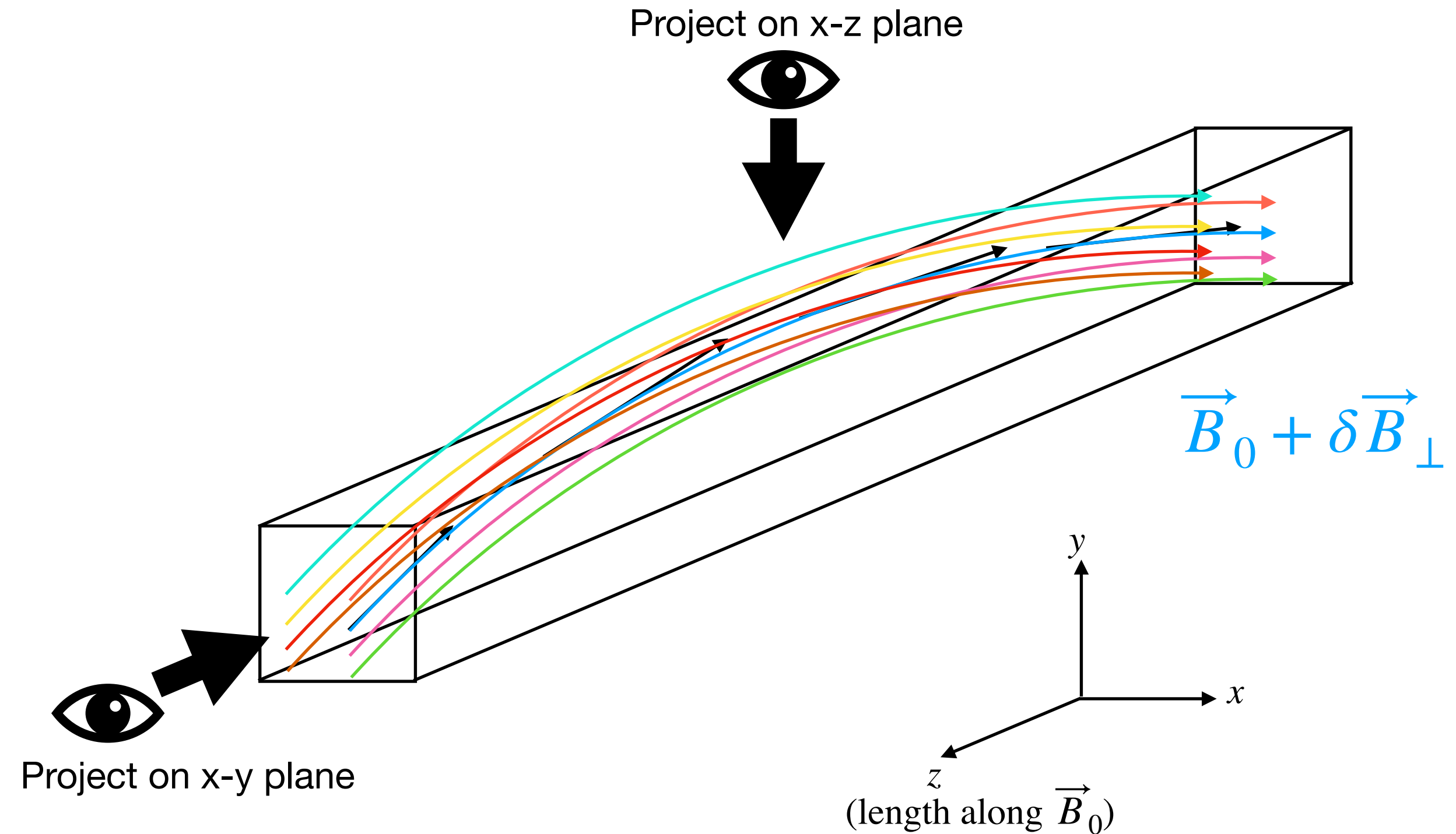
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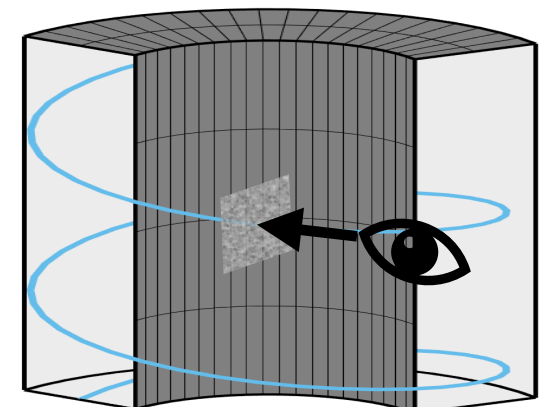
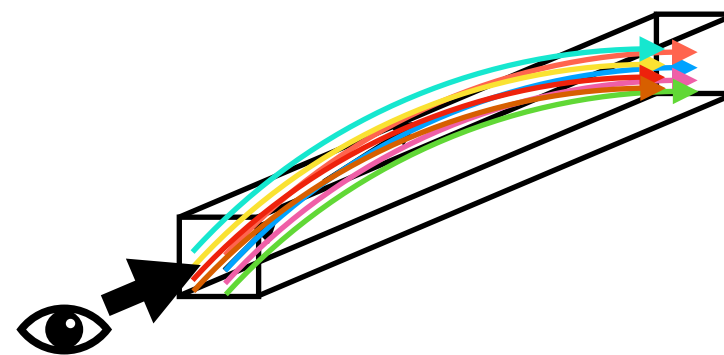
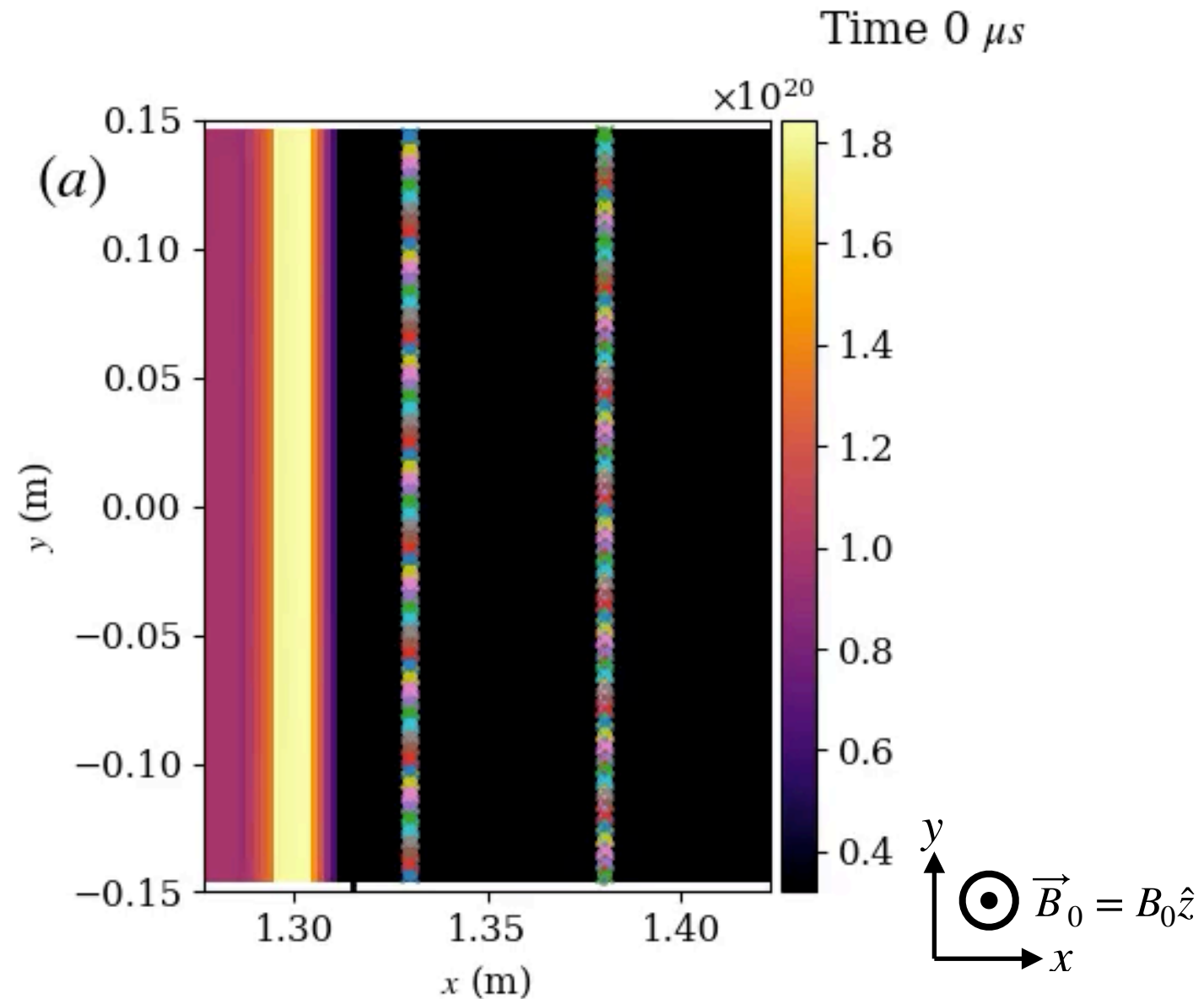
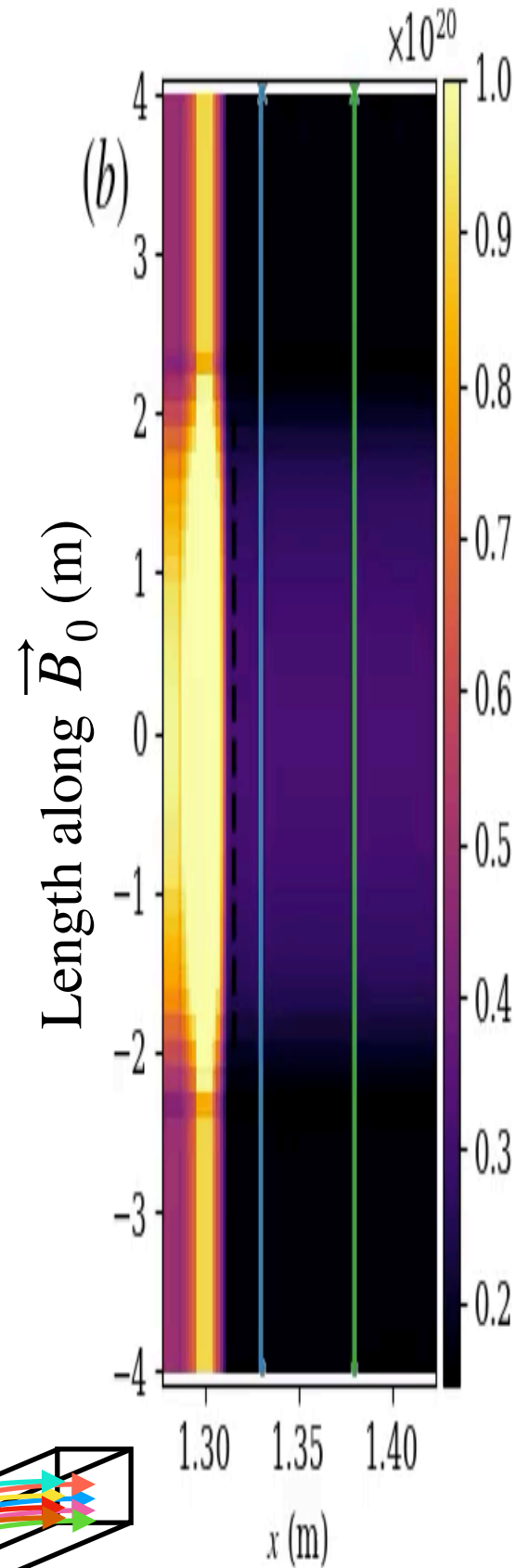
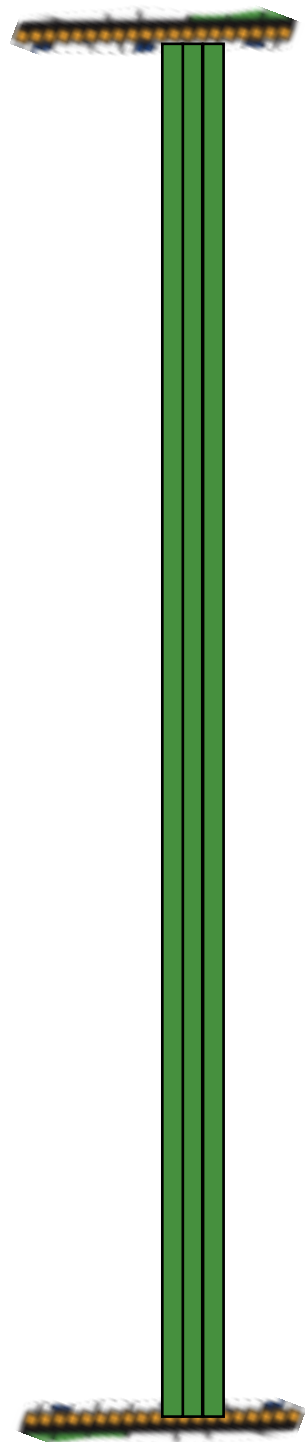


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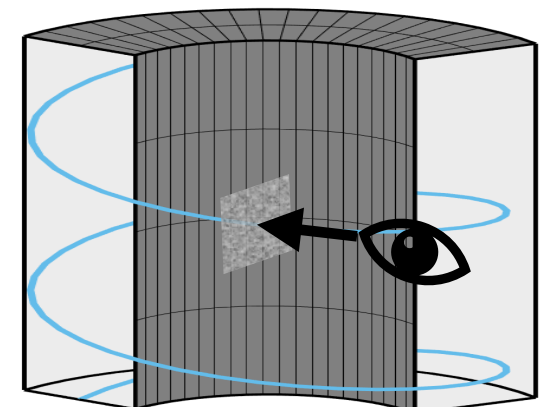
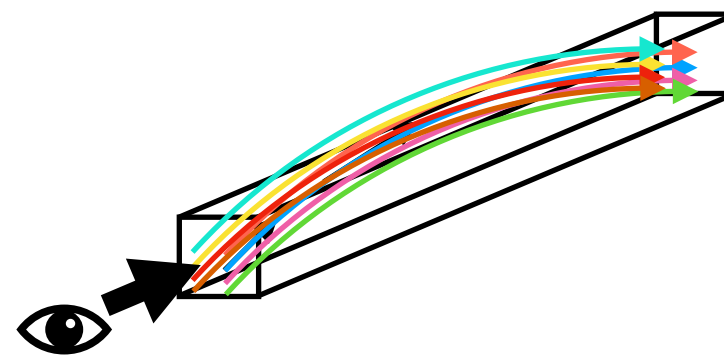
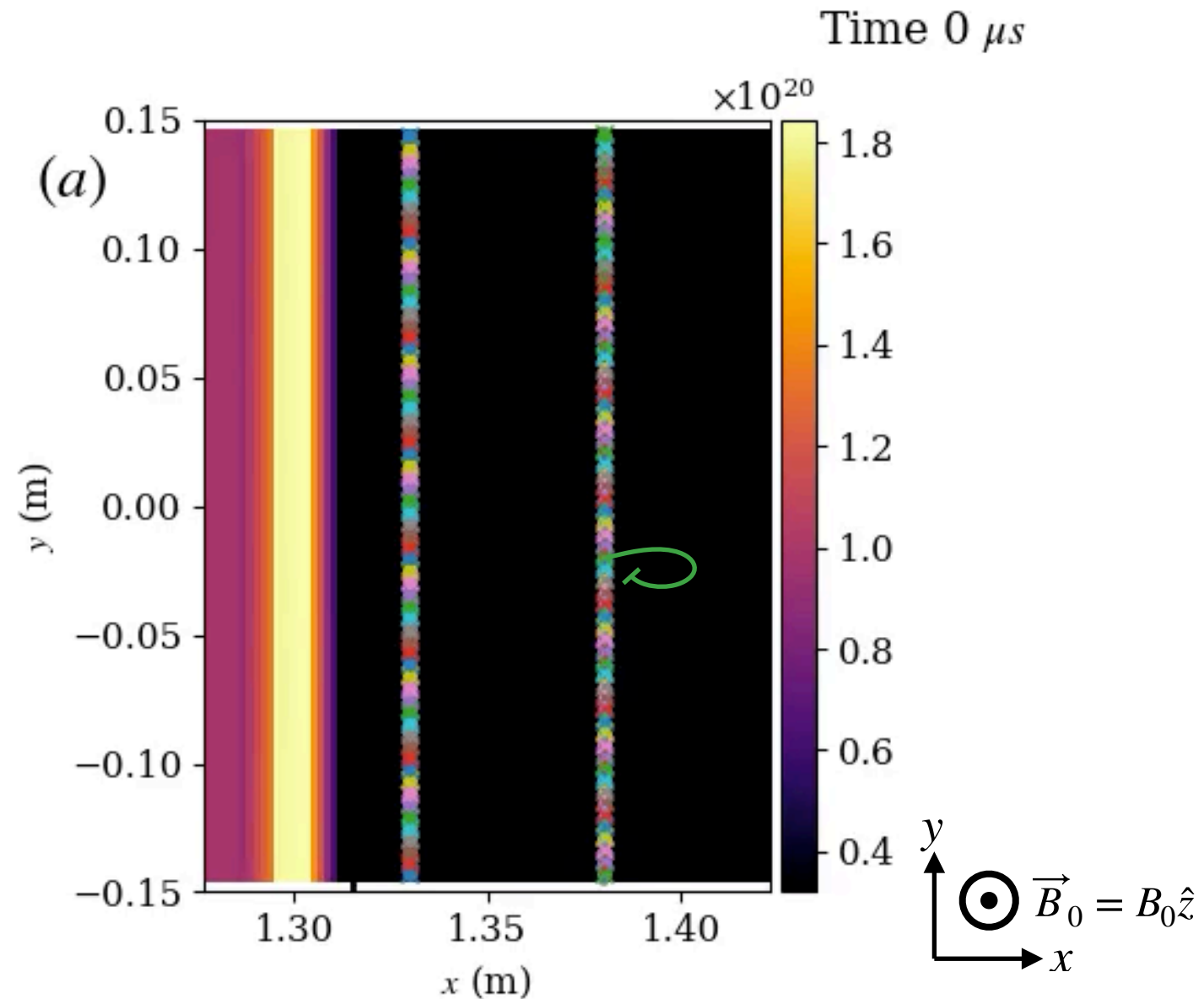
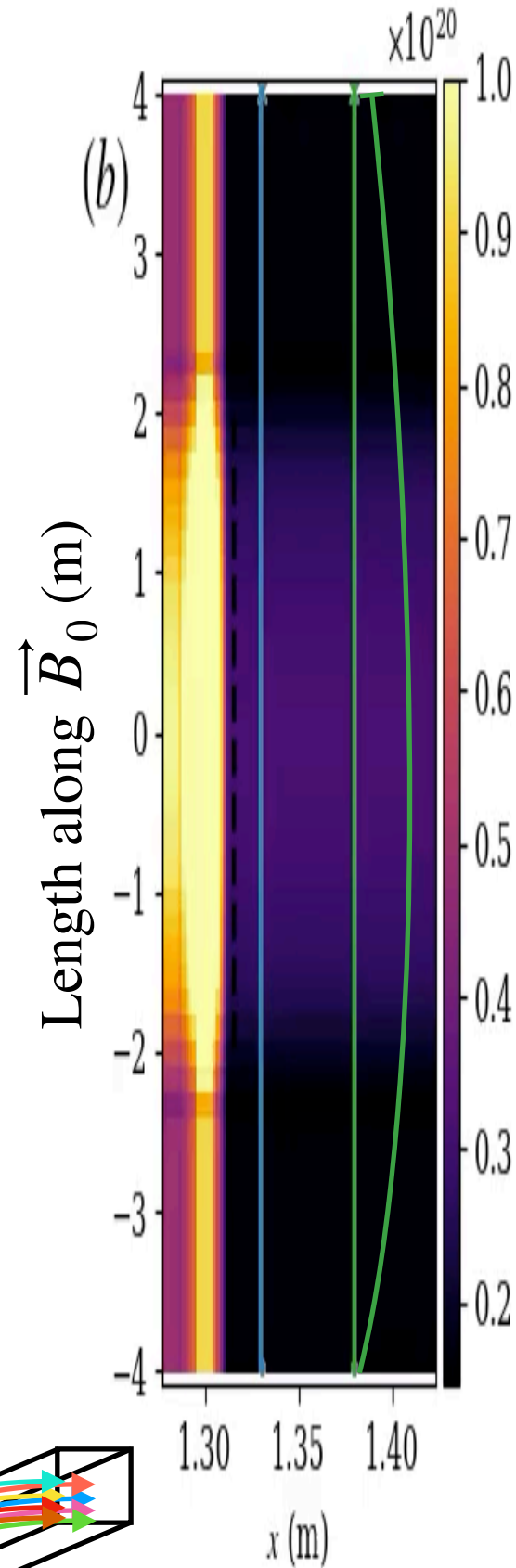
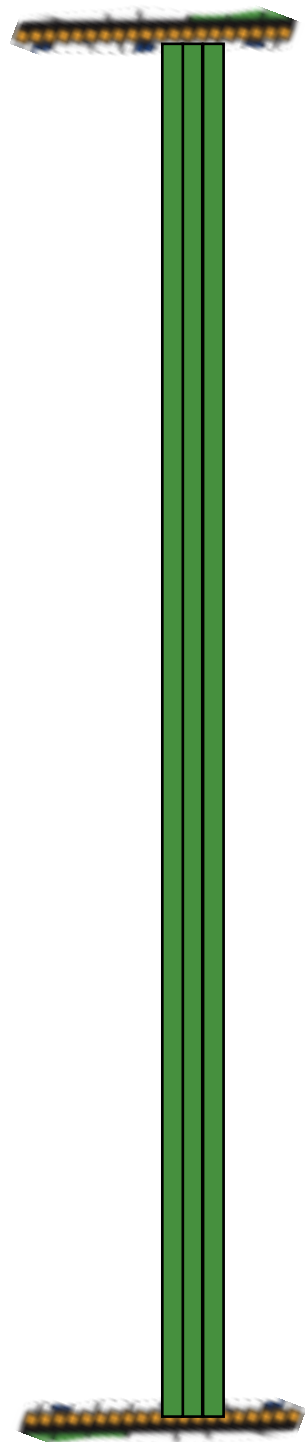


Demonstrating Gkeyll's EMGK capabilities:

dancing field lines

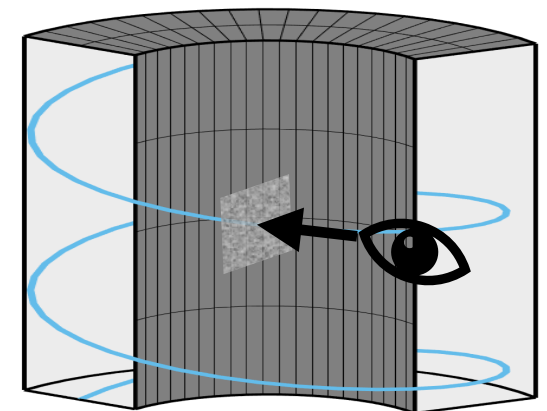
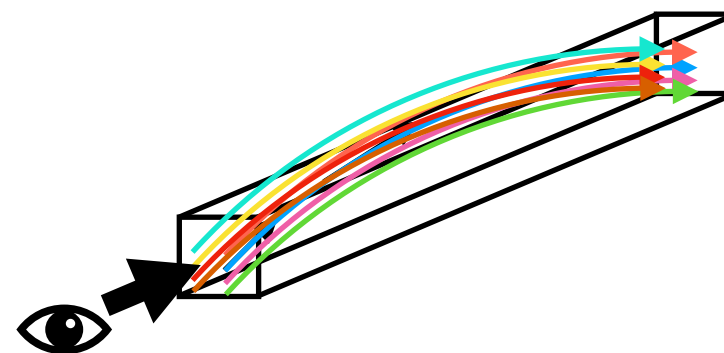
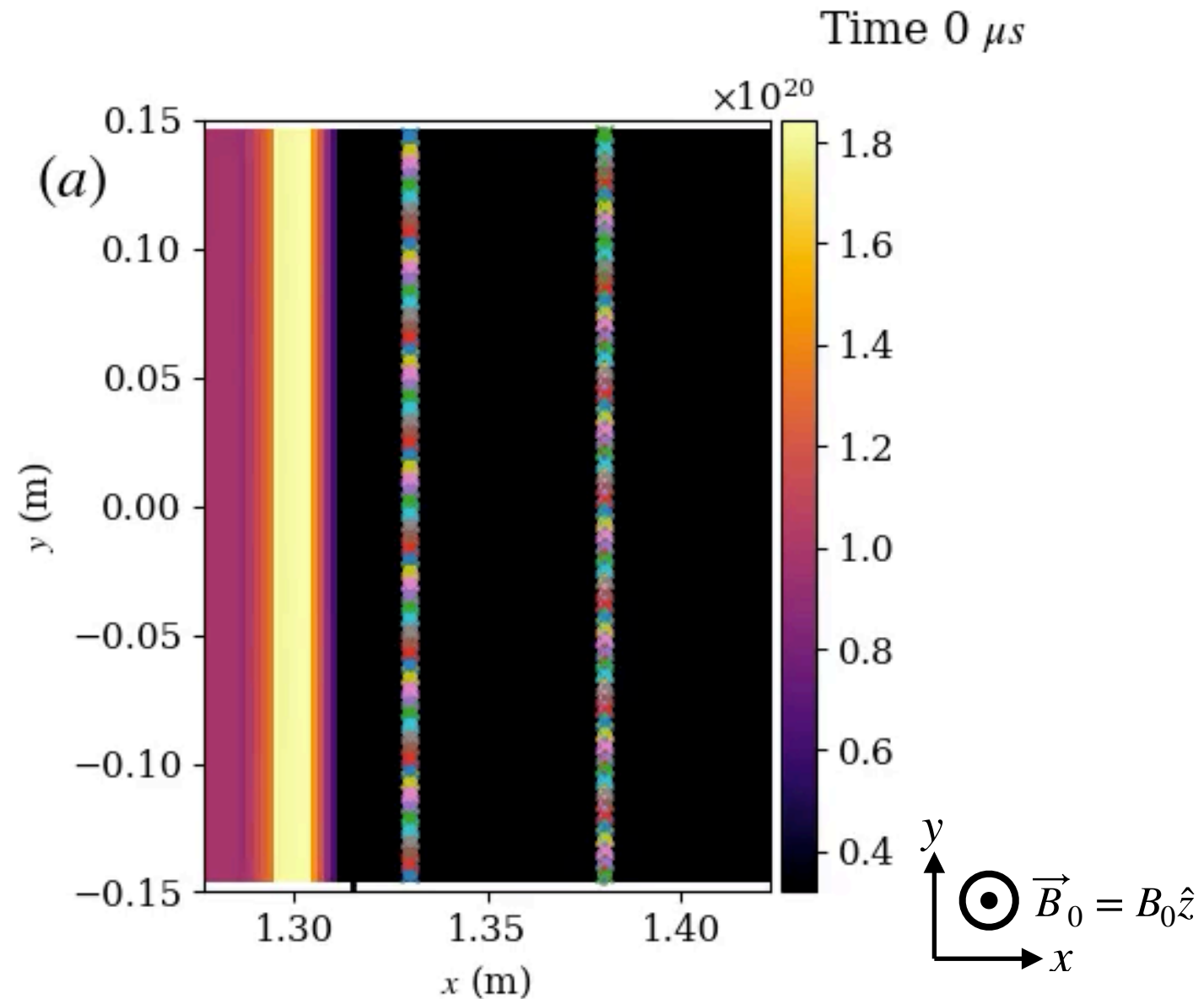
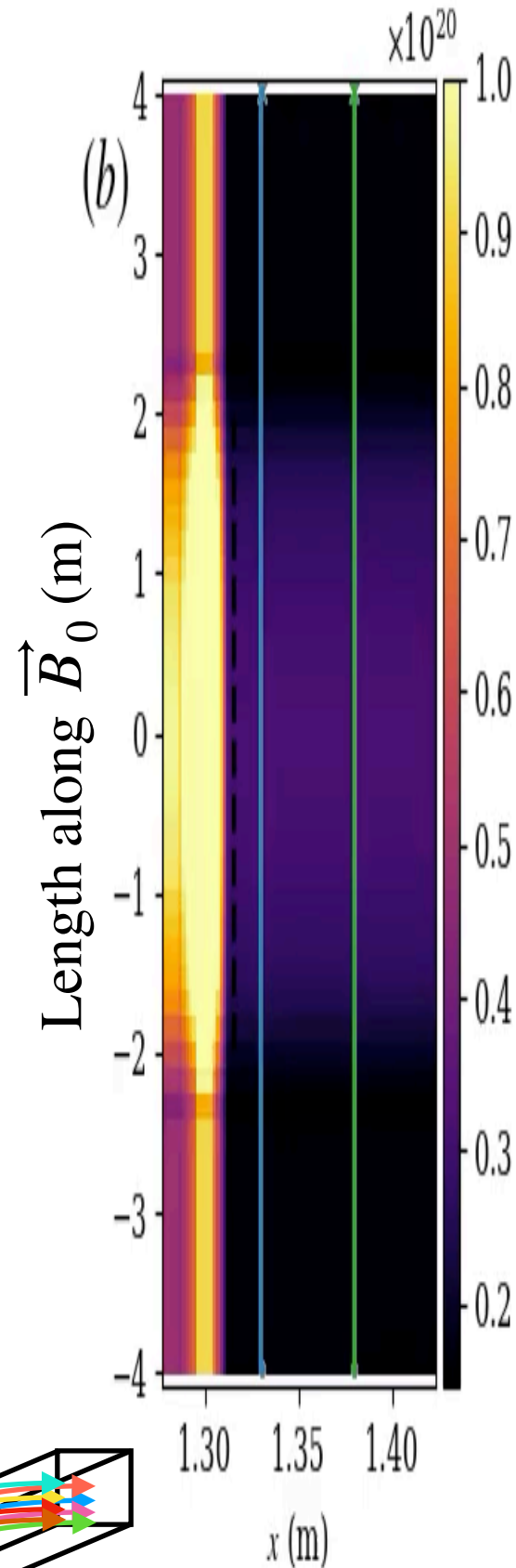
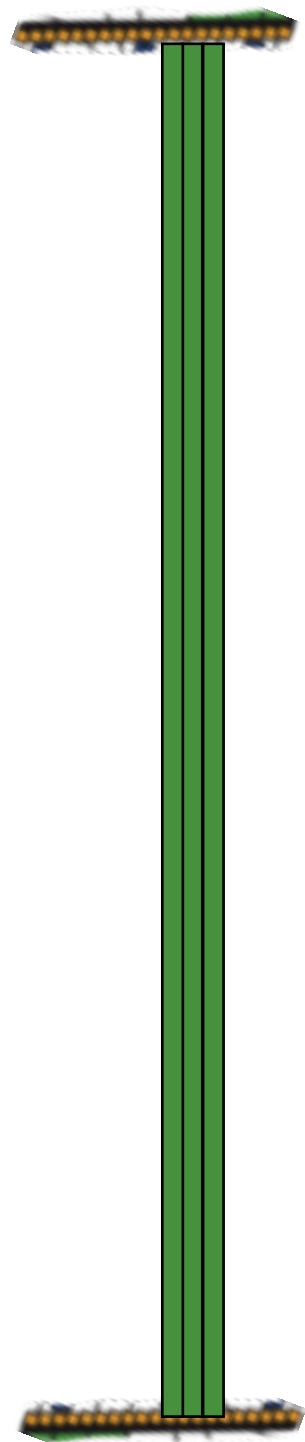


Demonstrating Gkeyll's EMGK capabilities: dancing field lines



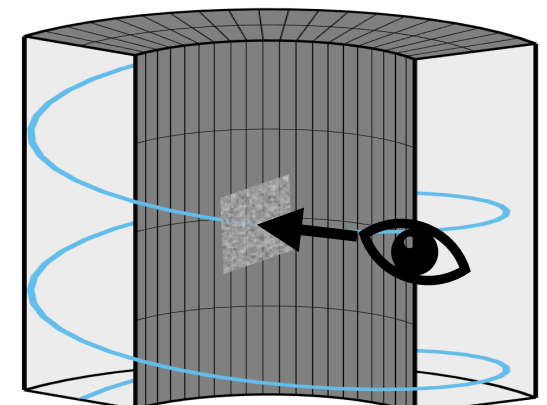
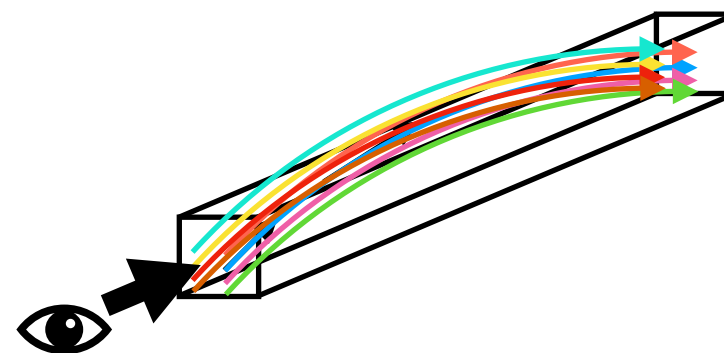
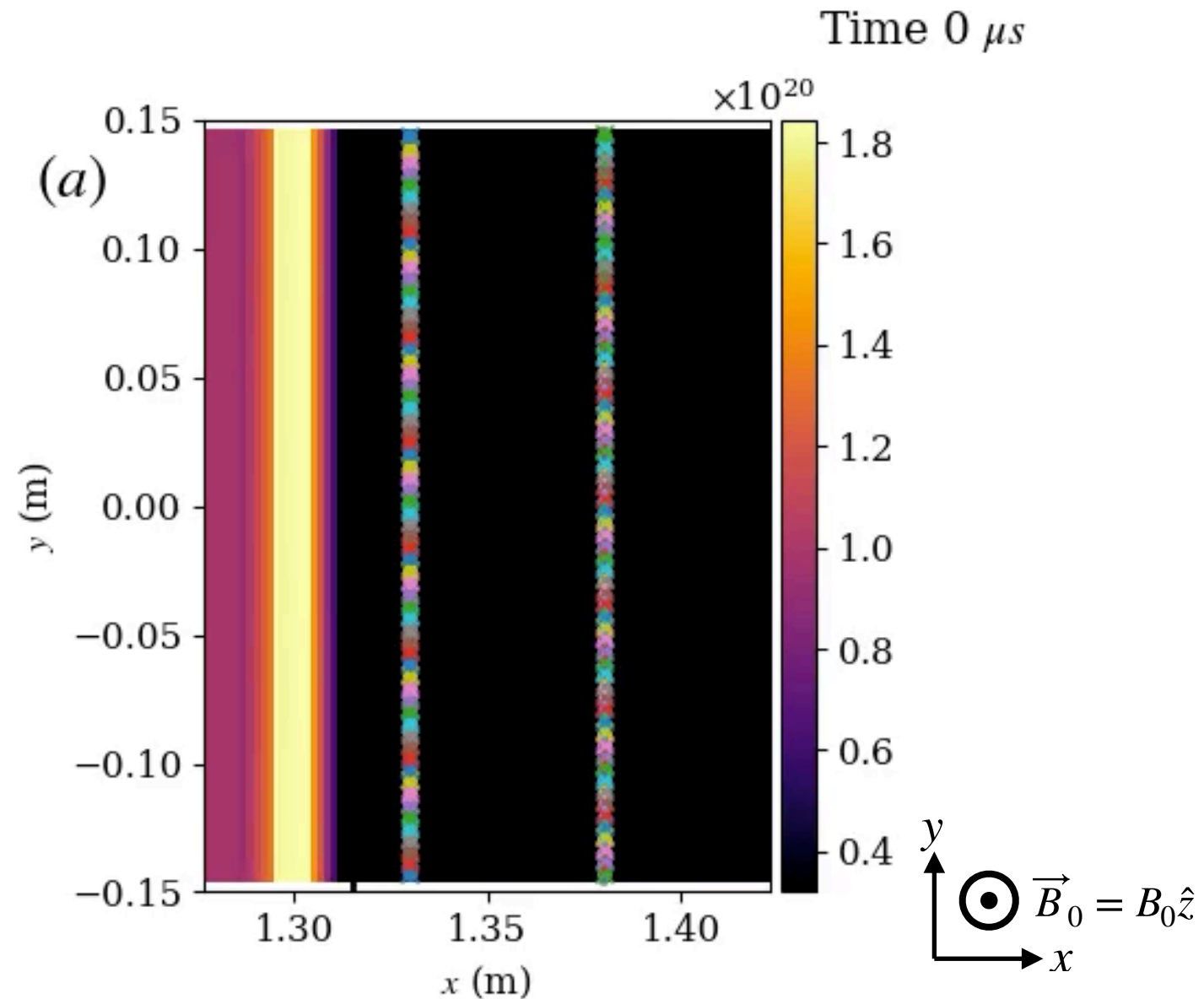
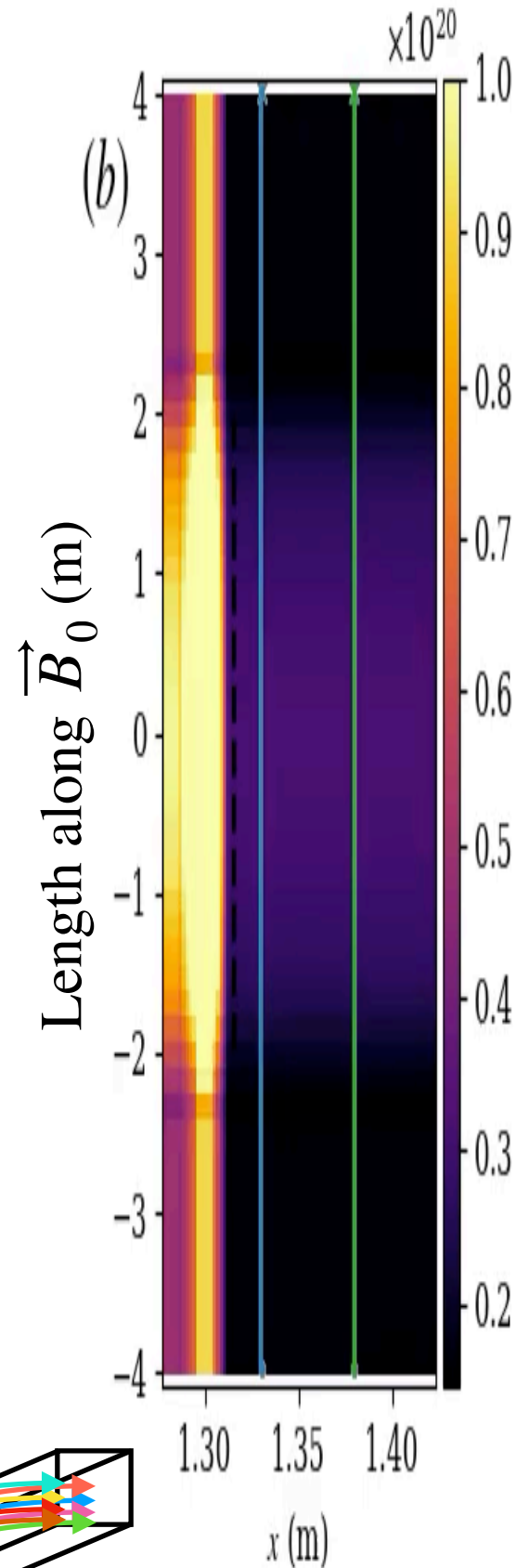
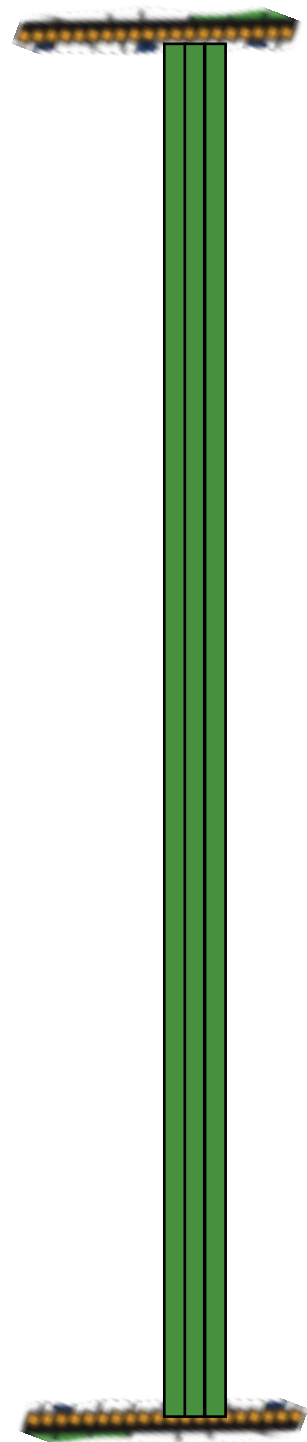
Demonstrating Gkeyll's EMGK capabilities:

dancing field lines



Demonstrating Gkeyll's EMGK capabilities:

dancing field lines



Blobs ($\beta \sim 1\%$) bend/stretch magnetic field lines

First EMGK simulations of SOL ✓

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- We've used simple helical geometry and parameters from NSTX H-mode SOL, but with $10\times n_0$ to stress-test EM effects (could happen locally in ELM?)
 - Results in magnetic fluctuations $\delta B_{\perp}/B_0 \sim 1\%$
 - **Gkeyll** can handle this strong magnetic turbulence robustly ✓
 - Mandell et al, JPP 2020; Hakim et al, PoP 2020

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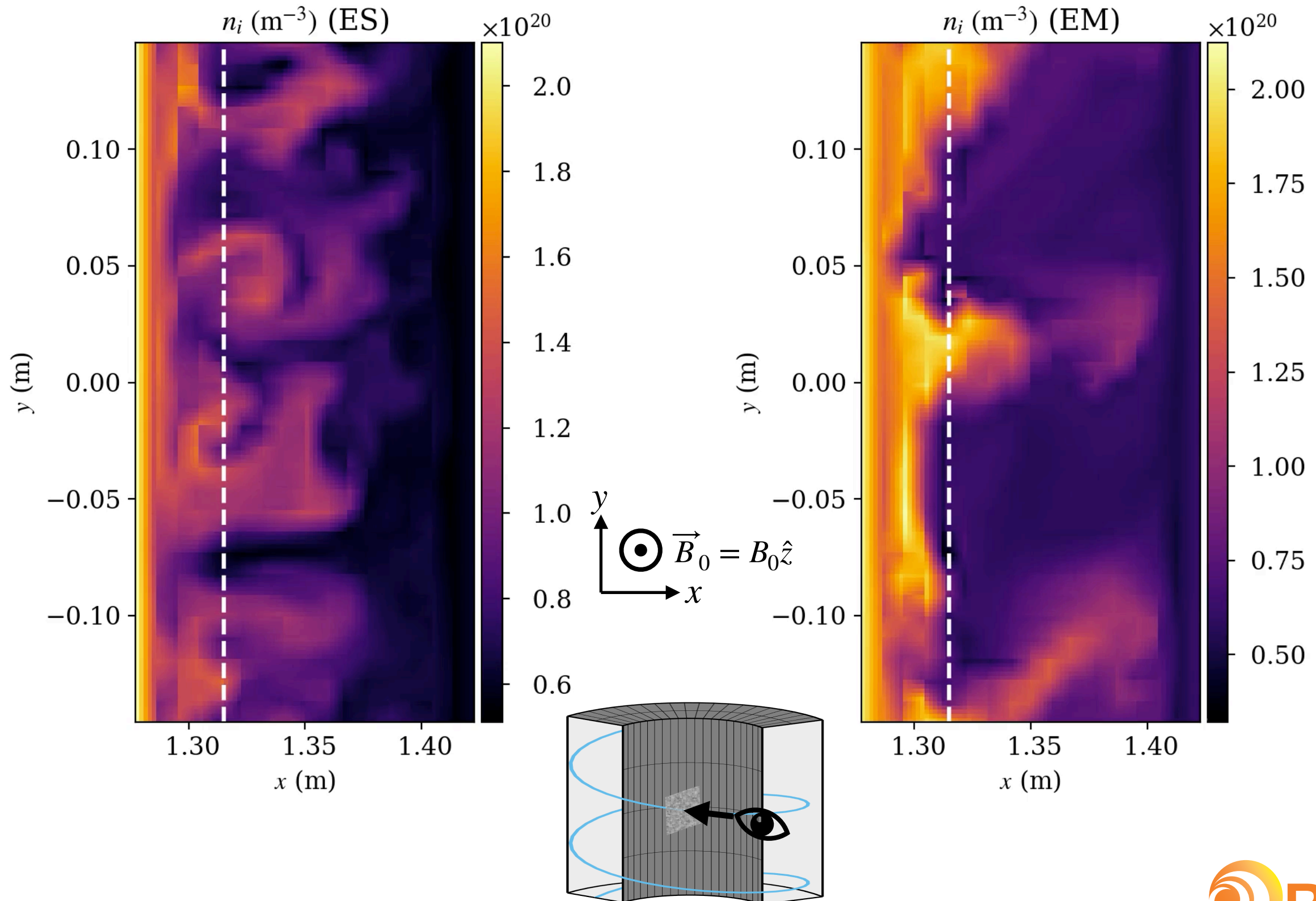
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Do EM fluctuations affect SOL turbulence dynamics?

- Now going to do side-by-side comparison of electrostatic and electromagnetic cases. Things to look for:
 - Changes in the blob structures, dynamics, frequency, etc

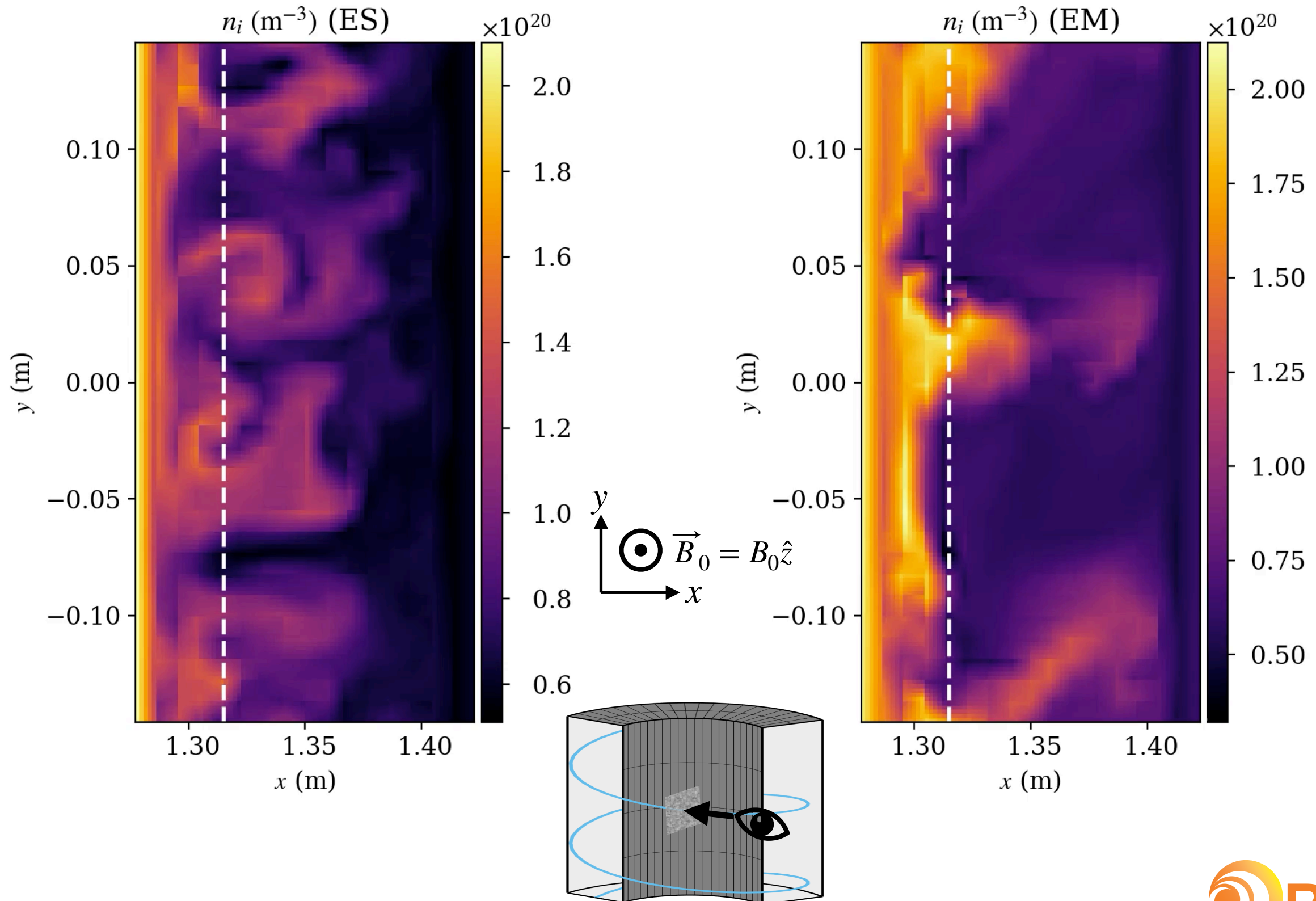
Electrostatic/electromagnetic comparison: midplane ion density

Time 500 μs

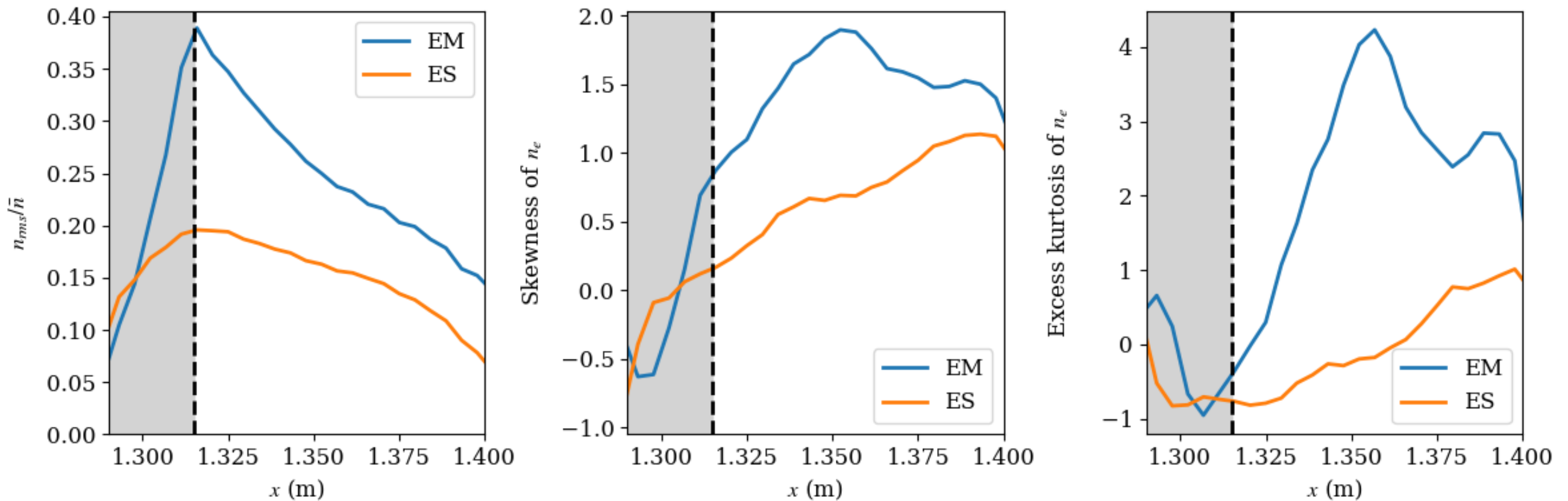


Electrostatic/electromagnetic comparison: midplane ion density

Time 500 μs



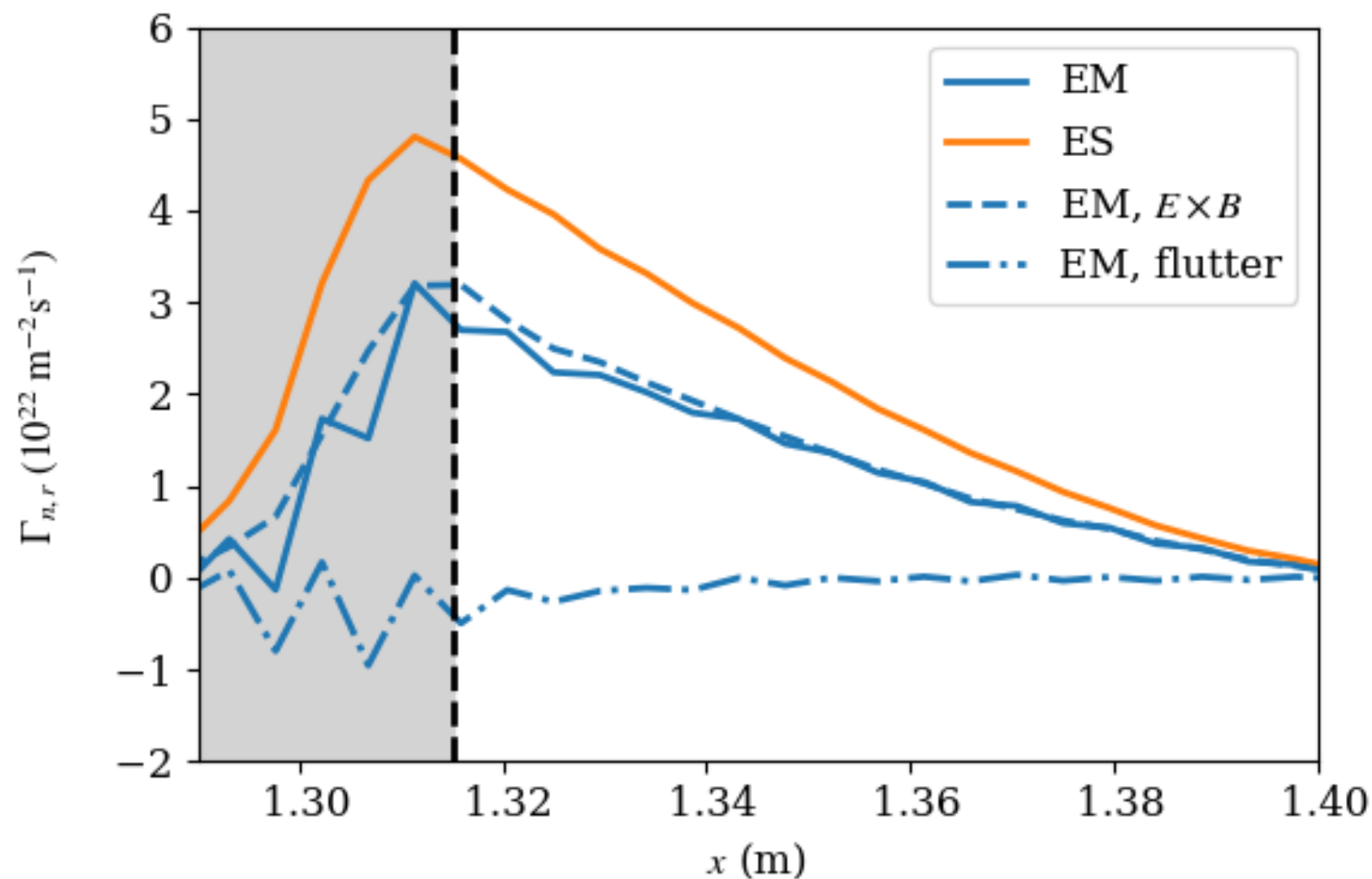
Electrostatic/electromagnetic comparison: density fluctuations statistics



EM has larger, more intermittent density fluctuations

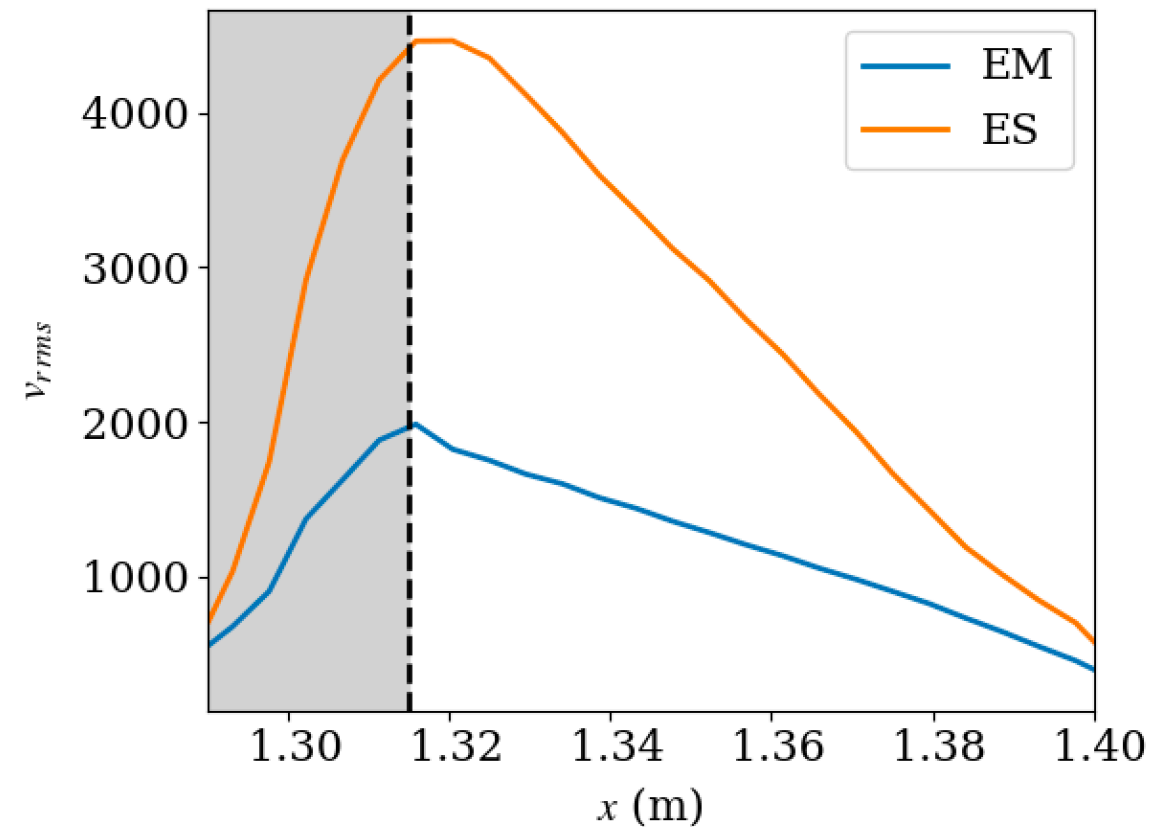
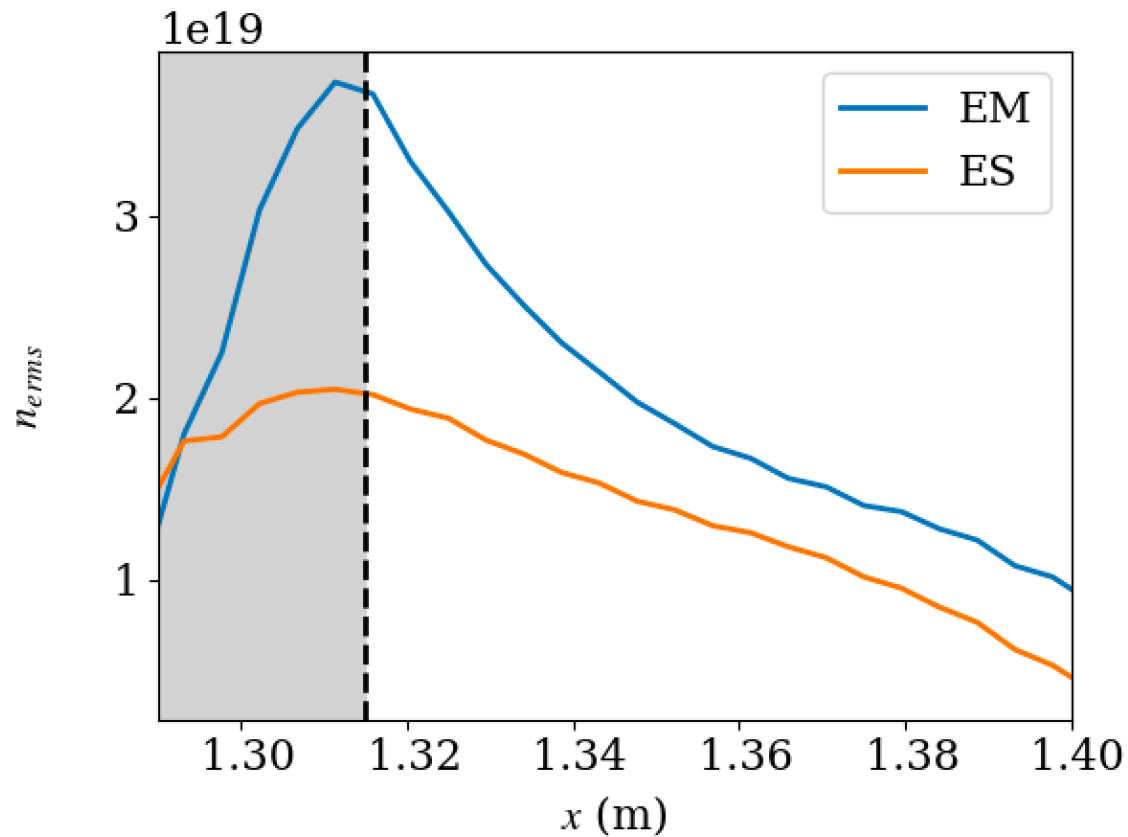
Electrostatic/electromagnetic comparison: radial particle flux (near midplane)

- Might expect larger density fluctuations means more transport, but...



Radial particle transport reduced in EM case by $\sim 40\%$

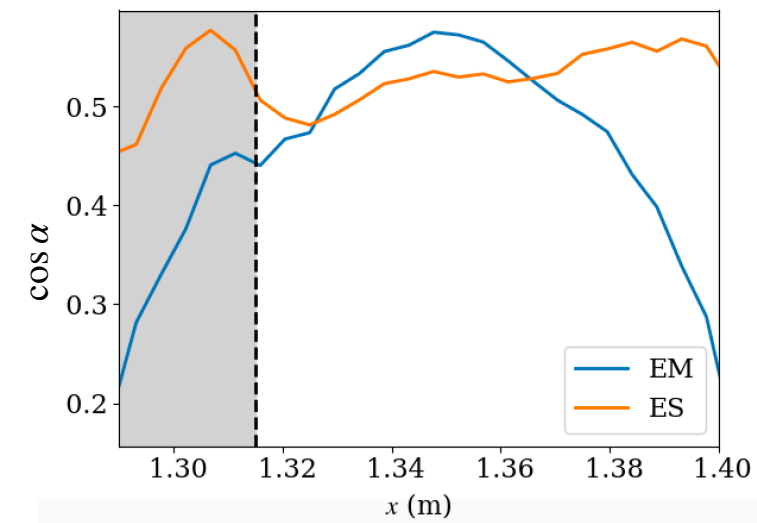
Electrostatic/electromagnetic comparison: radial particle flux (near midplane)



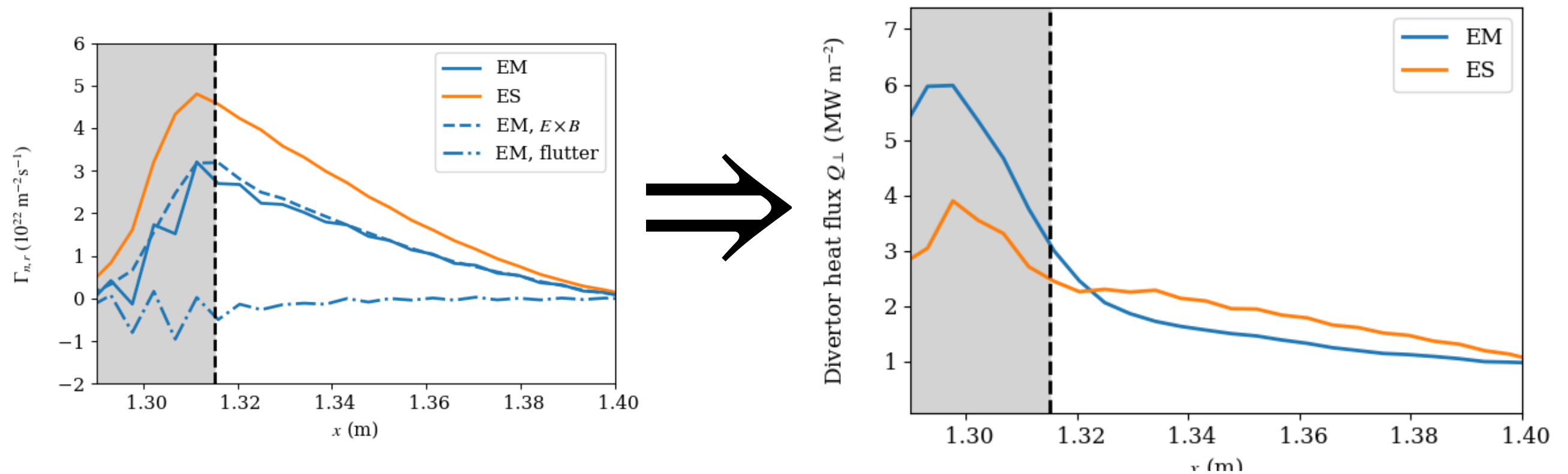
The radial $E \times B$ particle flux is defined as

$$\Gamma_r = \langle \tilde{n}_e \tilde{v}_r \rangle = n_{e,rms} v_{r,rms} \cos \alpha$$

In this case, using n_{rms} as a surrogate diagnostic for transport is not sufficient!



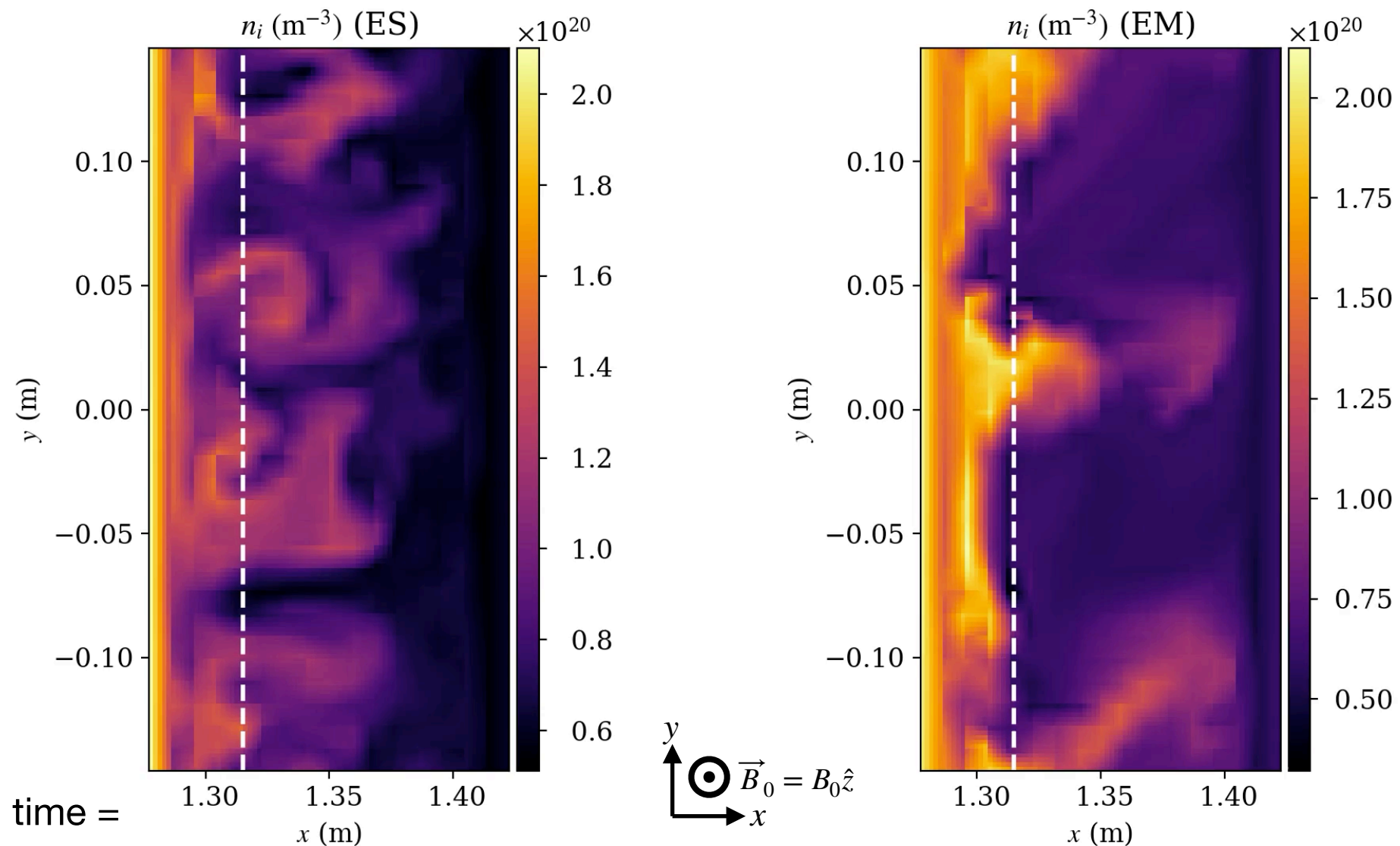
Electrostatic/electromagnetic comparison: divertor heat flux profile



- Because EM case has less radial transport, heat flux to divertor is more peaked in EM case
- ES case over-predicts transport, over-predicts heat flux width

Modest simulation cost (even for EM!)

Time 500 μs



Total simulation time =
1 ms \sim 20 ion transit times

(Nx, Ny, Nz, Nv, Nm)
 \sim (32, 64, 20, 20, 10)

Electrostatic

Electromagnetic

Total wall-clock time
(128 cores)

2.7 days

3.4 days

**only ~25%
more**

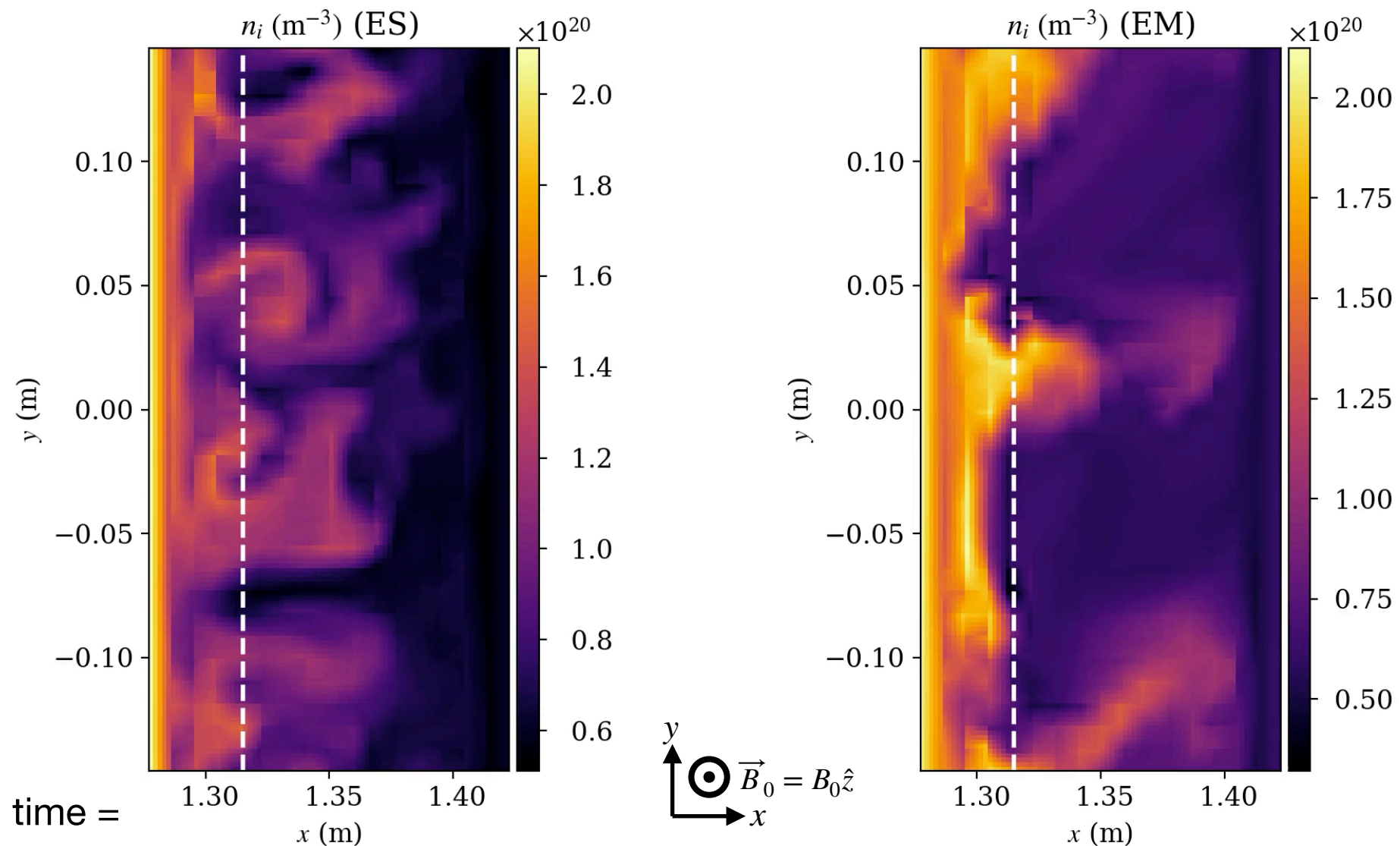
Time/timestep
(wall-clock)

0.41 s

0.68 s

Modest simulation cost (even for EM!)

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Time/timestep
(wall-clock)

0.41 s

0.68 s

Towards more realistic SOL geometry

- We know magnetic shear can be important in SOL, especially near X point
- All Gkeyll results to date (NSTX and Helimak) have used simplified helical geometry
 - Neglected most geometrical factors, no magnetic shear
- Even in SMT configuration (e.g. Helimak) with const vertical field, there should be some magnetic shear because toroidal field $\sim 1/R$

$$\mathbf{B} = \frac{B_0 R_0}{R} \hat{\boldsymbol{\phi}} + B_v \hat{\mathbf{Z}}, \quad q(R) = \frac{H B_\phi}{2\pi R B_v} = \frac{B_0 R_0 H}{2\pi R^2 B_v}, \quad \hat{s} = \frac{R}{q} \frac{dq}{dR} = -2$$

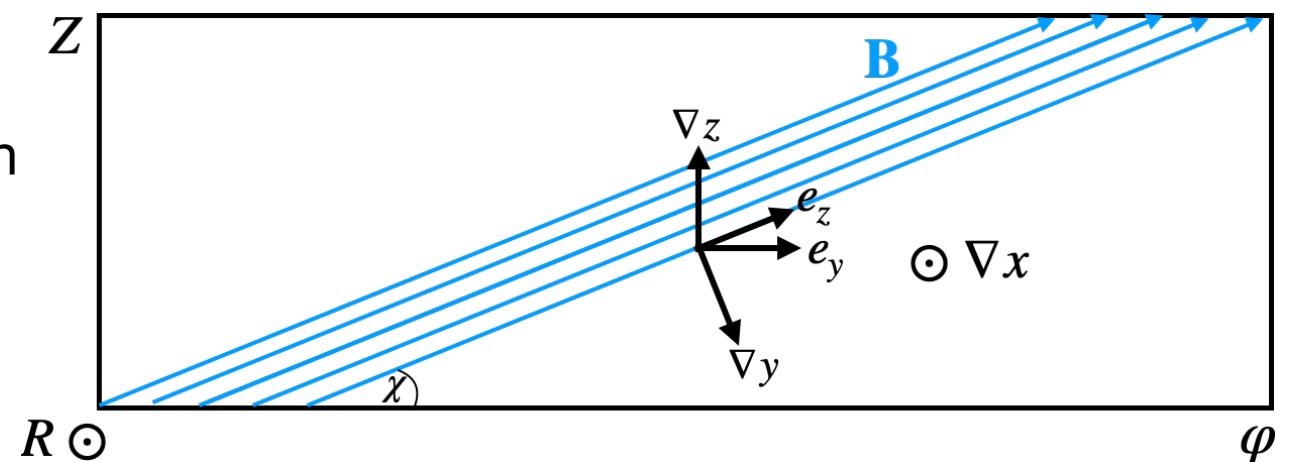
- Keeping helical configuration, can adjust shear by making $B_v = B_v(R) = B_{v0}(R/x_0)^n$, so that

$$\hat{s} = -2 - \frac{R}{B_v} \frac{dB_v}{dR} = -2 - n$$

- Take field-aligned helical coordinate system with

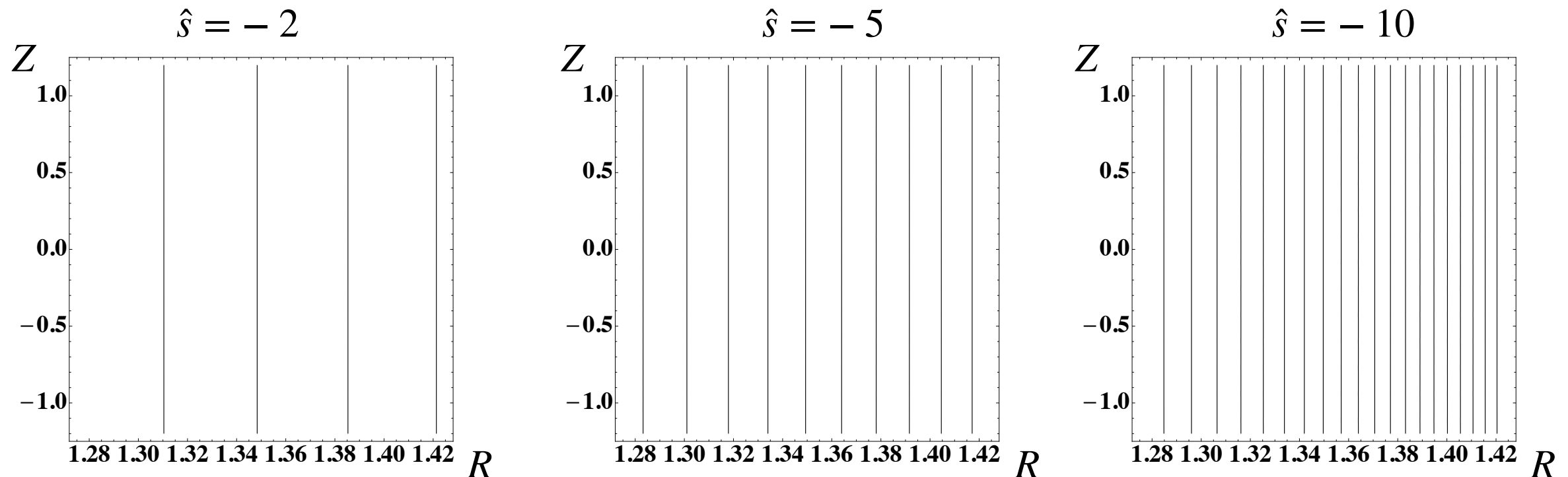
$$x = R, \quad z = Z, \quad y = x_0 \left(\varphi - \frac{2\pi q Z}{H} \right),$$

$$\mathbf{B} = \frac{R B_v}{x_0} \nabla x \times \nabla y$$

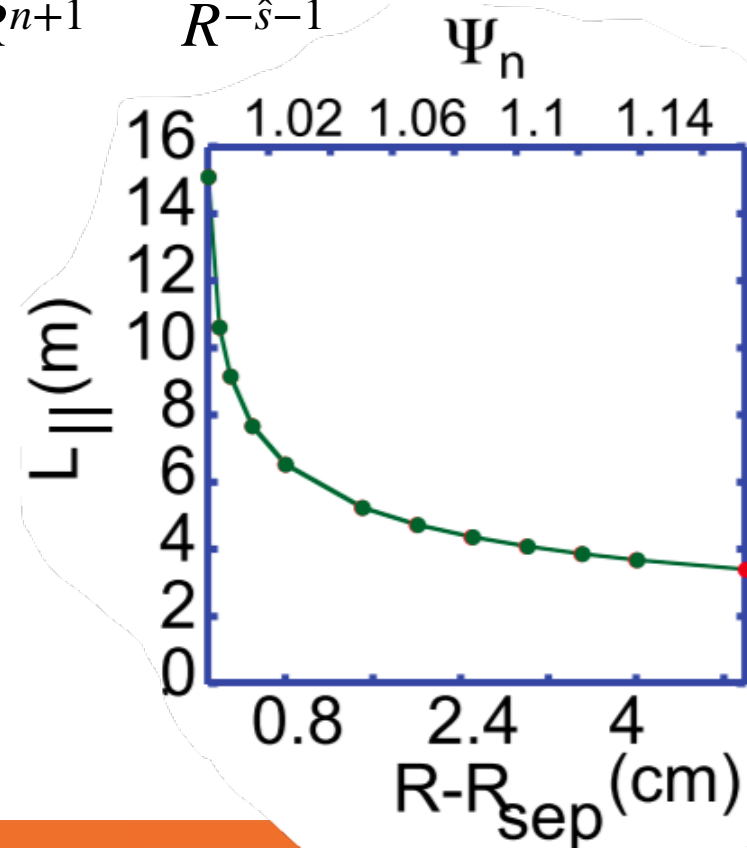
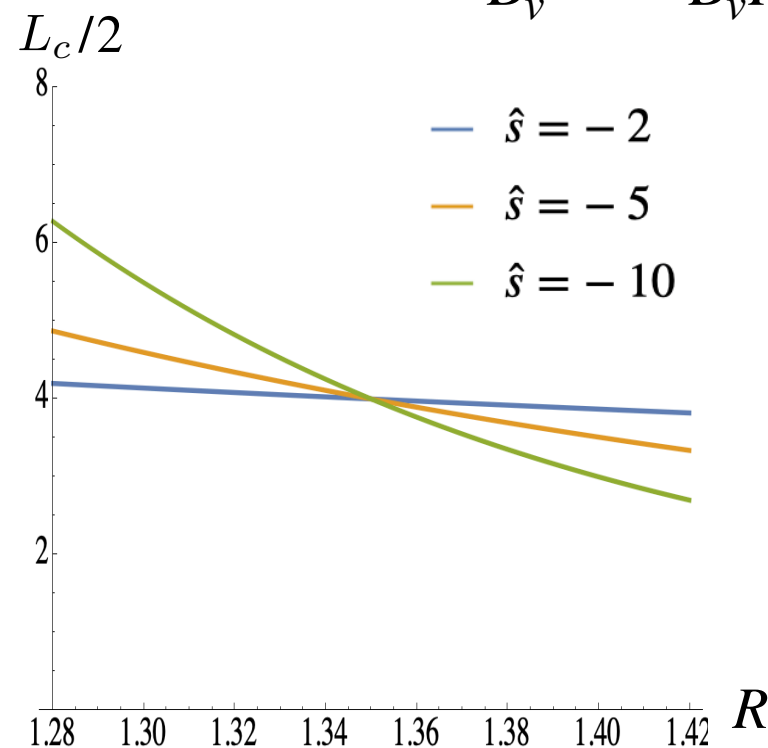


Helical geometry with magnetic shear

Vertical (~poloidal) flux: $\Psi(R, Z) = R^2 B_v / 2 \sim R^{-\hat{s}}$

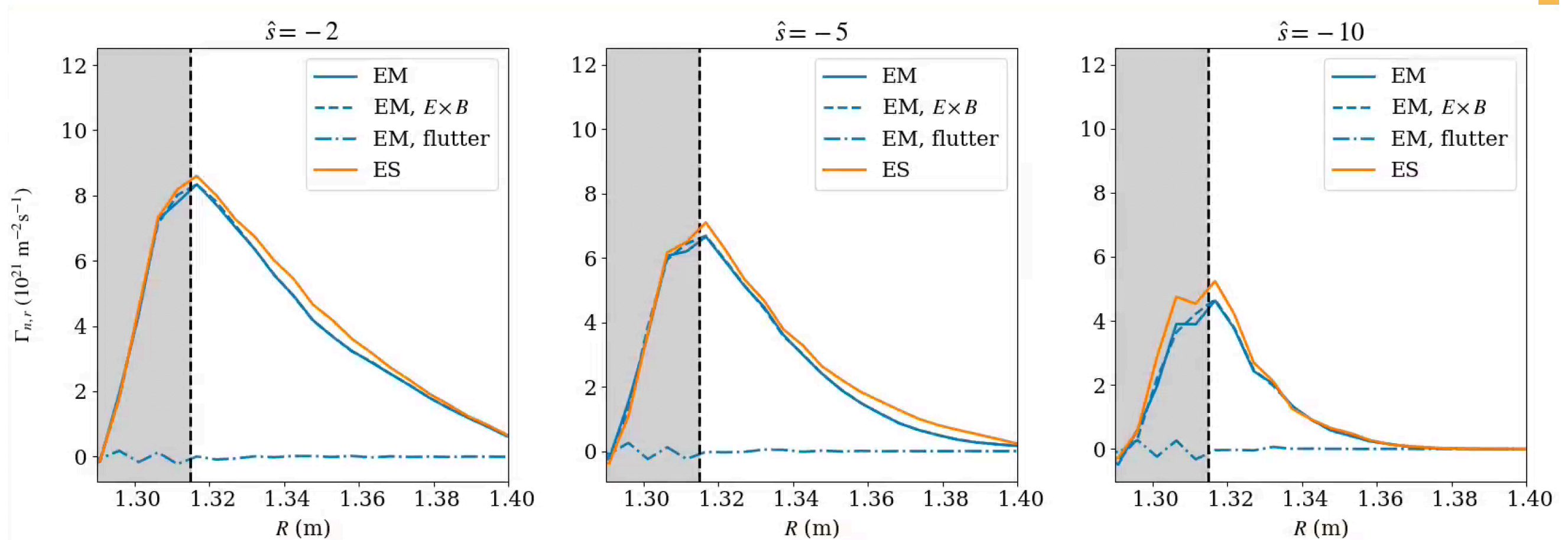


Connection length: $L_c = \frac{HB}{B_v} \approx \frac{HB_0 R_0}{B_v R} \sim \frac{1}{R^{n+1}} \sim \frac{1}{R^{-\hat{s}-1}}$



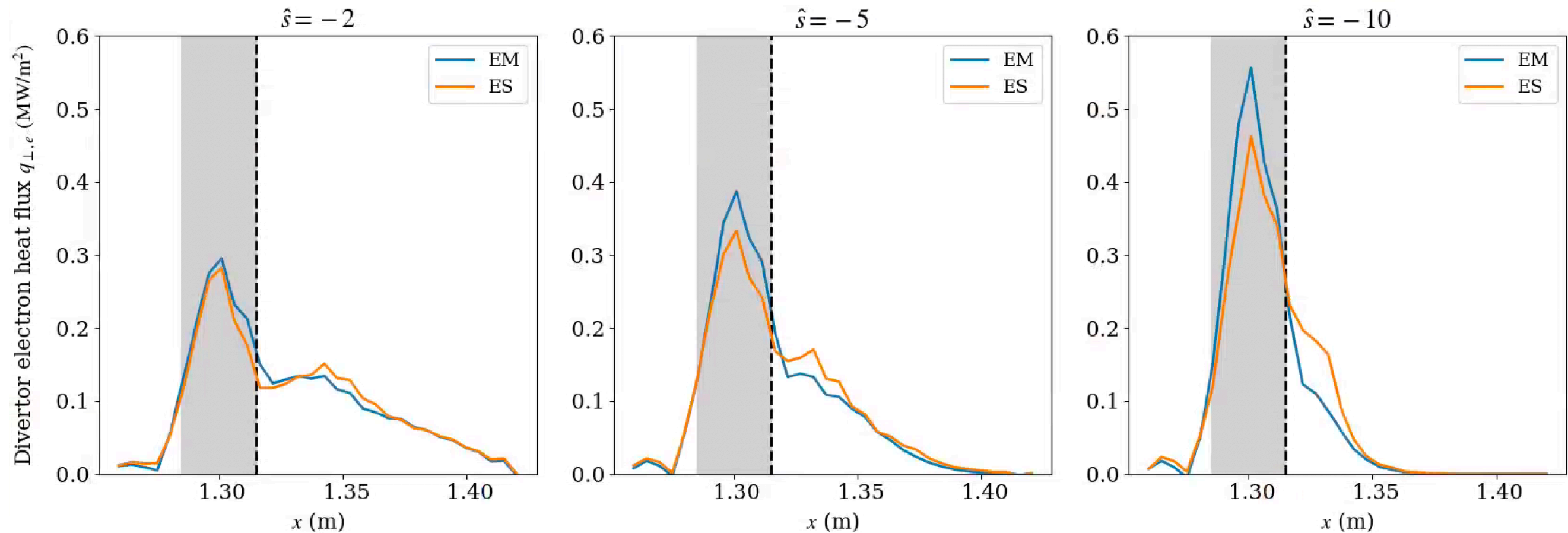
NSTX SOL
connection length
Boedo et al, PoP 2014

\hat{s} scan: radial particle flux (near midplane)



- Transport decreases as $|\hat{s}|$ increases
- EM cases again have less transport
- This is at \sim experimental $\beta \sim 0.1 \%$ (no more 10x)

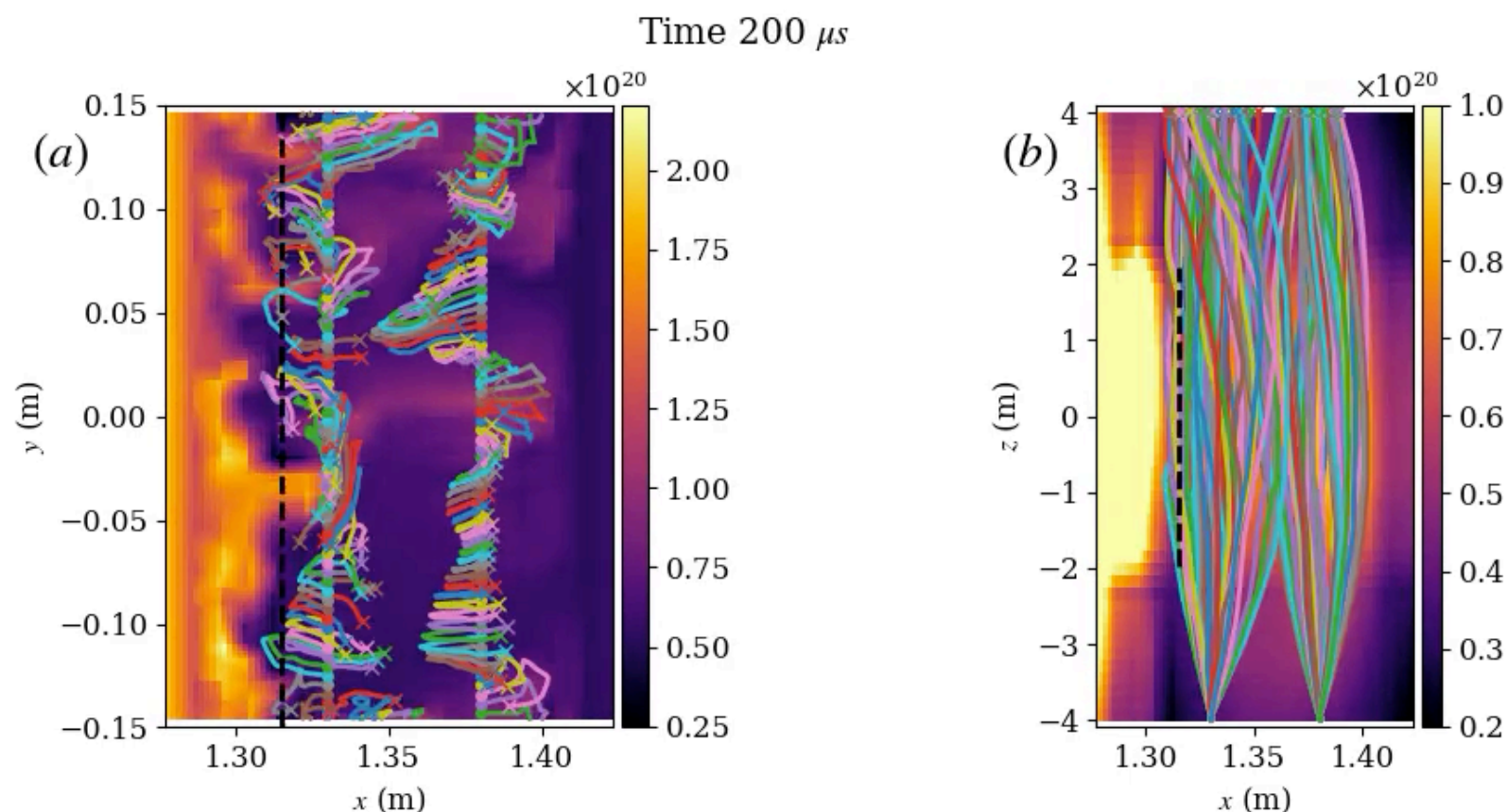
\hat{s} scan: divertor heat flux profiles



- Heat flux profile gets narrower as $|\hat{s}|$ increases
- EM effects more important as $|\hat{s}|$ increases

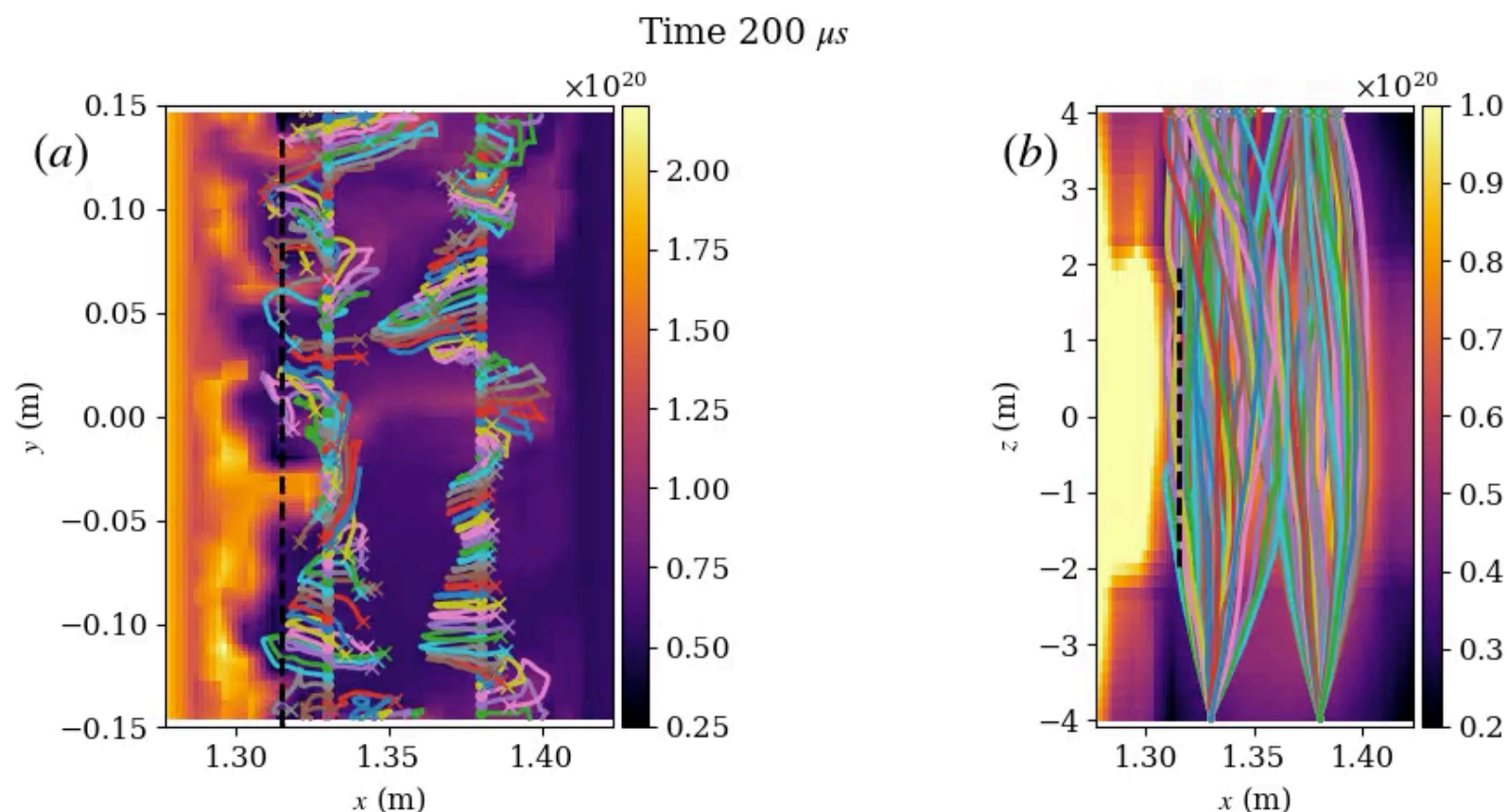
Summary

- **Gkeyll** is being used to study SOL turbulence in tokamaks like NSTX (only handles open field lines right now)
- **Gkeyll** has produced the first nonlinear electromagnetic gyrokinetic simulations in the SOL, can handle strong magnetic turbulence with $\delta B_{\perp}/B_0 \sim 1\%$
- In high β regime, including electromagnetic fluctuations results in larger, more intermittent fluctuations in SOL but less transport
- Moving towards more realistic SOL geometry, including magnetic shear



Summary

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Acknowledgements

Gkeyll team:

Ammar Hakim

Greg Hammett

Tess Bernard

Petr Cagas

Mana Francisquez

Jimmy Juno

Rupak Mukherjee

Liang Wang

Eric Shi

... and others!

<https://github.com/ammarrhakim/gkyl/>

<https://gkeyll.readthedocs.io>



Gkeyll GK references

E. L. Shi, Princeton Ph.D. thesis 2017

A. Hakim et al, arXiv:1908.01814

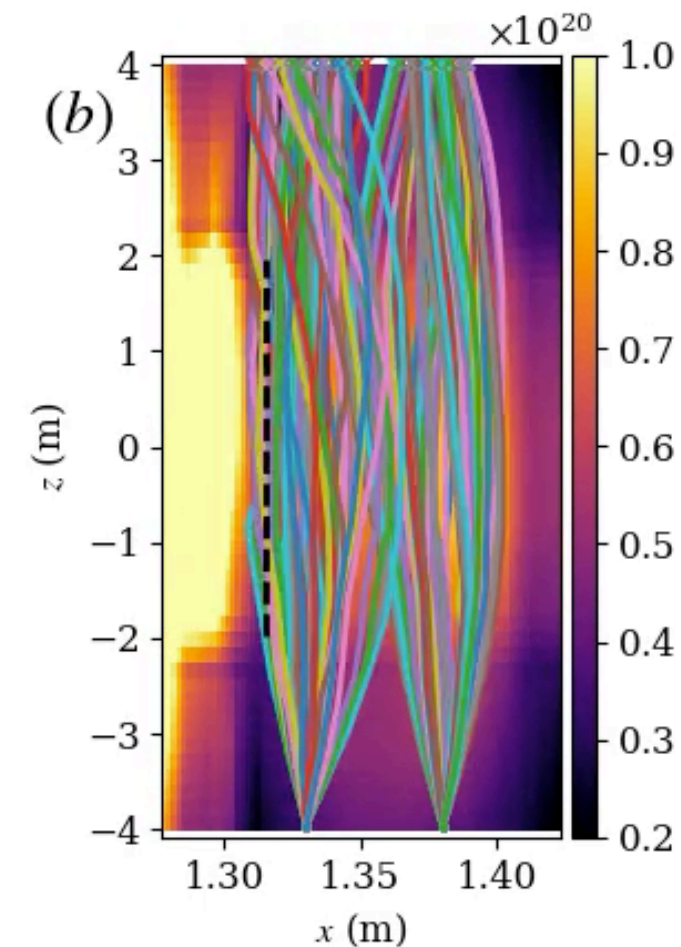
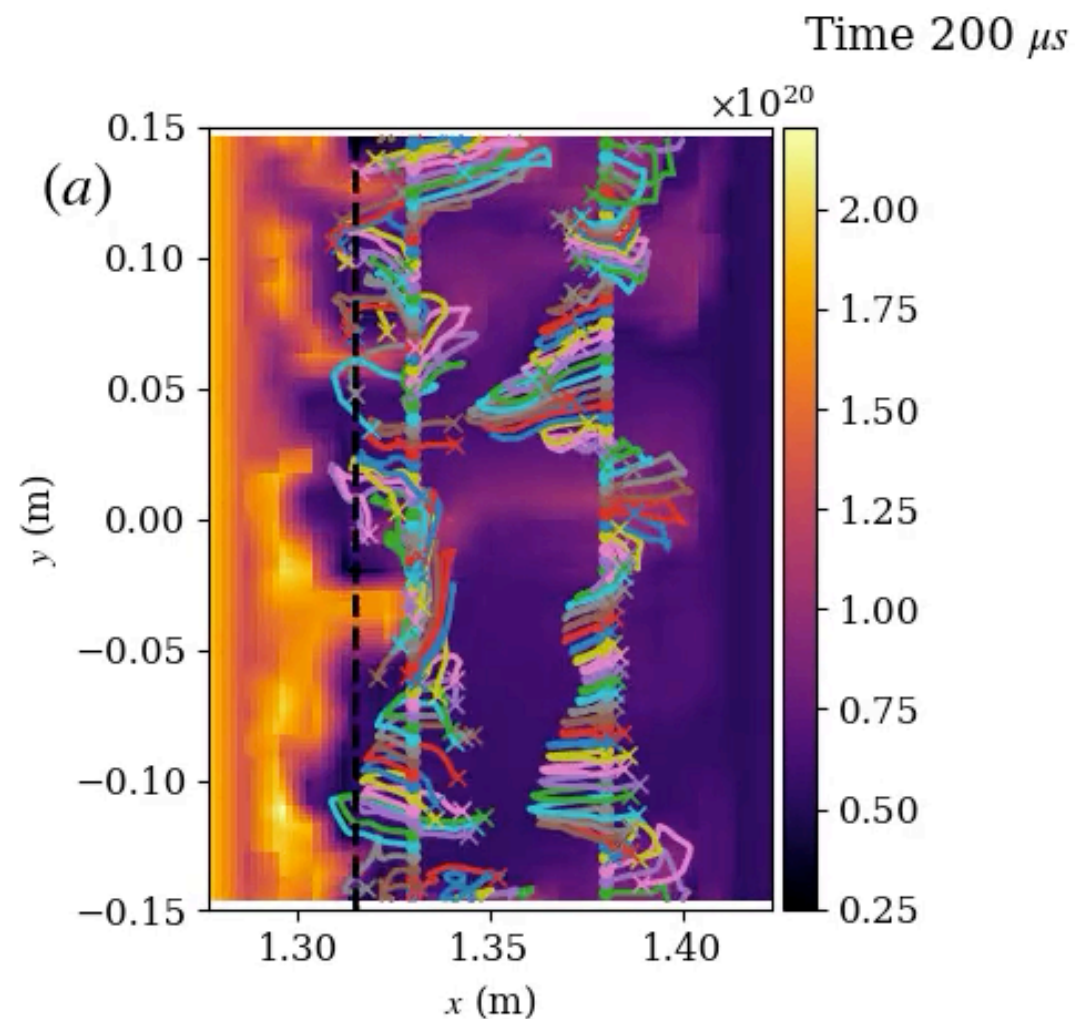
E. L. Shi et al, J. Plasma Phys. 83 (3) (2017)
905830304

**N. R. Mandell et al, J. Plasma Phys. 86 (1)
(2020)**

E. L. Shi et al, Phys. Plasmas 26 (1) (2019)
012307

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042301



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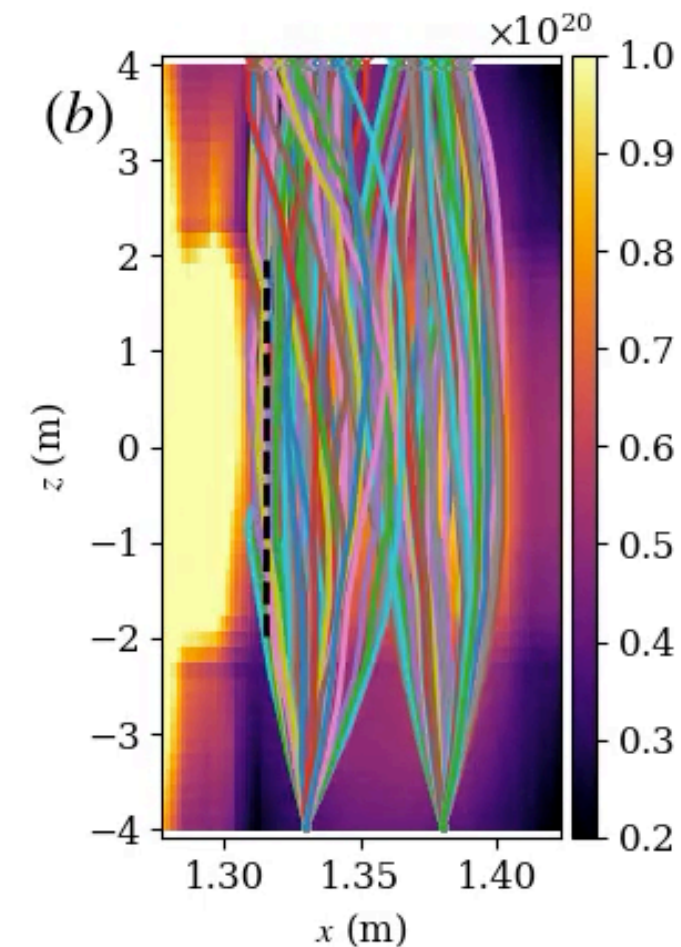
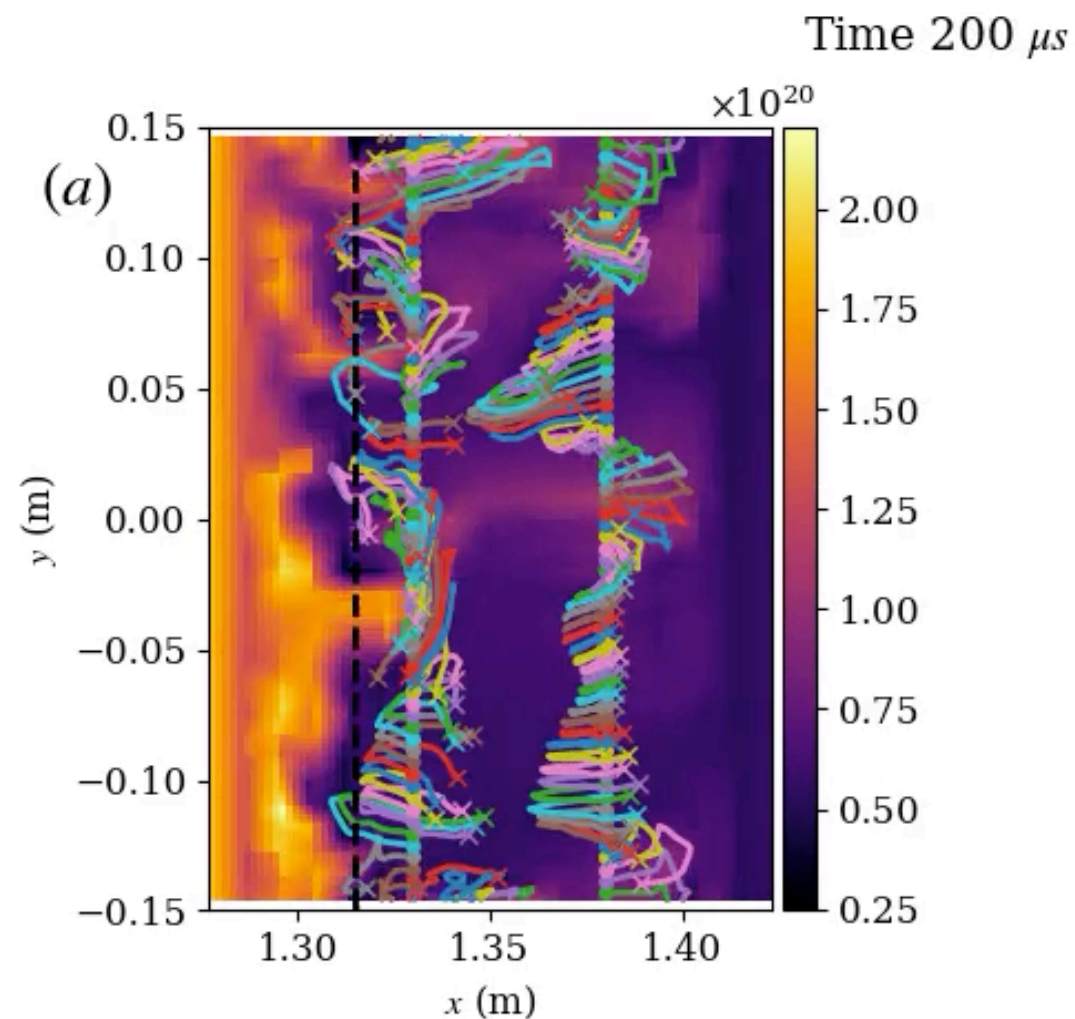
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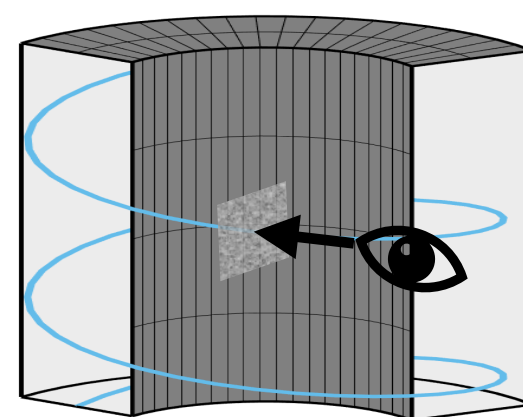
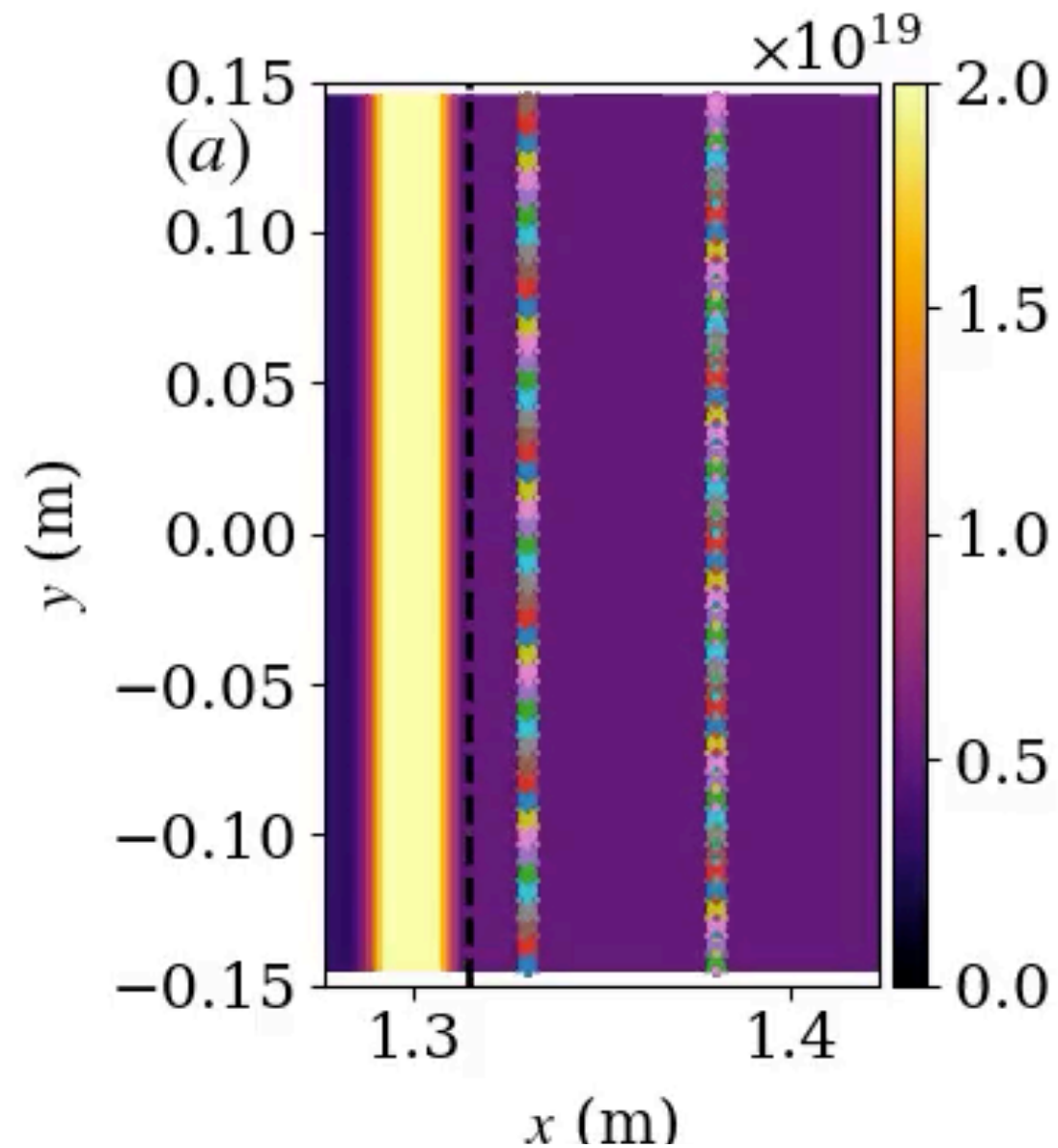
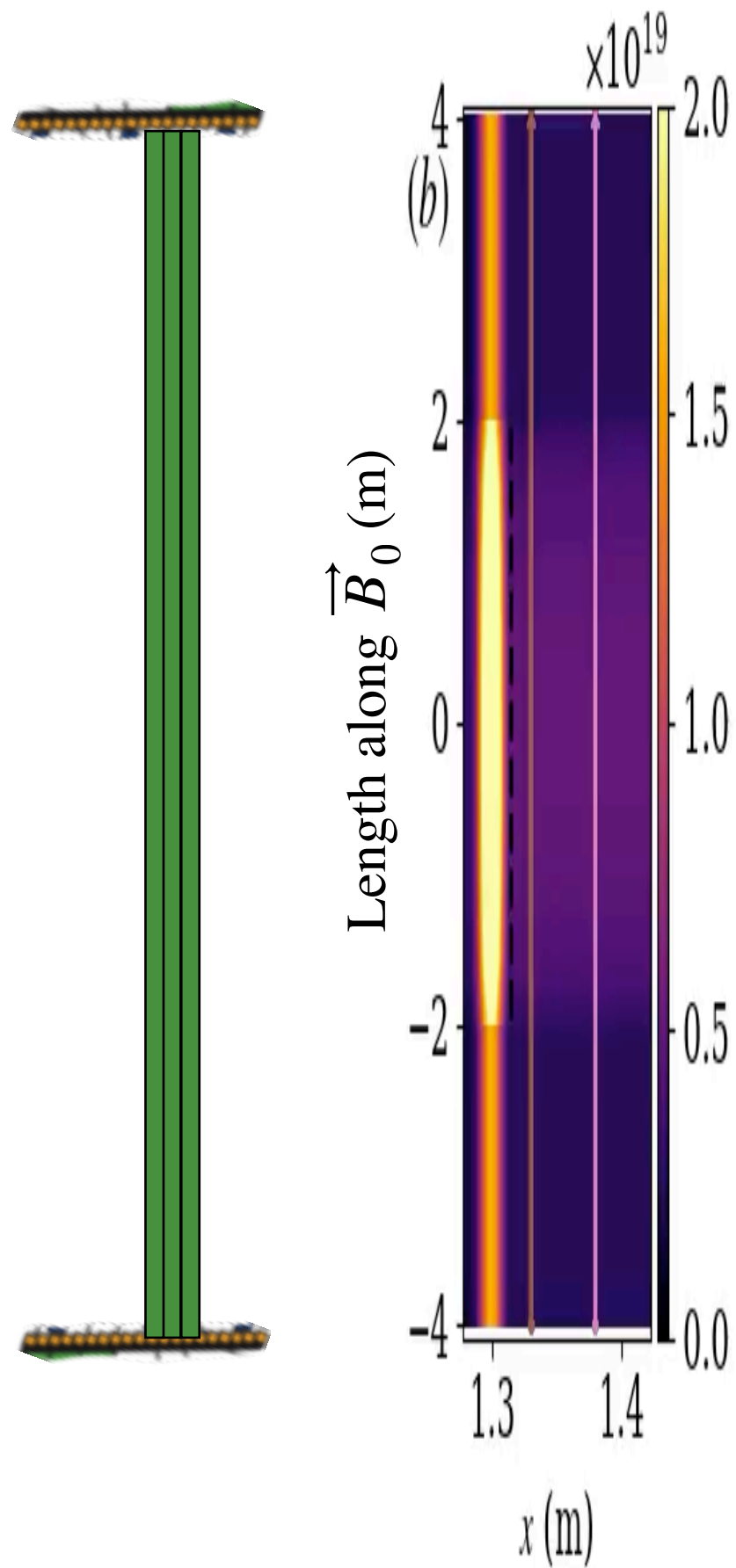
T. Bernard et al, Phys. Plasmas 26 (4) (2019)
042301



Back-up slides

Dancing field lines at \sim experimental β

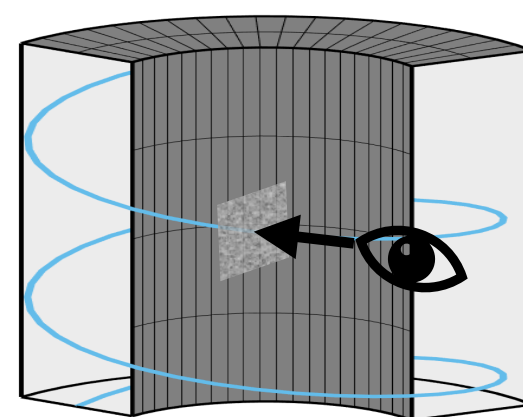
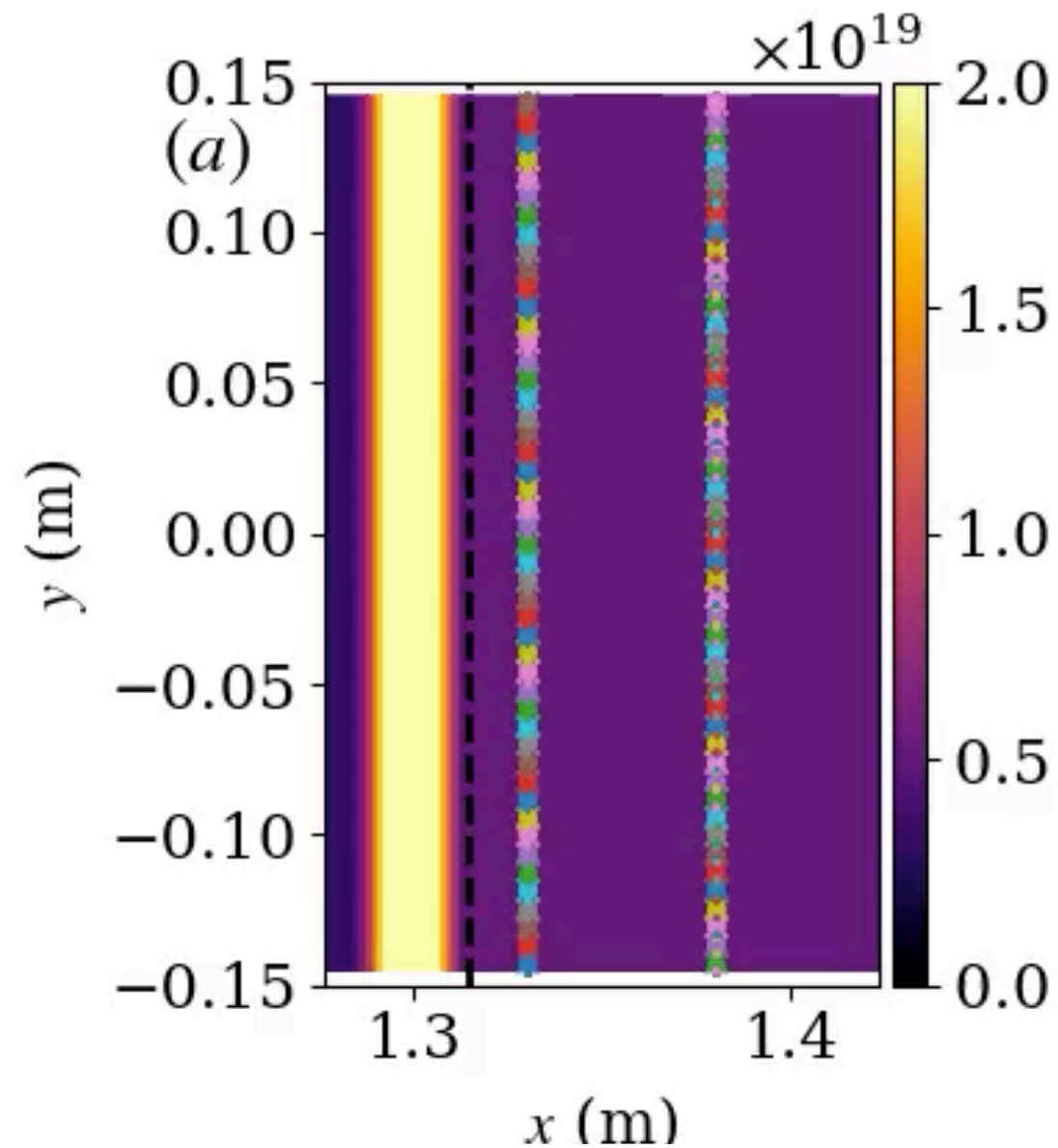
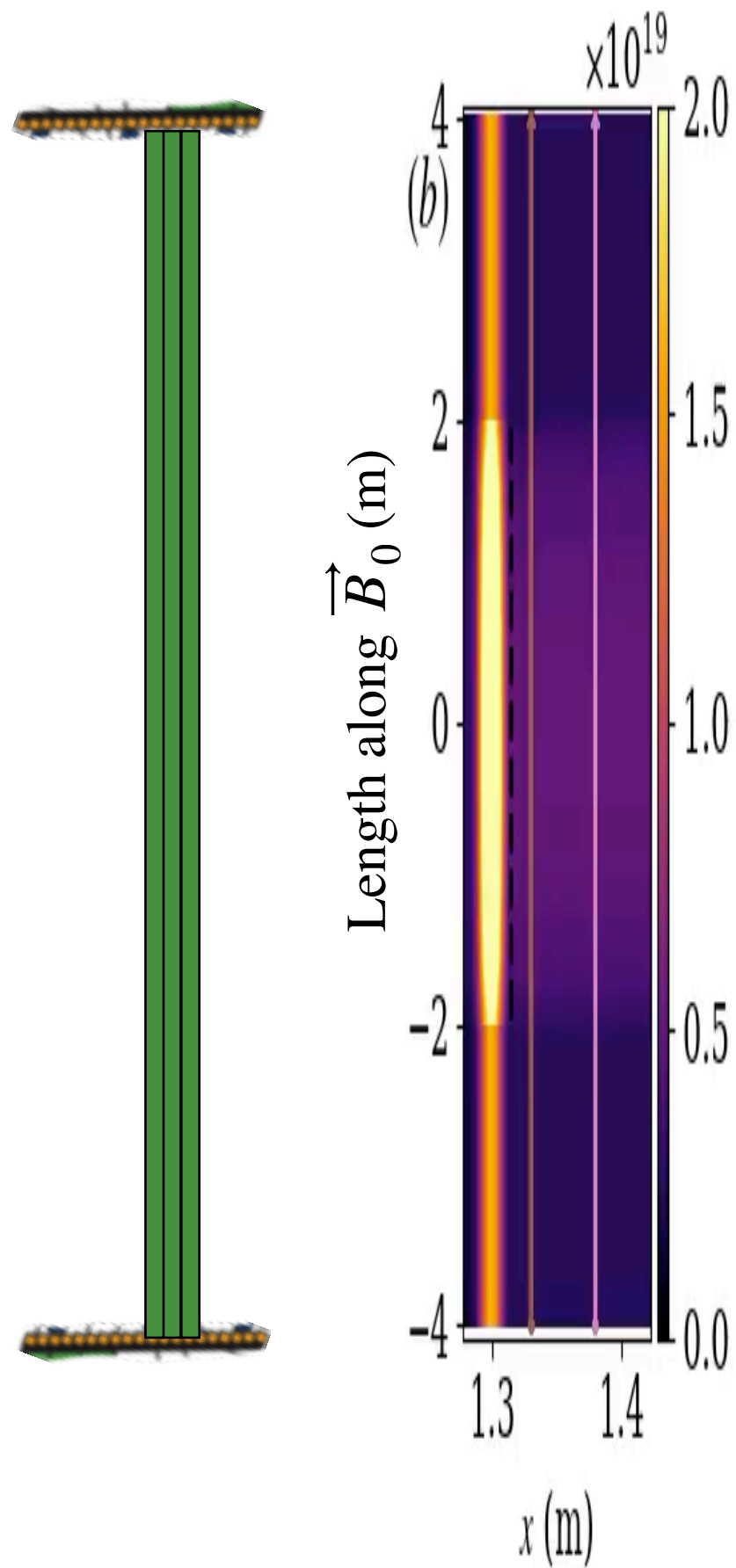
Time 0 μs



$\vec{B}_0 = B_0 \hat{z}$

Dancing field lines at \sim experimental β

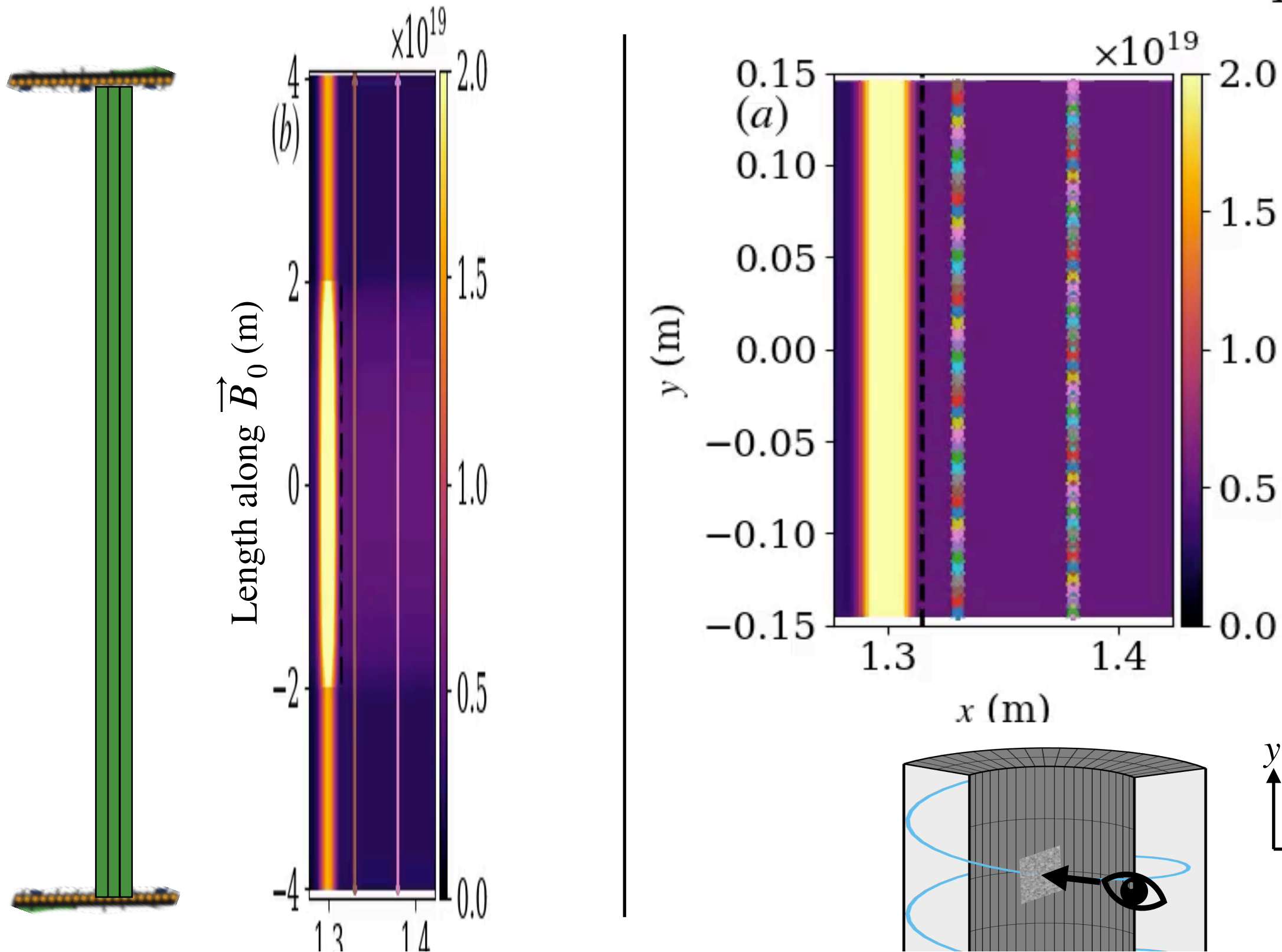
Time 0 μs



$$\vec{B}_0 = B_0 \hat{z}$$

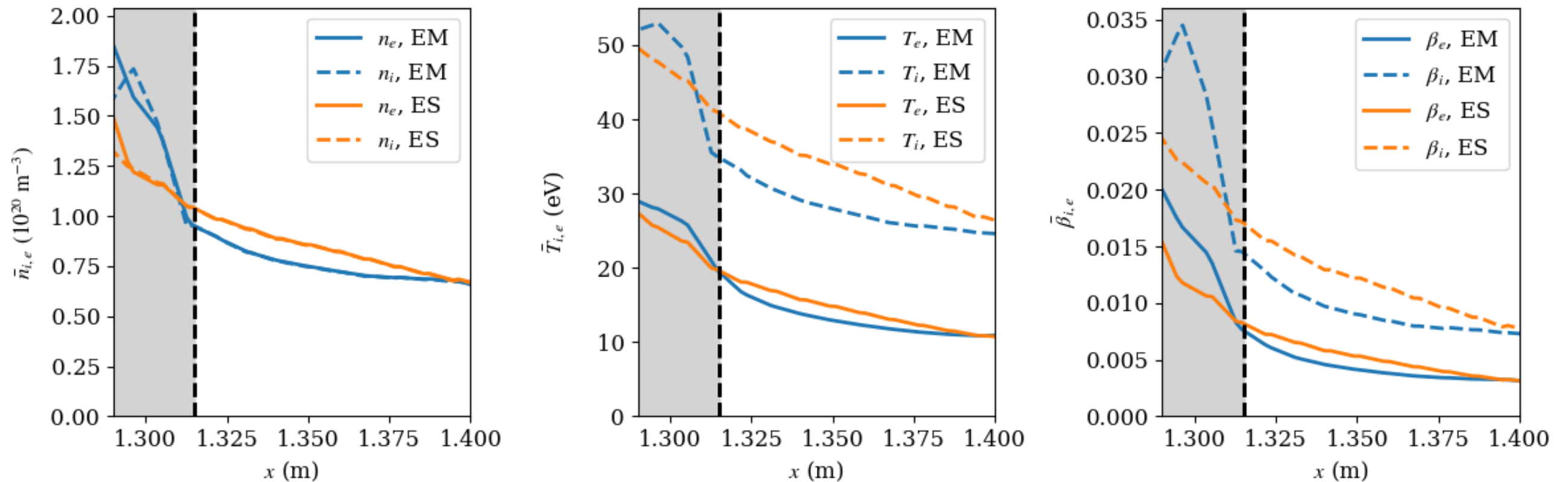
Dancing field lines at \sim experimental β

Time 0 μs



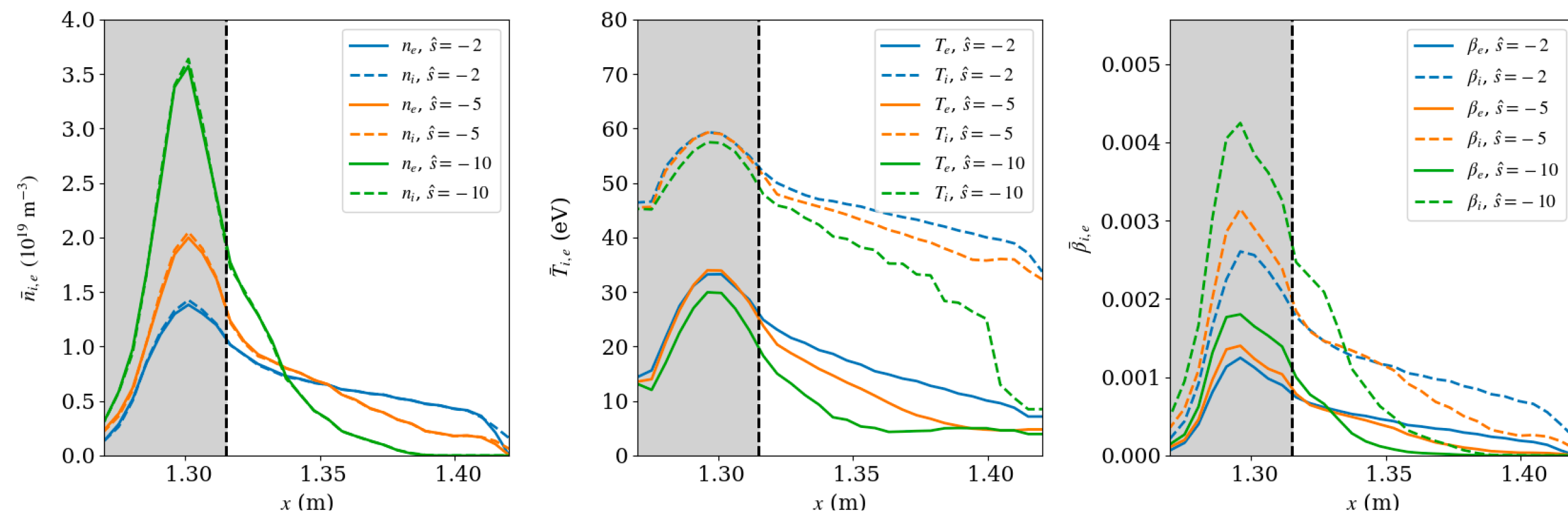
But increasing connection length, adding magnetic shear can also increase magnetic fluctuations and affect transport

Electrostatic/electromagnetic comparison: midplane profiles



- $\beta_e \sim 0.5 \%$, $\beta_i \sim 1 \%$ in SOL region ($\sim 10 \times$ NSTX SOL)
- Profiles are steeper in source region, shallower in SOL region in EM case

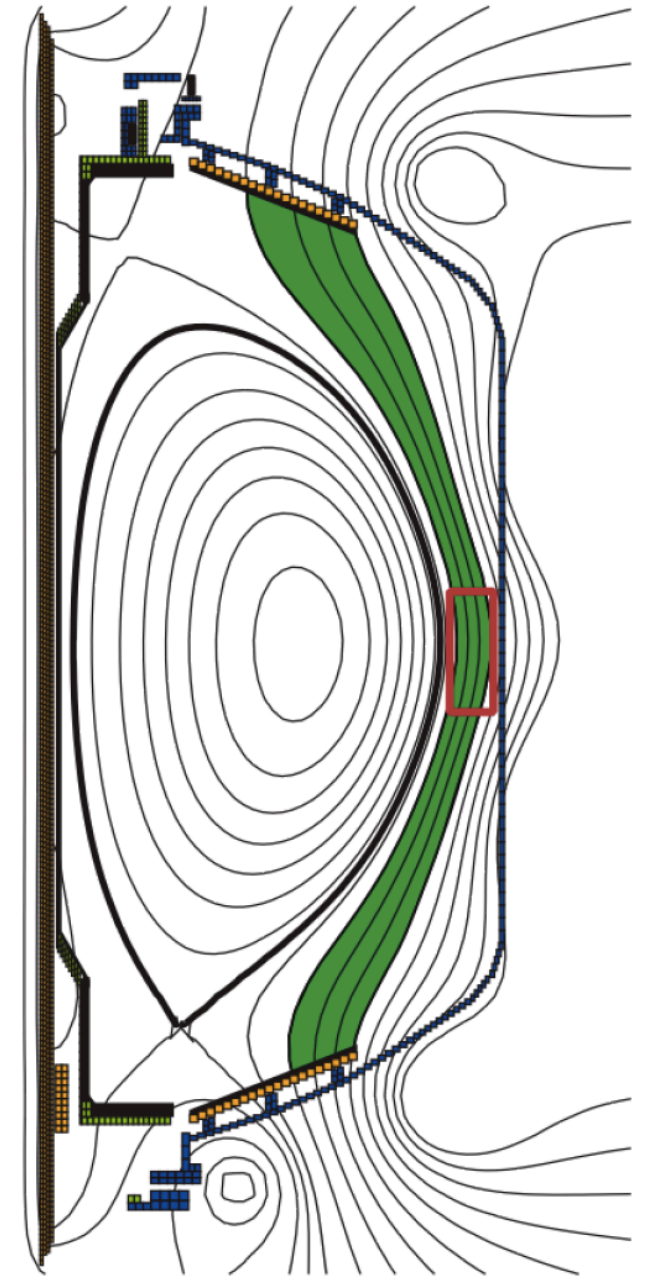
\hat{s} scan: midplane profiles (EM)



- $\beta \sim 0.1 \%$ (\sim NSTX SOL)
- Profiles drop off more quickly as $|\hat{s}|$ increases
- Previous simulations with simplified geometry similar to $\hat{s} = -2$ case

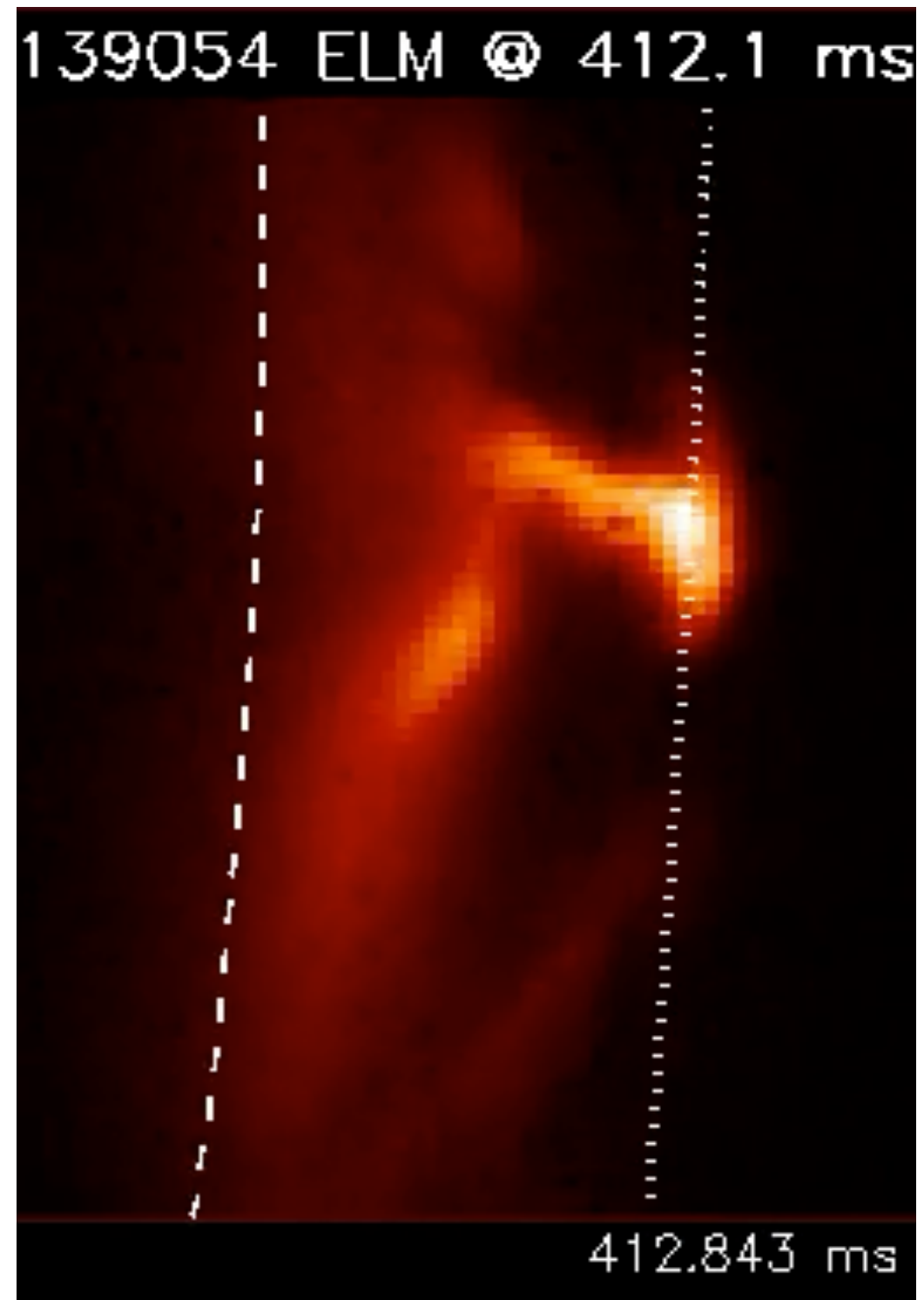
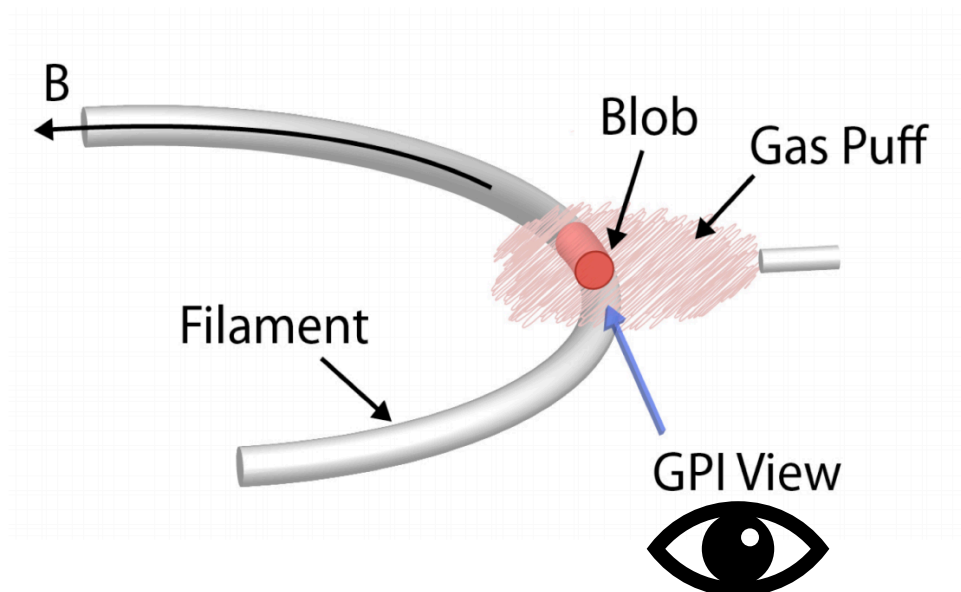
Current/Future Work

- Generalizing the magnetic geometry to include magnetic shear, non-constant curvature, closed field line regions, X-point
 - Non-orthogonal field-aligned coordinate system with magnetic shear now implemented
 - X-point is a singularity in these coordinates, challenging!
- More studies of EM effects on blobs/ELMs
 - Comparisons with magnetic fluctuation measurements in experiments?



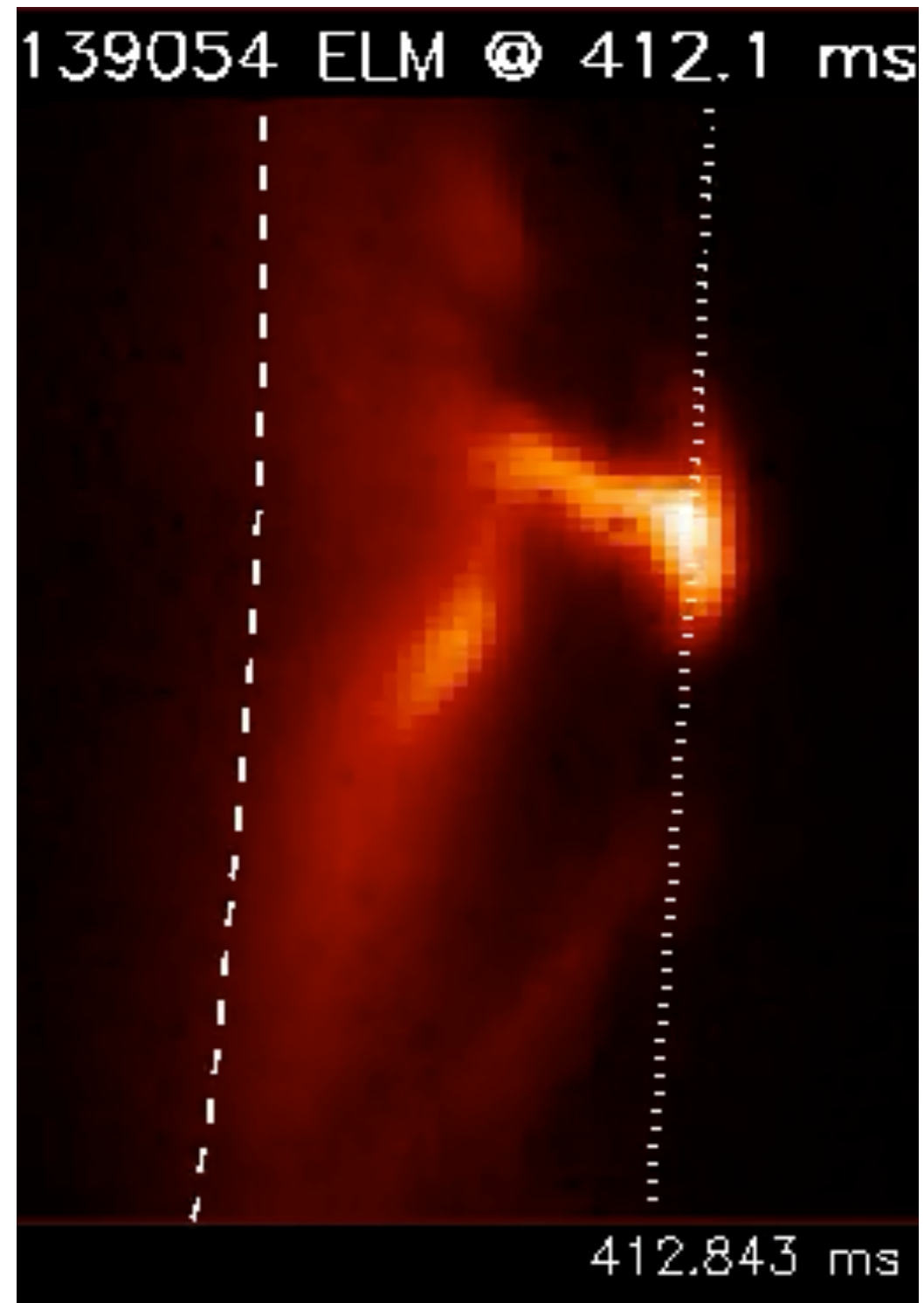
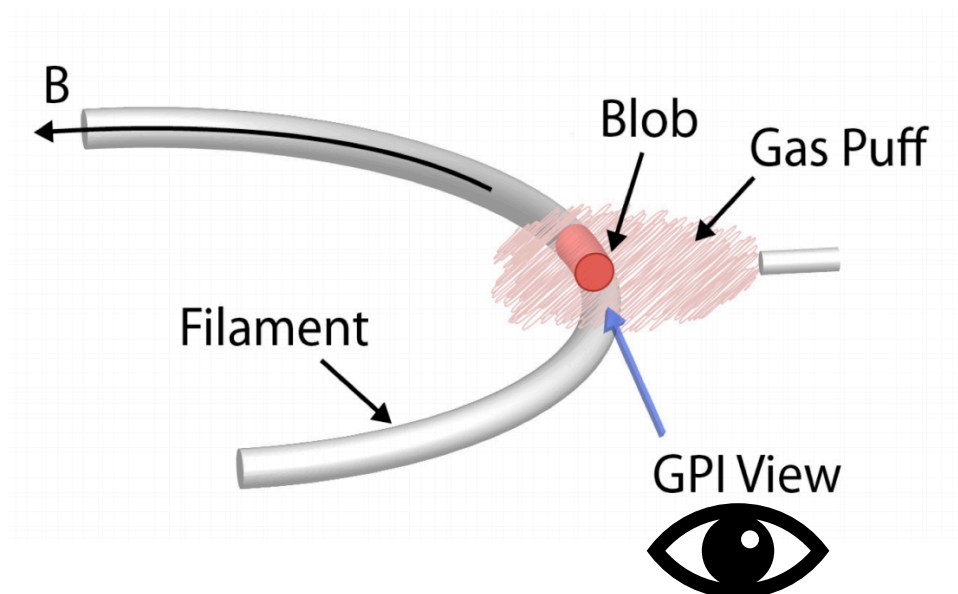
Imaging the SOL with GPI

- GPI = Gas-puff imaging diagnostic (S. Zweben)
- Real-time turbulence movies in NSTX SOL
- Data taken using fast camera (400,000 fr/s)
- $D\alpha$ intensity proportional to some combination of n and T

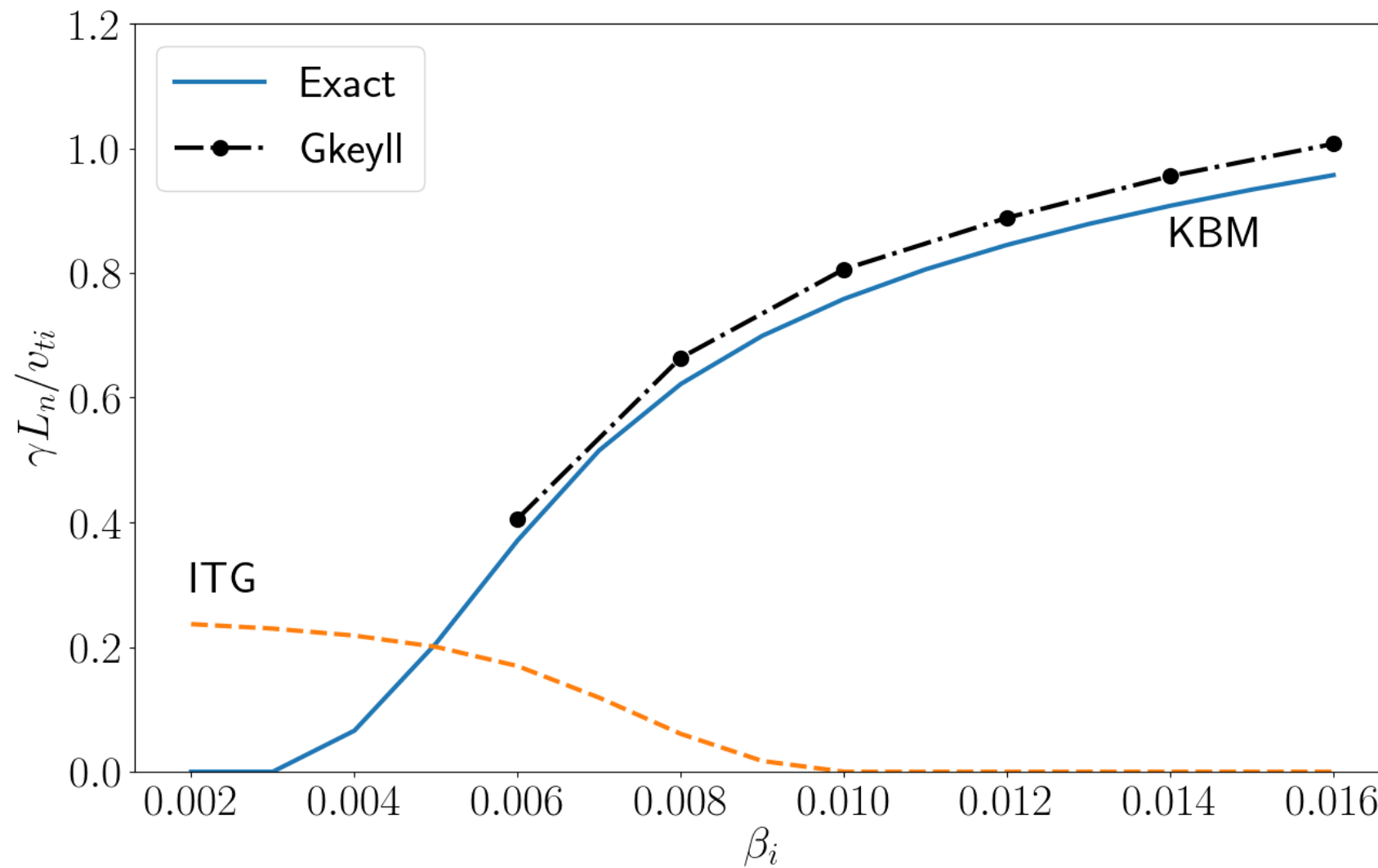


Imaging the SOL with GPI

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- $D\alpha$ intensity proportional to some combination of n and T



Linear benchmark: KBM instability (local limit)



$$k_{\perp} \rho_s = 0.5, \quad k_{\parallel} L_n = 0.1, \quad R/L_n = 5, \quad R/L_{Ti} = 12.5, \quad R/L_{Te} = 10, \quad \tau = 1$$